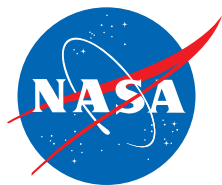


**NASA
Reference
Publication
1398**

March 1997

Total Solar Eclipse of 1999 August 11

Fred Espenak
and Jay Anderson



ECLIPSE PREDICTIONS

INTRODUCTION

On Wednesday, 1999 August 11, a total eclipse of the Sun will be visible from within a narrow corridor which traverses the Eastern Hemisphere. The path of the Moon's umbral shadow begins in the Atlantic and crosses central Europe, the Middle East, and India (Figures 3, 4 and 5) where it ends at sunset in the Bay of Bengal. A partial eclipse will be seen within the much broader path of the Moon's penumbral shadow, which includes northeastern North America, all of Europe, northern Africa and the western half of Asia (Figures 1 and 2).

UMBRAL PATH AND VISIBILITY

The last total solar eclipse of the 20th century begins in the North Atlantic about 300 kilometers south of Nova Scotia where the Moon's umbral shadow first touches down on Earth at 09:30:57 UT. Along the sunrise terminator, the maximum duration is a mere 47 seconds as seen from the center of the narrow 49 kilometers wide path. No major landfall occurs for the first forty minutes as the shadow sweeps across the North Atlantic. The umbra finally reaches the Isles of Scilly off the southwestern coast of England at 10:10 UT (Figure 6). At this locale, it is already mid-morning with the Sun 45° above the eastern horizon. The center line duration is 2 minutes and the path width has expanded to 103 kilometers as the shadow pursues its eastern track with a ground velocity of 0.91 km/s.

One minute later (10:11 UT), the umbra arrives along the shores of the Cornwall Peninsula. In the following four minutes, the shadow skirts the southern coast giving eager observers a brief taste of totality. Plymouth, the largest English city in the path, is north of center line and witnesses a total phase lasting 1 minute 39 seconds. London misses the total phase but experiences partiality with a maximum magnitude of 0.968. By 10:16 UT, the umbra leaves England as it quickly traverses the English Channel. The Channel Islands of Guernsey and Jersey lie just south of the path and witness a partial eclipse of magnitude >0.995. To the north, Alderney is deep in the path and enjoys over one and a half minutes of totality.

Not since 1961 has the Moon cast its dark shadow upon central Europe. The southern edge of the umbra first reaches the Normandy coast just as the northern edge leaves England (10:16 UT). But another four minutes elapse before the center line makes landfall in northern France (Figure 7). As the shadow sweeps through the French countryside, its southern edge passes 30 kilometers north of Paris. The City of Lights will bear witness to a partial event of magnitude 0.992 at 10:23 UT. Continuing on its eastward track, the path's northern limit crosses into southern Belgium, Luxembourg and Germany. Meanwhile, the center line cuts through Champagne where the citizens of Metz witness a total eclipse lasting 2 minutes 13 seconds (10:29 UT). Four minutes later, the entire umbra crosses into southern Germany (Figure 8) and the picturesque Rhine Valley. North of the path, Frankfurt witnesses a 0.979 magnitude partial eclipse, while Stuttgart lies near path center for 2 minutes 17 seconds of totality. At 10:35 UT, the Sun's altitude stands at 55°, the path width is 109 kilometers and the ground velocity is 0.74 km/s. Although München (Munich) lies 20 kilometers south of the center line, the city's two million citizens will still witness more than two minutes of totality, provided the winds of good fortune bring clear skies on eclipse day.

At 10:41 UT, the umbra leaves Germany and crosses into Austria where it encounters the Eastern Alps. Wien (Vienna) is almost 40 kilometers north of the path and experiences a 0.990 magnitude partial eclipse. The southern edge of the path grazes northeastern Slovenia as the shadow enters Hungary at 10:47 UT (Figure 9). Lake Balaton lies wholly within the path where the central duration lasts 2 minutes 22 seconds (10:50 UT). Like Wien, Budapest is also located about 40 kilometers north of the path where a 0.991 magnitude partial eclipse will occur. As the shadow leaves Hungary, the southern third briefly sweeps through northern Yugoslavia before continuing on into Romania.

The instant of greatest eclipse¹ occurs at 11:03:04 UT when the axis of the Moon's shadow passes closest to the center of Earth ($\gamma^2=0.506$). At that moment, the shadow's epicenter is located among the rolling hills of south-central Romania very near Rîmnicu-Vîlcea (Figure 10). The length of totality

¹ The instant of greatest eclipse occurs when the distance between the Moon's shadow axis and Earth's geocenter reaches a minimum. Although greatest eclipse differs slightly from the instants of greatest magnitude and greatest duration (for total eclipses), the differences are usually quite small.

² Minimum distance of the Moon's shadow axis from Earth's center in units of equatorial Earth radii.

reaches its maximum duration of 2 minutes 23 seconds, the Sun's altitude is 59°, the path width is 112 kilometers and the umbra's velocity is 0.680 km/s. Four minutes later (11:07 UT), Romania's capital city Bucuresti (Bucharest) is engulfed by the shadow. Since Bucharest lies on the center line near the instant of greatest eclipse, it enjoys a duration nearly as long at 2 minutes 22 seconds. Traveling south-southeast, the path encompasses the Romania-Bulgaria border before leaving land and heading out across the Black Sea.

The next landfall occurs along the Black Sea coast of northern Turkey at 11:21 UT (Figure 11). Ankara lies 150 kilometers south of the path and witnesses a 0.969 magnitude partial eclipse. The track diagonally bisects Turkey as it moves inland while the center line duration begins a gradual but steady decrease. At 11:29 UT, Turhal falls deep within the shadow for 2 minutes 15 seconds. The umbra reaches Turkey's southeastern border at 11:45 UT and briefly enters northwestern Syria as it crosses into Iraq (Figure 12). The center line duration is now 2 minutes 5 seconds with the Sun's altitude at 50°. Baghdad lies 220 kilometers south of the path and experiences a 0.940 magnitude partial eclipse (Figure 13). Arriving at Iran's western boundary at 11:52 UT, the shadow spends the next half hour crossing sparsely populated mountain ranges and deserts (Figures 13, 14 and 15). Tehran lies north of the path where its eight million inhabitants witness a 0.943 magnitude partial eclipse. At 12:22 UT, the shadow enters Pakistan and skirts the shores of the Arabian Sea (Figure 16). Karachi is near the center line and experiences 1 minute 13 seconds of total eclipse with the Sun 22° above the western horizon. The path width has shrunk to 85 kilometers while the shadow's speed has increased to 2 km/s.

The umbra arrives in India, the last nation in its path, at 12:28 UT (Figure 17). As the shadow sweeps across the sub-continent, its velocity rapidly increases while the center line duration drops below one minute and the Sun's altitude decreases to 7° (Figures 18 and 19). The eleven million inhabitants of Calcutta will witness a 0.879 magnitude partial eclipse with the Sun a scant 2° above the western horizon. Leaving India just north of Vishakhapatnam at 12:36 UT, the shadow sweeps into the Bay of Bengal where it departs Earth and races back into space (12:36:23 UT), not to return until the next millennium. Over the course of 3 hours and 7 minutes, the Moon's umbra travels along a path approximately 14000 kilometers long and covering 0.2% of Earth's surface area.

GENERAL MAPS OF THE ECLIPSE PATH

ORTHOGRAPHIC PROJECTION MAP OF THE ECLIPSE PATH

Figure 1 is an orthographic projection map of Earth [adapted from Espenak, 1987] showing the path of penumbral (partial) and umbral (total) eclipse. The daylight terminator is plotted for the instant of greatest eclipse with north at the top. The sub-Earth point is centered over the point of greatest eclipse and is indicated with an asterisk-like symbol. The sub-solar point at that instant is also shown.

The limits of the Moon's penumbral shadow define the region of visibility of the partial eclipse. This saddle shaped region often covers more than half of Earth's daylight hemisphere and consists of several distinct zones or limits. At the northern and/or southern boundaries lie the limits of the penumbra's path. Partial eclipses have only one of these limits, as do central eclipses when the shadow axis falls no closer than about 0.45 radii from Earth's center. Great loops at the western and eastern extremes of the penumbra's path identify the areas where the eclipse begins/ends at sunrise and sunset, respectively. If the penumbra has both a northern and southern limit, the rising and setting curves form two separate, closed loops. Otherwise, the curves are connected in a distorted figure eight. Bisecting the 'eclipse begins/ends at sunrise and sunset' loops is the curve of maximum eclipse at sunrise (western loop) and sunset (eastern loop). The exterior tangency points **P1** and **P4** mark the coordinates where the penumbral shadow first contacts (partial eclipse begins) and last contacts (partial eclipse ends) Earth's surface. The path of the umbral shadow bisects the penumbral path from west to east and is shaded dark gray.

A curve of maximum eclipse is the locus of all points where the eclipse is at maximum at a given time. They are plotted at each half hour Universal Time (UT), and generally run from northern to southern penumbral limits, or from the maximum eclipse at sunrise or sunset curves to one of the limits. The outline of the umbral shadow is plotted every ten minutes in UT. Curves of constant eclipse magnitude³ delineate the locus of all points where the magnitude at maximum eclipse is constant. These curves run

³ Eclipse magnitude is defined as the fraction of the Sun's diameter occulted by the Moon. It is strictly a ratio of *diameters* and should not be confused with eclipse obscuration which is a measure of the Sun's surface *area* occulted by the Moon. Eclipse magnitude may be expressed as either a percentage or a decimal fraction (e.g.: 50% or 0.50).

exclusively between the curves of maximum eclipse at sunrise and sunset. Furthermore, they are quasi-parallel to the northern/southern penumbral limits and the umbral paths of central eclipses. Northern and southern limits of the penumbra may be thought of as curves of constant magnitude of 0%, while adjacent curves are for magnitudes of 20%, 40%, 60% and 80%. The northern and southern limits of the path of total eclipse are curves of constant magnitude of 100%.

At the top of Figure 1, the Universal Time of geocentric conjunction between the Moon and Sun is given followed by the instant of greatest eclipse. The eclipse magnitude is given for greatest eclipse. For central eclipses (both total and annular), it is equivalent to the geocentric ratio of diameters of the Moon and Sun. Gamma is the minimum distance of the Moon's shadow axis from Earth's center in units of equatorial Earth radii. The shadow axis passes south of Earth's geocenter for negative values of Gamma. Finally, the Saros series number of the eclipse is given along with its relative sequence in the series.

STEREOGRAPHIC PROJECTION MAP OF THE ECLIPSE PATH

The stereographic projection of Earth in Figure 2 depicts the path of penumbral and umbral eclipse in greater detail. The map is oriented with the north up with the point of greatest eclipse near the center. International political borders are shown and circles of latitude and longitude are plotted at 20° increments. The region of penumbral or partial eclipse is identified by its northern and southern limits, curves of eclipse begins or ends at sunrise and sunset, and curves of maximum eclipse at sunrise and sunset. Curves of constant eclipse magnitude are plotted for 20%, 40%, 60% and 80%, as are the limits of the path of total eclipse. Also included are curves of greatest eclipse at every half hour Universal Time.

Figures 1 and 2 may be used to determine quickly the approximate time and magnitude of maximum eclipse at any location within the eclipse path.

EQUIDISTANT CONIC PROJECTION MAPS OF THE ECLIPSE PATH

Figures 3, 4 and 5 are equidistant conic projection maps chosen to minimize distortion, and which isolate specific regions of the umbral path. Once again, curves of maximum eclipse and constant eclipse magnitude are plotted and labeled. A linear scale is included for estimating approximate distances (kilometers). Within the northern and southern limits of the path of totality, the outline of the umbral shadow is plotted at ten minute intervals. The duration of totality (minutes and seconds) and the Sun's altitude correspond to the local circumstances on the center line at each shadow position.

The scale used in these figures is ~1:15,904,000. The positions of larger cities and metropolitan areas in and near the umbral path are depicted as black dots. The size of each city is logarithmically proportional to its population using 1990 census data (Rand McNally, 1991).

EQUIDISTANT CYLINDRICAL PROJECTION MAPS OF THE ECLIPSE PATH

Figures 6 through 19 all use a simple equidistant cylindrical projection scaled for the central latitude of each map. They all use high resolution coastline data from the World Data Base II (WDB) and World Vector Shoreline (WVS) data bases and have a scale of 1:2,784,000. These maps were chosen to isolate small regions along the entire land portion of the eclipse path. Once again, curves of maximum eclipse and constant eclipse magnitude are included as well as the outline of the umbral shadow. A special feature of these maps are the curves of constant umbral eclipse duration (i.e., totality) which are plotted within the path. These curves permit fast determination of approximate durations without consulting any tables. Furthermore, city data from a recently enlarged geographic data base of over 90,000 positions are plotted to give as many locations as possible in the path of totality. Local circumstances have been calculated for these positions and can be found in Tables 9 through 35.

ELEMENTS, SHADOW CONTACTS AND ECLIPSE PATH TABLES

The geocentric ephemeris for the Sun and Moon, various parameters, constants, and the Besselian elements (polynomial form) are given in Table 1. The eclipse elements and predictions were derived from the DE200 and LE200 ephemerides (solar and lunar, respectively) developed jointly by the Jet Propulsion Laboratory and the U. S. Naval Observatory for use in the *Astronomical Almanac* for 1984 and thereafter. Unless otherwise stated, all predictions are based on center of mass positions for the Moon and Sun with no corrections made for center of figure, lunar limb profile or atmospheric refraction. The predictions depart from normal IAU convention through the use of a smaller constant for the mean lunar radius k for all umbral contacts (see: LUNAR LIMB PROFILE). Times are expressed in either Terrestrial Dynamical Time (TDT) or in Universal Time (UT), where the best value of ΔT^4 available at the time of preparation is used.

From the polynomial form of the Besselian elements, any element can be evaluated for any time t_1 (in decimal hours) via the equation:

$$\mathbf{a} = \mathbf{a}_0 + \mathbf{a}_1 * \mathbf{t} + \mathbf{a}_2 * \mathbf{t}^2 + \mathbf{a}_3 * \mathbf{t}^3 \quad (\text{or } \mathbf{a} = \sum [\mathbf{a}_n * \mathbf{t}^n]; n = 0 \text{ to } 3)$$

where: $\mathbf{a} = x, y, d, l_1, l_2, \text{ or } \mu$
 $\mathbf{t} = t_1 - t_0$ (decimal hours) and $t_0 = 11.000$ TDT

The polynomial Besselian elements were derived from a least-squares fit to elements rigorously calculated at five separate times over a six hour period centered at t_0 . Thus, the equation and elements are valid over the period $8.00 \leq t_0 \leq 14.00$ TDT.

Table 2 lists all external and internal contacts of penumbral and umbral shadows with Earth. They include TDT times and geodetic coordinates with and without corrections for ΔT . The contacts are defined:

- P1** - Instant of first external tangency of penumbral shadow cone with Earth's limb.
(partial eclipse begins)
- P4** - Instant of last external tangency of penumbral shadow cone with Earth's limb.
(partial eclipse ends)
- U1** - Instant of first external tangency of umbral shadow cone with Earth's limb.
(umbral eclipse begins)
- U2** - Instant of first internal tangency of umbral shadow cone with Earth's limb.
- U3** - Instant of last internal tangency of umbral shadow cone with Earth's limb.
- U4** - Instant of last external tangency of umbral shadow cone with Earth's limb.
(umbral eclipse ends)

Similarly, the northern and southern extremes of the penumbral and umbral paths, and extreme limits of the umbral center line are given. The IAU longitude convention is used throughout this publication (i.e., for longitude, east is positive and west is negative; for latitude, north is positive and south is negative).

The path of the umbral shadow is delineated at five minute intervals in Universal Time in Table 3. Coordinates of the northern limit, the southern limit and the center line are listed to the nearest tenth of an arc-minute (~185 m at the Equator). The Sun's altitude, path width and umbral duration are calculated for the center line position. Table 4 presents a physical ephemeris for the umbral shadow at five minute intervals in UT. The center line coordinates are followed by the topocentric ratio of the apparent diameters of the Moon and Sun, the eclipse obscuration⁵, and the Sun's altitude and azimuth at that instant. The central path width, the umbral shadow's major and minor axes and its instantaneous velocity with respect to Earth's surface are included. Finally, the center line duration of the umbral phase is given.

Local circumstances for each center line position listed in Tables 3 and 4 are presented in Table 5. The first three columns give the Universal Time of maximum eclipse, the center line duration of totality and the altitude of the Sun at that instant. The following columns list each of the four eclipse contact times followed by their related contact position angles and the corresponding altitude of the Sun. The four contacts identify significant stages in the progress of the eclipse. They are defined as follows:

⁴ ΔT is the difference between Terrestrial Dynamical Time and Universal Time.

⁵ Eclipse obscuration is defined as the fraction of the Sun's surface area occulted by the Moon.

- First Contact** – Instant of first external tangency between the Moon and Sun.
(partial eclipse begins)
- Second Contact** – Instant of first internal tangency between the Moon and Sun.
(central or umbral eclipse begins; total or annular eclipse begins)
- Third Contact** – Instant of last internal tangency between the Moon and Sun.
(central or umbral eclipse ends; total or annular eclipse ends)
- Fourth Contact** – Instant of last external tangency between the Moon and Sun.
(partial eclipse ends)

The position angles **P** and **V** identify the point along the Sun's disk where each contact occurs⁶. Second and third contact altitudes are omitted since they are always within 1° of the altitude at maximum eclipse.

Table 6 presents topocentric values from the central path at maximum eclipse for the Moon's horizontal parallax, semi-diameter, relative angular velocity with respect to the Sun, and libration in longitude. The altitude and azimuth of the Sun are given along with the azimuth of the umbral path. The northern limit position angle identifies the point on the lunar disk defining the umbral path's northern limit. It is measured counter-clockwise from the north point of the Moon. In addition, corrections to the path limits due to the lunar limb profile are listed. The irregular profile of the Moon results in a zone of 'grazing eclipse' at each limit that is delineated by interior and exterior contacts of lunar features with the Sun's limb. This geometry is described in greater detail in the section **LIMB CORRECTIONS TO THE PATH LIMITS: GRAZE ZONES**. Corrections to center line durations due to the lunar limb profile are also included. When added to the durations in Tables 3, 4, 5 and 7, a slightly longer central total phase is predicted along most of the path.

To aid and assist in the plotting of the umbral path on large scale maps, the path coordinates are also tabulated at 1° intervals in longitude in Table 7. The latitude of the northern limit, southern limit and center line for each longitude is tabulated to the nearest hundredth of an arc-minute (~18.5 m at the Equator) along with the Universal Time of maximum eclipse at each position. Finally, local circumstances on the center line at maximum eclipse are listed and include the Sun's altitude and azimuth, the umbral path width and the central duration of totality.

In applications where the zones of grazing eclipse are needed in greater detail, Table 8 lists these coordinates over land based portions of the path at 30' intervals in longitude. The time of maximum eclipse is given at both northern and southern limits as well as the path's azimuth. The elevation and scale factors are also given (See: **LIMB CORRECTIONS TO THE PATH LIMITS: GRAZE ZONES**).

⁶ **P** is defined as the contact angle measured counter-clockwise from the *north* point of the Sun's disk.
V is defined as the contact angle measured counter-clockwise from the *zenith* point of the Sun's disk.

LOCAL CIRCUMSTANCES TABLES

Local circumstances for approximately 1460 cities, metropolitan areas and places in North America, Europe, Africa and Asia are presented in Tables 9 through 36. These tables give the local circumstances at each contact and at maximum eclipse⁷ for every location. The coordinates are listed along with the location's elevation (meters) above sea-level, if known. If the elevation is unknown (i.e., not in the data base), then the local circumstances for that location are calculated at sea-level. In any case, the elevation does not play a significant role in the predictions unless the location is near the umbral path limits and the Sun's altitude is relatively small ($<10^\circ$). The Universal Time of each contact is given to a tenth of a second, along with position angles **P** and **V** and the altitude of the Sun. The position angles identify the point along the Sun's disk where each contact occurs and are measured counter-clockwise (i.e., eastward) from the north and zenith points, respectively. Locations outside the umbral path miss the umbral eclipse and only witness first and fourth contacts. The Universal Time of maximum eclipse (either partial or total) is listed to a tenth of a second. Next, the position angles **P** and **V** of the Moon's disk with respect to the Sun are given, followed by the altitude and azimuth of the Sun at maximum eclipse. Finally, the corresponding eclipse magnitude and obscuration are listed. For umbral eclipses (both annular and total), the eclipse magnitude is identical to the topocentric ratio of the Moon's and Sun's apparent diameters. The eclipse magnitude is always less than 1 for annular eclipses and equal to or greater than 1 for total eclipses. The final column gives the duration of totality if this location lies in the path of the Moon's umbral shadow. The effects of refraction have not been included in these calculations, nor have there been any corrections for center of figure or the lunar limb profile.

Locations were chosen based on general geographic distribution, population, and proximity to the path. The primary source for geographic coordinates is *The New International Atlas* (Rand McNally, 1991). Elevations for major cities were taken from *Climates of the World* (U. S. Dept. of Commerce, 1972). In this rapidly changing political world, it is often difficult to ascertain the correct name or spelling for a given location. Therefore, the information presented here is for location purposes only and is not meant to be authoritative. Furthermore, it does not imply recognition of status of any location by the United States Government. Corrections to names, spellings, coordinates and elevations is solicited in order to update the geographic data base for future eclipse predictions.

For countries in the path of totality, expanded versions of the local circumstances tables listing many more locations are available via a special web site of supplemental material for the total solar eclipse of 1999 (<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>). This site also has tables of local circumstances for all astronomical observatories listed in the *Astronomical Almanac for 1996*.

DETAILED MAPS OF THE UMBRAL PATH

The path of totality has been plotted by hand on a set of ten detailed maps appearing in the last section of this publication. The maps are Global Navigation and Planning Charts or GNC's from the Defense Mapping Agency, which use a Lambert conformal conic projection. More specifically, GNC-4 covers Europe and western Asia while GNC-12 covers the Middle East and south Asia. GNC's have a scale of 1:5,000,000 (1 inch ~ 69 nautical miles), which is adequate for showing major cities, highways, airports, rivers, bodies of water and basic topography required for eclipse expedition planning including site selection, transportation logistics and weather contingency strategies.

Northern and southern limits, as well as the center line of the path, are plotted using data from Table 7. Although no corrections have been made for center of figure or lunar limb profile, they have little or no effect at this scale. Atmospheric refraction has not been included, as its effects play a significant role only at very low solar altitudes. In any case, refraction corrections to the path are uncertain since they depend on the atmospheric temperature-pressure profile, which cannot be predicted in advance. If observations from the graze zones are planned, then the zones of grazing eclipse must be plotted on higher scale maps using coordinates in Table 8. See PLOTTING THE PATH ON MAPS for sources and more information. The GNC paths also depict the curve of maximum eclipse at five and/or ten minute increments in UT from Table 3.

Selected sections of the path are plotted on higher resolution maps (i.e., ONC maps with a scale of 1:1,000,000) which are available via a special web site of supplemental material for the total solar eclipse of 1999 (<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>).

⁷ For partial eclipses, maximum eclipse is the instant when the greatest fraction of the Sun's diameter is occulted. For total eclipses, maximum eclipse is the instant of mid-totality.

ESTIMATING TIMES OF SECOND AND THIRD CONTACTS

The times of second and third contact for any location not listed in this publication can be estimated using the detailed maps found in the final section. Alternatively, the contact times can be estimated from maps on which the umbral path has been plotted. Table 7 lists the path coordinates conveniently arranged in 1° increments of longitude to assist plotting by hand. The path coordinates in Table 3 define a line of maximum eclipse at five minute increments in time. These lines of maximum eclipse each represent the projection diameter of the umbral shadow at the given time. Thus, any point on one of these lines will witness maximum eclipse (i.e., mid-totality) at the same instant. The coordinates in Table 3 should be added to the map in order to construct lines of maximum eclipse.

The estimation of contact times for any one point begins with an interpolation for the time of maximum eclipse at that location. The time of maximum eclipse is proportional to a point's distance between two adjacent lines of maximum eclipse, measured along a line parallel to the center line. This relationship is valid along most of the path with the exception of the extreme ends, where the shadow experiences its largest acceleration. The center line duration of totality **D** and the path width **W** are similarly interpolated from the values of the adjacent lines of maximum eclipse as listed in Table 3. Since the location of interest probably does not lie on the center line, it is useful to have an expression for calculating the duration of totality **d** as a function of its perpendicular distance **a** from the center line:

$$\mathbf{d} = \mathbf{D} (1 - (2 \mathbf{a}/\mathbf{W})^2)^{1/2} \text{ seconds} \quad [1]$$

where: **d** = duration of totality at desired location (seconds)
D = duration of totality on the center line (seconds)
a = perpendicular distance from the center line (kilometers)
W = width of the path (kilometers)

If **t_m** is the interpolated time of maximum eclipse for the location, then the approximate times of second and third contacts (**t₂** and **t₃**, respectively) are:

$$\begin{array}{ll} \text{Second Contact:} & \mathbf{t}_2 = \mathbf{t}_m - \mathbf{d}/2 \quad [2] \\ \text{Third Contact:} & \mathbf{t}_3 = \mathbf{t}_m + \mathbf{d}/2 \quad [3] \end{array}$$

The position angles of second and third contact (either **P** or **V**) for any location off the center line are also useful in some applications. First, linearly interpolate the center line position angles of second and third contacts from the values of the adjacent lines of maximum eclipse as listed in Table 5. If **X₂** and **X₃** are the interpolated center line position angles of second and third contacts, then the position angles **x₂** and **x₃** of those contacts for an observer located **a** kilometers from the center line are:

$$\begin{array}{ll} \text{Second Contact:} & \mathbf{x}_2 = \mathbf{X}_2 - \arcsin (2 \mathbf{a}/\mathbf{W}) \quad [4] \\ \text{Third Contact:} & \mathbf{x}_3 = \mathbf{X}_3 + \arcsin (2 \mathbf{a}/\mathbf{W}) \quad [5] \end{array}$$

where: **x_n** = interpolated position angle (either **P** or **V**) of contact **n** at location
X_n = interpolated position angle (either **P** or **V**) of contact **n** on center line
a = perpendicular distance from the center line (kilometers)
 (use negative values for locations south of the center line)
W = width of the path (kilometers)

MEAN LUNAR RADIUS

A fundamental parameter used in eclipse predictions is the Moon's radius k , expressed in units of Earth's equatorial radius. The Moon's actual radius varies as a function of position angle and libration due to the irregularity in the limb profile. From 1968 through 1980, the Nautical Almanac Office used two separate values for k in their predictions. The larger value ($k=0.2724880$), representing a mean over topographic features, was used for all penumbral (exterior) contacts and for annular eclipses. A smaller value ($k=0.272281$), representing a mean minimum radius, was reserved exclusively for umbral (interior) contact calculations of total eclipses [*Explanatory Supplement*, 1974]. Unfortunately, the use of two different values of k for umbral eclipses introduces a discontinuity in the case of hybrid or annular-total eclipses.

In August 1982, the International Astronomical Union (IAU) General Assembly adopted a value of $k=0.2725076$ for the mean lunar radius. This value is now used by the Nautical Almanac Office for all solar eclipse predictions [Fiala and Lukac, 1983] and is currently the best mean radius, averaging mountain peaks and low valleys along the Moon's rugged limb. The adoption of one single value for k eliminates the discontinuity in the case of annular-total eclipses and ends confusion arising from the use of two different values. However, the use of even the best 'mean' value for the Moon's radius introduces a problem in predicting the true character and duration of umbral eclipses, particularly total eclipses. A total eclipse can be defined as an eclipse in which the Sun's disk is completely occulted by the Moon. This cannot occur so long as any photospheric rays are visible through deep valleys along the Moon's limb [Meeus, Grosjean and Vanderleen, 1966]. But the use of the IAU's mean k guarantees that some annular or annular-total eclipses will be misidentified as total. A case in point is the eclipse of 3 October 1986. Using the IAU value for k , the *Astronomical Almanac* identified this event as a total eclipse of 3 seconds duration when it was, in fact, a beaded annular eclipse. Since a smaller value of k is more representative of the deeper lunar valleys and hence the minimum solid disk radius, it helps ensure the correct identification of an eclipse's true nature.

Of primary interest to most observers are the times when umbral eclipse begins and ends (second and third contacts, respectively) and the duration of the umbral phase. When the IAU's value for k is used to calculate these times, they must be corrected to accommodate low valleys (total) or high mountains (annular) along the Moon's limb. The calculation of these corrections is not trivial but must be performed, especially if one plans to observe near the path limits [Herald, 1983]. For observers near the center line of a total eclipse, the limb corrections can be more closely approximated by using a smaller value of k which accounts for the valleys along the profile.

This publication uses the IAU's accepted value of $k=0.2725076$ for all penumbral (exterior) contacts. In order to avoid eclipse type misidentification and to predict central durations which are closer to the actual durations at total eclipses, we depart from standard convention by adopting the smaller value of $k=0.272281$ for all umbral (interior) contacts. This is consistent with predictions in *Fifty Year Canon of Solar Eclipses: 1986 - 2035* [Espenak, 1987]. Consequently, the smaller k produces shorter umbral durations and narrower paths for total eclipses when compared with calculations using the IAU value for k . Similarly, predictions using a smaller k result in longer umbral durations and wider paths for annular eclipses than do predictions using the IAU's k .

LUNAR LIMB PROFILE

Eclipse contact times, magnitude and duration of totality all depend on the angular diameters and relative velocities of the Moon and Sun. Unfortunately, these calculations are limited in accuracy by the departure of the Moon's limb from a perfectly circular figure. The Moon's surface exhibits a rather dramatic topography, which manifests itself as an irregular limb when seen in profile. Most eclipse calculations assume some mean radius that averages high mountain peaks and low valleys along the Moon's rugged limb. Such an approximation is acceptable for many applications, but if higher accuracy is needed, the Moon's actual limb profile must be considered. Fortunately, an extensive body of knowledge exists on this subject in the form of Watts' limb charts [Watts, 1963]. These data are the product of a photographic survey of the marginal zone of the Moon and give limb profile heights with respect to an adopted smooth reference surface (or datum). Analyses of lunar occultations of stars by Van Flandern [1970] and Morrison [1979] have shown that the average cross-section of Watts' datum is slightly elliptical rather than circular. Furthermore, the implicit center of the datum (i.e., the center of figure) is displaced from the Moon's center of mass. In a follow-up analysis of 66,000 occultations, Morrison and Appleby [1981] have found that the radius of the datum appears to vary with libration. These variations produce systematic errors in Watts' original limb profile heights that attain 0.4 arc-seconds at some position angles. Thus, corrections to Watts' limb data are necessary to ensure that the reference datum is a sphere with its center at the center of mass.

The Watts charts have been digitized by Her Majesty's Nautical Almanac Office in Herstmonceux, England, and transformed to grid-profile format at the U. S. Naval Observatory. In this computer readable form, the Watts limb charts lend themselves to the generation of limb profiles for any lunar libration. Ellipticity and libration corrections may be applied to refer the profile to the Moon's center of mass. Such a profile can then be used to correct eclipse predictions which have been generated using a mean lunar limb.

Along the path, the Moon's topocentric libration (physical + optical) in longitude ranges from $l=+5.5^\circ$ to $l=+4.0^\circ$. Thus, a limb profile with the appropriate libration is required in any detailed analysis of contact times, central durations, etc.. But a profile with an intermediate value is useful for planning purposes and may even be adequate for most applications. The lunar limb profile presented in Figure 20 includes corrections for center of mass and ellipticity [Morrison and Appleby, 1981]. It is generated for 11:00 UT, which corresponds to western Romania just before greatest eclipse. The Moon's topocentric libration is $l=+4.81^\circ$, and the topocentric semi-diameters of the Sun and Moon are 946.8 and 973.9 arc-seconds, respectively. The Moon's angular velocity with respect to the Sun is 0.379 arc-seconds per second.

The radial scale of the limb profile in Figure 20 (at bottom) is greatly exaggerated so that the true limb's departure from the mean lunar limb is readily apparent. The mean limb with respect to the center of figure of Watts' original data is shown (dashed) along with the mean limb with respect to the center of mass (solid). Note that all the predictions presented in this publication are calculated with respect to the latter limb unless otherwise noted. Position angles of various lunar features can be read using the protractor marks along the Moon's mean limb (center of mass). The position angles of second and third contact are clearly marked along with the north pole of the Moon's axis of rotation and the observer's zenith at mid-totality. The dashed line with arrows at either end identifies the contact points on the limb corresponding to the northern and southern limits of the path. To the upper left of the profile are the Sun's topocentric coordinates at maximum eclipse. They include the right ascension *R.A.*, declination *Dec.*, semi-diameter *S.D.* and horizontal parallax *H.P.*. The corresponding topocentric coordinates for the Moon are to the upper right. Below and left of the profile are the geographic coordinates of the center line at 11:00 UT while the times of the four eclipse contacts at that location appear to the lower right. Directly below the profile are the local circumstances at maximum eclipse. They include the Sun's altitude and azimuth, the path width, and central duration. The position angle of the path's northern/southern limit axis is *PA(N.Limit)* and the angular velocity of the Moon with respect to the Sun is *A.Vel.(M:S)*. At the bottom left are a number of parameters used in the predictions, and the topocentric lunar librations appear at the lower right.

In investigations where accurate contact times are needed, the lunar limb profile can be used to correct the nominal or mean limb predictions. For any given position angle, there will be a high mountain (annular eclipses) or a low valley (total eclipses) in the vicinity that ultimately determines the true instant of contact. The difference, in time, between the Sun's position when tangent to the contact point on the mean limb and tangent to the highest mountain (annular) or lowest valley (total) at actual contact is the desired correction to the predicted contact time. On the exaggerated radial scale of Figure 20, the Sun's limb can be represented as an epicyclic curve that is tangent to the mean lunar limb at the point of contact and departs from the limb by **h** through:

$$\mathbf{h} = \mathbf{S} (\mathbf{m}-1) (1-\cos[\mathbf{C}]) \quad [6]$$

where: **h** = departure of Sun's limb from mean lunar limb
S = Sun's semi-diameter
m = eclipse magnitude
C = angle from the point of contact

Herald [1983] has taken advantage of this geometry to develop a graphical procedure for estimating correction times over a range of position angles. Briefly, a displacement curve of the Sun's limb is constructed on a transparent overlay by way of equation [6]. For a given position angle, the solar limb overlay is moved radially from the mean lunar limb contact point until it is tangent to the lowest lunar profile feature in the vicinity. The solar limb's distance **d** (arc-seconds) from the mean lunar limb is then converted to a time correction Δ by:

$$\Delta = \mathbf{d} \mathbf{v} \cos[\mathbf{X} - \mathbf{C}] \quad [7]$$

where: Δ = correction to contact time (seconds)
d = distance of Solar limb from Moon's mean limb (arc-sec)
v = angular velocity of the Moon with respect to the Sun (arc-sec/sec)
X = center line position angle of the contact
C = angle from the point of contact

This operation may be used for predicting the formation and location of Baily's beads. When calculations are performed over a large range of position angles, a contact time correction curve can then be constructed.

Since the limb profile data are available in digital form, an analytical solution to the problem is possible that is quite straightforward and robust. Curves of corrections to the times of second and third contact for most position angles have been computer generated and are plotted in Figure 20. The circular protractor scale at the center represents the nominal contact time using a mean lunar limb. The departure of the contact correction curves from this scale graphically illustrates the time correction to the mean predictions for any position angle as a result of the Moon's true limb profile. Time corrections external to the circular scale are added to the mean contact time; time corrections internal to the protractor are subtracted from the mean contact time. The magnitude of the time correction at a given position angle is measured using any of the four radial scales plotted at each cardinal point.

For example, Table 17 gives the following data for München, Germany:

Second Contact = 10:37:12.3 UT $P_2=130^\circ$

Third Contact = 10:39:20.1 UT $P_3=263^\circ$

Using Figure 20, the measured time corrections and the resulting contact times are:

$C_2=-3.1$ seconds; Second Contact = 10:37:12.3 -3.1s = 10:37:09.2 UT

$C_3=-1.7$ seconds; Third Contact = 10:39:20.1 -1.7s = 10:39:18.4 UT

The above corrected values are within 0.1 seconds of a rigorous calculation using the actual limb profile.

Lunar limb profile diagrams for a number of other positions/times along the path of totality are available via a special web site of supplemental material for the total solar eclipse of 1999 (<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>).

LIMB PROFILE EFFECTS ON THE DURATION OF TOTALITY

As was previously discussed, the Moon's center of figure (i.e., the geometric center of the Watts' datum) is displaced from the Moon's center of mass. A case in point is the lunar limb geometry at 11:00 UT (Figure 20) where the center of figure is displaced -0.28 and $+0.38$ arc-seconds in ecliptic latitude and longitude, respectively. This shift is fairly characteristic along much of the 1999 umbral path but varies considerably between eclipses due to different libration geometries. Since most predictions appearing in this publication are calculated with respect to the Moon's center of mass, the center of figure offset has a small but significant consequence on the duration of totality. When compounded with the irregularities of the lunar limb profile, the overall result is to shift the maximum duration of totality south of the center line by 2-6 kilometers along the European path, and 3-10 kilometers along the Middle Eastern path.

Furthermore, the true limb profile of the 1999 eclipse actually produces a longer duration of totality than the one calculated using a mean limb. This results in maximum durations lasting 1 to 3 seconds longer than the nominal center line durations along much of the path.

Figure 21 shows a series of calculations for the duration of totality within ± 30 kilometers of the center line and spaced at ten minute intervals along the path through Europe (a) and the Middle East (b). For a given time, the duration of totality is calculated at 1 kilometer intervals perpendicular to the path within a 60 kilometer zone centered on the center line. Predictions using the Moon's center of mass and mean limb are represented by the dotted curves. Predictions using the actual limb profile to calculate corrected contact times and the resulting duration of totality are plotted as solid curves. What becomes immediately apparent upon inspection of Figure 21, is the asymmetry of the true limb duration curves and is a consequence of the complex Sun/Moon limb geometry which changes quickly with path position.

Observers wishing to witness the maximum possible duration of totality from a given section of the path can use Figure 21 to optimize their location with respect to the center line.

LIMB CORRECTIONS TO THE PATH LIMITS: GRAZE ZONES

The northern and southern umbral limits provided in this publication were derived using the Moon's center of mass and a mean lunar radius. They have not been corrected for the Moon's center of figure or the effects of the lunar limb profile. In applications where precise limits are required, Watts' limb data must be used to correct the nominal or mean path. Unfortunately, a single correction at each limit is not possible since the Moon's libration in longitude and the contact points of the limits along the Moon's limb each vary as a function of time and position along the umbral path. This makes it necessary to calculate a unique correction to the limits at each point along the path. Furthermore, the northern and southern limits of the umbral path are actually paralleled by a relatively narrow zone where the eclipse is neither penumbral nor umbral. An

observer positioned here will witness a slender solar crescent that is fragmented into a series of bright beads and short segments whose morphology changes quickly with the rapidly varying geometry between the limbs of the Moon and the Sun. These beading phenomena are caused by the appearance of photospheric rays that alternately pass through deep lunar valleys and hide behind high mountain peaks as the Moon's irregular limb grazes the edge of the Sun's disk. The geometry is directly analogous to the case of grazing occultations of stars by the Moon. The graze zone is typically five to ten kilometers wide and its interior and exterior boundaries can be predicted using the lunar limb profile. The interior boundaries define the actual limits of the umbral eclipse (both total and annular) while the exterior boundaries set the outer limits of the grazing eclipse zone.

Table 6 provides topocentric data and corrections to the path limits due to the true lunar limb profile. At five minute intervals, the table lists the Moon's topocentric horizontal parallax, semi-diameter, relative angular velocity of the Moon with respect to the Sun and lunar libration in longitude. The Sun's center line altitude and azimuth is given, followed by the azimuth of the umbral path. The position angle of the point on the Moon's limb which defines the northern limit of the path is measured counter-clockwise (i.e., eastward) from the north point on the limb. The path corrections to the northern and southern limits are listed as interior and exterior components in order to define the graze zone. Positive corrections are in the northern sense while negative shifts are in the southern sense. These corrections (minutes of arc in latitude) may be added directly to the path coordinates listed in Table 3. Corrections to the center line umbral durations due to the lunar limb profile are also included and they are mostly positive. Thus, when added to the central durations given in Tables 3, 4, 5 and 7, a slightly longer central total phase is predicted.

Detailed coordinates for the zones of grazing eclipse at each limit for all land based sections of the path are presented in Table 8. Given the uncertainties in the Watts data, these predictions should be accurate to ± 0.3 arc-seconds. The interior graze coordinates take into account the deepest valleys along the Moon's limb which produce the simultaneous second and third contacts at the path limits. Thus, the interior coordinates define the true edge of the path of totality. They are calculated from an algorithm which searches the path limits for the extreme positions where no photospheric beads are visible along a $\pm 30^\circ$ segment of the Moon's limb, symmetric about the extreme contact points at the instant of maximum eclipse. The exterior graze coordinates are somewhat arbitrarily defined and calculated for the geodetic positions where an unbroken photospheric crescent of 60° in angular extent is visible at maximum eclipse.

In Table 8, the graze zone latitudes are listed every 1° in longitude (at sea level) and include the time of maximum eclipse at the northern and southern limits as well as the path's azimuth. To correct the path for locations above sea level, *Elev Fact*⁸ is a multiplicative factor by which the path must be shifted north perpendicular to itself (i.e., perpendicular to path azimuth) for each unit of elevation (height) above sea level. To calculate the shift, a location's elevation is multiplied by the *Elev Fact* value. Negative values (usually the case for eclipses in the Northern Hemisphere) indicate that the path must be shifted south. For instance, if one's elevation is 1000 meters above sea level and the *Elev Fact* value is -0.20 , then the shift is -200m ($= 1000\text{m} \times -0.20$). Thus, the observer must shift the path coordinates 200 meters in a direction perpendicular to the path and in a negative or southerly sense.

The final column of Table 8 lists the *Scale Fact* (km/arc-second). This scaling factor provides an indication of the width of the zone of grazing phenomena, due to the topocentric distance of the Moon and the projection geometry of the Moon's shadow on Earth's surface. Since the solar chromosphere has an apparent thickness of about 3 arc-seconds, and assuming a *Scale Fact* value of 2 km/arc-seconds, then the chromosphere should be visible continuously during totality for any observer in the path who is within 6 kilometers ($= 2 \times 3$) of each interior limit. However, the most dynamic beading phenomena occurs within 1.5 arc-seconds of the Moon's limb. Using the above Scale Factor, this translates into the first 3 kilometers inside the interior limits. But observers should position themselves at least 1 kilometer inside the interior limits (south of the northern interior limit or north of the southern interior limit) in order to ensure that they are inside the path due to small uncertainties in Watts' data and the actual path limits.

For applications where the zones of grazing eclipse are needed at a higher frequency in longitude interval, tables of coordinates every 7.5' in longitude are available via a special web site for the total solar eclipse of 1999 (<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>).

⁸ is the product, $\tan(90-A) * \sin(D)$, where A is the altitude of the Sun and D is the difference between the azimuth of the Sun and the azimuth of the limit line, with the sign selected to be positive if the path should be shifted north with positive elevations above sea level.

SAROS HISTORY

The total eclipse of 1999 August 11 is the twenty-first member of Saros series 145 (Table 37), as defined by van den Bergh [1955]. All eclipses in the series occur at the Moon's ascending node and gamma⁹ decreases with each member in the family. The series is a young one which began with a minuscule partial eclipse at high northern hemisphere latitudes on 1639 Jan 04. After fourteen partial eclipses each of increasing magnitude, the first central eclipse occurred on 1891 Jun 06. The event was a six second annular eclipse with a path sweeping through eastern Siberia and the Arctic Ocean. Although the vertex of the umbral shadow fell just short of Earth's surface, the Moon's distance was gradually decreasing with each subsequent eclipse in the series. In fact, the very next eclipse was a hybrid or annular/total eclipse on 1909 Jun 17. Greatest eclipse occurred in the Arctic Ocean and lasted 24 seconds.

The third central eclipse of Saros 145 occurred on 1927 Jun 29. It was the first total eclipse of the family and coincidentally passed through England in addition to Scandinavia and Siberia. On 1945 Jul 09, the path of totality began in Idaho and quickly swept northeast through Montana, Saskatchewan and Manitoba. After crossing Hudson Bay, Greenland and the North Atlantic, the umbra returned to Scandinavia and Siberia. The fifth central eclipse occurred on 1963 Jul 20 and is well known to many eclipse observers. Its path crossed Alaska, central and eastern Canada and Maine. The event drew a great deal of media attention and a beautiful article about the eclipse appeared months later in the pages of NATIONAL GEOGRAPHIC [November 1963]. In fact, one of the authors (Espenak) has fond memories of watching the partial phases of this eclipse as a boy from his grandmother's home in Long Island.

The most recent eclipse of the series took place on 1981 Jul 31 and its path crossed central Siberia, Sakhalin Island and the Pacific Ocean where it ended north of Hawaii. After 1999, the following member occurs on 2017 Aug 21. *This* is the first total solar eclipse visible from the continental United States since 1979 Feb 26. The path of totality stretches from Oregon through Idaho, Wyoming, Nebraska, Missouri, Illinois, Kentucky, Tennessee and the Carolinas and has a greatest duration of 2m 40s (<http://planets.gsfc.nasa.gov/eclipse/eclipse/SEmap/TSENorAm2001.GIF>).

During the 21st through 24th centuries, Saros 145 continues to produce total solar eclipses of increasing duration as the path of each event shifts southward. By the time the midpoint of the series is reached (2324 Feb 25), the duration of totality exceeds four minutes. The duration continues to increase into the 25th and 26th centuries. The maximum duration of totality peaks at 7m 12s on 2522 Jun 25. In the remaining six umbral eclipses, the duration rapidly drops but still lasts almost three minutes during the final total eclipse on 2648 Sep 09.

For the next three and a half centuries, twenty partial eclipses of progressively decreasing magnitude occur. The final event takes place on 3009 Apr 17 from the polar regions of the Southern Hemisphere. A detailed list of eclipses in Saros series 145 appears in Table 37.

In summary, Saros series 145 includes 77 eclipses with the following distribution:

Saros 145	<u>Partial</u>	<u>Annular</u>	<u>Ann/Total</u>	<u>Total</u>
Non-Central	34	0	0	0
Central	—	1	1	41

⁹ Minimum distance of the Moon's shadow axis from Earth's center in units of equatorial Earth radii. Gamma defines the instant of greatest eclipse and takes on negative values south of the Earth's center.

WEATHER PROSPECTS FOR THE ECLIPSE

INTRODUCTION

The lunar shadow begins its three hour trek off the coast of Nova Scotia in the cloudy skies of the North Atlantic, just beyond Canada's Sable Island. A third of the globe later it reaches land in southwest England. While this is the best location in Great Britain, better weather can be found farther along the track where the frequency of sunny weather improves steadily into the Middle East. Beyond Iran, the Indian monsoon rapidly diminishes the chances of a view of this spectacle and sunset finds the shadow over the very cloudy skies of the Bay of Bengal.

THE BRITISH ISLES

Of all of the countries of Europe, the British Isles are most exposed to the varying weather from the mobile westerly air flow off of the North Atlantic. While blessed with mild winters, the dampness, cool summers and meteorological exposure to the westerlies has led to the pronouncement that the Isles have "no climate, only weather." Historically, Britain has appealed more to invading northern peoples rather than those from the south. The former found the warm moist climate an improvement on their homelands, while to the latter, especially the French, it appeared illogical to covet such a cloudy, damp and windy land (Manley, 1970). While good for the "English lawn," the damp gray skies will not find favor at eclipse time.

But fairness dictates that we note the eclipse comes at the height of summer over the sunniest parts of the English countryside. Land's End and the south coast of England offer the best prospects for sunshine anywhere in the Isles, and so by good fortune the eclipse track is located to its best advantage.

The upper air currents which carry the weather systems onto the islands are predominantly westerly. The main track of invading low pressure centers is north of Scotland, but trailing cold fronts frequently drag across England as these disturbances pass, bringing persistent cloudiness and precipitation. The cooler, drier high pressure systems following these fronts do not always improve sky conditions, as the colder airmass is often unstable and speckles the afternoon landscape with showers or thundershowers.

Periodic blocking patterns occasionally interrupt the alternating stream of highs and lows, bringing episodes of stable unchanging weather. This interruption in the variable weather pattern may or may not favor the eclipse observer depending on where the block develops. If the circulation over the islands is anticyclonic, as when a high builds northward from the Azores anticyclone or takes up residence over western France, then sunny dry weather will dominate, likely for several days. However, when cyclonic circulations settle in, especially if the low center lingers over England, dull gray skies will dominate.

Long term studies suggest that a variable westerly flow can be found about 28% of the time, cyclonic blocking for 24% and anticyclonic patterns for 23%. Suitable eclipse-viewing weather will be found during anticyclonic days and possibly in the wake of cold fronts during the changeable westerly weather. In the wake of a cold frontal passage, coastal areas are more likely to have sunshine than interior locations as the cooling effect of the nearby ocean suppresses the development of convective cloud. This distinction is often enhanced by the formation of sea breeze circulations which can bring a zone of clear skies along the waterfront to an otherwise cloudy day. The zone of clearing will extend inland at best for only a few kilometers depending on the precise nature of the onshore winds.

Realistically, however, climate statistics suggest that England is the least suitable land location from which to view the eclipse except for parts of India and Pakistan where monsoon cloudiness dominates. The satellite-derived climatology drawn in Figure 22 shows a mean cloud cover of 55 to 65% across southern England. Surface observations show an average of 5.8 to 6.5 days of the month with scattered cloud and good visibility (Table 38), slightly lower than much of the rest of western Europe, but only a third of the cloud-free days available in Bulgaria and Romania, and a quarter to a fifth the number of days in the best locations in Turkey and Iran. This rather heavy cloudiness culminates in a meager 43% of the possible sunshine in England.

Land's End and The Lizard are among the sunniest spots in the England thanks to the suppression of convective cloud by sea breezes and a slightly closer position to the semi-permanent anticyclone over the Azores. The effect of these two factors is minor but detectable and can be seen by comparing statistics for Plymouth and The Lizard in Table 38. Data collected at the former show a frequency of scattered cloud and good visibilities of 5.8 days per month; at The Lizard this statistic improves slightly to 6.5 days per month. Most climatological studies concede a five to ten percent increase in sunshine to the coastal sites over those inland, but much of this advantage is only realized during unstable convective days, such as those found

behind a cold front. Figure 23 promises a stationary observer only a 45% chance of seeing the eclipse in southern England.

Whatever the cause, low level stratiform clouds are the most likely type of cloud along the south coast (Warren et. al, 1986). These clouds are formed as warm maritime air approaches from the south and moves over the colder waters of the nearby English Channel. Such clouds typically cover three-quarters of the sky when present, making a clear patch difficult to find. The best location to find a gap in this kind of cloud cover is to seek shelter against the southerly flow behind the low range of hills which define the spine of Cornwall so that low level winds are flowing downhill and drying the air. The hills near the center line between Penzance and The Lizard do not accord much protection, in part because they are not quite oriented to lie across the likely direction of the wind. In the event of a heavy layer of low cloud, the eclipse watcher is best advised to sacrifice some of the duration of the eclipse and head northeast along the highway toward Bodmin, leaving the center line behind. It's a small chance, but the best that can be offered under the circumstances.

Western Europe

During the summer months, the "European monsoon" sets in across the western half of the continent – a period of cloudy and showery weather which begins with the onset of westerly winds in mid June and continues into September. At other times of the year the prevailing wind direction aloft is more variable, bringing alternating spells of warm and cold weather. This does not imply that cold air does not spread southward in summer, for cold spells with unstable weather come at intervals of a week or two, just as in North America. In fact, the weather along the eclipse track is similar in character to that found in the northern plains States and Canadian prairie provinces, with frequent weak disturbances bringing showers and thundershowers. Intervals between disturbances often come with fine and dry weather, though these are less frequent in Europe than in North America. Prolonged dry weather in Europe comes when a high pressure system builds northward from the Azores anticyclone.

The varied topography through Germany and Austria creates a wealth of climatic sub-regions, but these are defined more by temperature and precipitation differences than in cloudiness. Cloud patterns are strongly modified by the higher mountain ranges, particularly the Alps and Jura Mountains to the south of the eclipse track, but the main effect of these barriers is to create cloud rather than dissipate it. The prevailing westerly and north-westerly winds are forced to rise up these slopes, cooling and causing clouds to form. Southerly winds, flowing over the mountains toward the eclipse track, would tend to dissipate the cloud and bring dry sunny weather on the downhill side (a process known as a chinook in North America and a foehn wind in Europe) but southerlies are relatively rare during August and the benefits of the mountain barrier are not likely to be realized.

In spite of the influence of the Alps, there is a steady trend to sunnier skies as the eclipse track proceeds across Europe, and the distance from the Atlantic moisture source increases. The percent frequency of clear skies is about 18% over Normandy on the Channel coast and barely rises above 20% throughout the length of the eclipse track over the rest of France. However, beyond this point the maritime character of the westerly winds becomes more continental, and sunshine frequency begins a slow climb through Germany and Austria. This improving trend is interrupted briefly through central Austria where a branch of the Alps reaches down to the eclipse track and brings an increase in cloudiness. The effect is quite dramatic – the number of hours of sunshine in summer drops nearly 25% across middle of Austria in comparison with its eastern and western borders.

Figure 22 and Table 38 show that the mean August cloudiness decreases from 60% near Cornwall to 50% near Paris. The mean number of hours of sunshine grows from 6.5 over France to just over 8 hours in Austria, with a matching increase in the frequency of scattered cloud conditions. Figure 23 shows little difference in the prospects between Land's End and Paris, but then a slowly growing probability of seeing the eclipse through Germany and Austria.

There is not much to recommend one part of the track over another other than to suggest that eclipse viewers head eastward to take advantage of the slow climatological improvement. A much more reliable option is to simply await the weather forecasts in the days ahead of the eclipse and pick a site which is forecast to be sunny. While long range forecasts are available out to ten days, they do not become particularly reliable until about five days in advance of the event. More will be said about this later.

Summing up then, and extracting the smallest details from the statistical record, the best sites along the center line in France will be found from Compiègne, past Reims, to Metz, though the advantage gained is very small. In Germany the most suitable climatology is found from Ulm past Munich to the Austrian border. In Austria, climatology favors a location near the Hungarian border south of Vienna, though a location near the German border comes in as a close second.

EASTERN EUROPE TO THE BLACK SEA

As the track leaves Austria, it also draws away from the influence of the westerly winds which have controlled the meteorology up to this point. In summer, the Danubian plains over southern Hungary are affected more by the Mediterranean climate advancing northward from the Adriatic Sea than from the march of Atlantic disturbances. The pronounced effect on cloud cover and the amount of sunshine leaves little doubt that the best European eclipse conditions will be found in Hungary, Romania and Bulgaria.

For the most part, the path follows the lowlands along the Danube River and is protected from stronger weather systems by the Carpathian Mountains in the north and the Balkan Mountains in the south. Prevailing winds blow lightly from the north or northwest, being drawn into a large low pressure system which forms over Iran in the summer. These etesian winds bring dry invigorating air which is constant in direction and speed. Precipitation is mostly in the form of showers and thundershowers, and tends to be greatest where the winds blow upslope – generally on the northern slopes of the Balkan Mountains.

Figure 22 shows that the eclipse track begins to cross the cloud isopleths at a sharper angle eastward from Hungary, with the result that the mean August cloud amount has dropped to about 45% at the shores of the Black Sea. The amount of sunshine climbs above ten hours per day, more than 70% of the maximum possible. The number of days with scattered cloud or less at eclipse time rises from about half the month near the Austrian border to nearly two-thirds over Bulgaria. The probability of seeing the eclipse reaches 63% at the Black Sea ports of Varna and Constanta, popular summer destinations with beaches and fine Roman ruins to attract visitors and eclipse-seekers.

Thunderstorms have a relatively high frequency of occurrence in eastern Europe and bring considerable cloudiness when present. But this convective cloud relies on the heating of the ground for much of its development, unless pushed by cold fronts or other weather disturbances. Because of this, cloud cover statistics tend to underestimate the chances of seeing an eclipse which takes place in the morning hours or in the early afternoon. The maximum time of thunderstorm development occurs at 6 PM local time, well after the noon hour date for the eclipse. Indeed, the cooling associated with the eclipse may delay the onset of convection for another hour or two.

The path crosses Hungary's Lake Balaton, a popular resort area of warm water and sandy beaches. At one time the border between the Ottoman and Hapsburg empires ran down the center of the lake, and a number of ruined castles dot the northern hills above the lake. The location is easily reached from Budapest.

According to Dr. Szécsényi-Nagy of Loránd Eötvös University, the weather in Hungary is normally stable in summer with occasional long periods (3-4 weeks) of high pressure, cloudless skies and a dry atmosphere. Dr. Szécsényi-Nagy studied 20 years of weather records for the week on either side of the eclipse date and noted that only three days of the 300 were without sunshine at stations along the track. According to his results, sunshine is the overwhelming character of the day, with nearly two thirds of the records showing more than ten hours each day.

His conclusions are mirrored by the statistics for Keszthely ("cast-eye"), a resort and service town on the western shore of Lake Balaton. The 14 days of the month in which scattered cloud is reported at eclipse time at this location is the best in Hungary and comparable with sites in Romania and Bulgaria. Lake Balaton seems to have more than its share of sunny weather, likely a result of the protection afforded by the surrounding heights. The eclipse viewing probabilities in Figure 23 do not show this advantage, but this is probably caused by the smoothing of the climatological record of cloud cover by the computer model. The real differences are best seen in the actual record of observations.

Just after its point of maximum eclipse, the Moon's shadow crosses Bucharest, the capital of Romania. This city of two million promises to be a prime eclipse-viewing site, in part because of the comfort and ease of access, and partly because of the excellent weather prospects and the long eclipse duration. Since the center line neatly bisects the city, eclipse-viewing can be done from the wide boulevards or one of the many city parks.

The sum of all of the climatological measures points to Hungary, Bulgaria and Romania as the choicest locations in Europe for viewing this event. The shores of the Black Sea offer the greatest prospects of success, with a generous sunny climate enhanced by sea breeze circulations. Other suitable spots can be found all along the track from Lake Balaton eastward. To get any better weather prospects, the eclipse location will have to move to the other side of the Black Sea.

TURKEY, IRAQ AND IRAN

Once across the Black Sea, the lunar shadow's eastward track moves into the best weather conditions. Weather in the region is dominated by an extension of the large monsoon low over India and Pakistan.

While low pressure systems are normally associated with cloud and rain, this particular low lies beneath an upper level high which suppresses the formation of cloud. A weak frontal system extends from northern Iran to the Mediterranean coast, separating moderate temperatures and moisture on the north side from the semi-arid airmasses to the south. Upper level winds tend to flow from the west in northern Turkey and from the east in more southerly parts along the track.

The narrow coastal plain along the Black Sea coast of Turkey has a Mediterranean climate and represents the airmasses north of the front. This area has sufficient moisture to grow figs, olives, tea and tobacco. Most of the rain falls in winter, but steady northwest winds in the summer season bring occasional convective clouds with showers and thundershowers.

The eclipse comes ashore at a remote part of the Turkish coastline near the town of Cide. To the east of the shadow path is Sinop, an ancient city of Greek and Byzantine origins. Its most famous native son was Diogenes the Cynic, who is reputed to have replied to Alexander the Great when asked what Alexander could do for him, "Yes, stand aside, you're blocking my light," – a perfect quotation (though not sentiment) to go with a solar eclipse [Ayliffe et al. 1994]. The scenic road to the center line winds closely along the coast, forced to the edge of the sea by a range of 2000 meter mountains which line the Black Sea coast. Cide, a few kilometers inland from the sea, has a 10 km long pebbly beach nearby which could provide a site to watch the eclipse. Travel time from Sinop to Cide is about two hours. West of the center line is the regional capital Zonguldak, slightly closer than Sinop, and with good transportation connections to Ankara.

The steady onshore winds along the coast promote the development of convective cloudiness as they rise up the mountain slopes which line the sea, but cloud statistics for Zonguldak and Sinop promise weather at least as favorable as on the Bulgarian coast. Farther south, the track reaches Sivas on the eastern edge of the Anatolian Plateau. This city of magnificent Selçuk monuments has a history extending from Hittite times to the formation of modern Turkey. The frequency of days with scattered cloud rises rapidly here as downslope winds from the surrounding mountains dry the air. Much of the meager cloudiness is due to occasional summer thundershowers, though the incidence of rainy days is dwindling rapidly. Nevertheless, some of these thunderstorms can bring violent weather with widespread cloudiness. As in eastern Europe, they are primarily an afternoon event, but because the eclipse is later in Turkey and the heating of the ground more pronounced, thunderstorms are more likely here than in Europe at eclipse time. Beyond Sivas the shadow bounces across the eastern limb of the Taurus Mountains, with gradually improving prospects for good eclipse weather.

When the shadow path finally departs the Taurus Mountains it descends into the Tigris Valley and reaches Dyrbakir, a sprawling city with a large Kurdish population. The political turmoil of the area has converted Dyrbakir into a large military outpost, enhanced by the presence of a large NATO air base with shrieking jets and thumping helicopter rotors. Weather prospects are excellent (in spite of the anomalous "% of Possible Sunshine" in Table 38) as August is the hottest and driest month of the year. Statistics for the area show that only scattered clouds can be found at eclipse time on 24 to 28 days of the month, and nearly 30 days at the border with Syria. The viewing probability in Figure 23 rises above 80% as thunderstorms almost cease to be a threat, and the eclipse track moves into its most promising arena.

Mountainous terrain intervenes again as the eclipse path moves across Kurdistan into Iraq. It is a volatile area and best left alone, in spite of the clear skies. Travelers here must contend with a region in turmoil and will have to have a special quest for adventure. Even better weather and a more stable political climate comes in Iran where the track leaves the Zagros Mountains and moves onto the desert plateau. This is an arid area with temperatures approaching 50° C. The low humidity makes the heat somewhat more bearable (as will the cooling with an eclipse), but by the time the track reaches the Pakistan border, sultry humidities from the Gulf of Oman promise less pleasant observing conditions.

Skies are nearly cloudless over Iran for the most part, except for the occasional patches of scattered convective clouds and light showers. These are usually weak enough that they will succumb readily to the cooling which comes with the eclipse. Figure 23 shows that the apex of the eclipse viewing prospects is reached at Esfahan in southern Iran. Sites here have a 96% chance of seeing the Sun on August 11.

PAKISTAN AND INDIA

In August the southwest monsoon season over India is just past its peak, though still in full sway. This humid, wet and cloudy season does not retreat from eastern Pakistan until early September, too late to affect the eclipse. Cloud conditions over western Pakistan are heavier than over Iran because of the abundant moisture available from the Arabian Sea and the convection-promoting influence of the monsoon. Statistics for Karachi show a dramatic decline in the frequency of days with scattered cloud, though not in the number of days with rain.

As the track moves into India, the low Sun angle, late hour, and the extensive cloudiness of the Indian monsoon bring the poorest conditions of anywhere along the track. Satellite pictures reveal a land cloaked in cloud day after day and the probability of seeing the eclipse declines rapidly. By the time the shadow reaches the sunset terminator, the chances of seeing the eclipse have dropped almost to zero.

THE ATLANTIC OCEAN AND THE BLACK SEA

The proximity of the sunrise eclipse to Nova Scotia will undoubtedly tempt some observers to try a ship-board expedition from the eastern seaboard of North America. Though skies have a high frequency of cloud cover, the mobility offered by a ship should be able to overcome this deficiency, to some extent, provided good weather advice is available. The very low Sun angle at the start of the eclipse will seriously impede the search for a hole in any cloud cover which might be there, but provided the excursion is not just a day trip with little time for exploration, the effort has a good chance of being rewarded.

Black Sea prospects are comparable to nearby lands enjoying favorable climate statistics and the diminished influence of the variable westerlies. Mobility offers advantages similar to those described above.

Mean wave heights in the western Atlantic range between 1 and 1.5 meters off of the Nova Scotia coast, making time-exposure photography a little challenging, particularly through a telescope. Waves on the much smaller Black Sea tend to keep under 0.5 meters except near the Turkish coast where the prevailing winds have the entire length of the sea to build wave heights.

COPING WITH THE WEATHER ON ECLIPSE DAY

From England to Romania and Bulgaria, eclipse day will be subject to the vagaries of the westerly winds which carry highs and lows across the European continent. While climatology offers some advice for long range planning, mobile eclipse observers will almost certainly be able to find a view of the Sun on August 11. Particularly in western Europe, where the collection of climate statistics over the decades does not favor one location over another, the ability to respond to short and long-range weather forecasts will be very helpful. Modern meteorology is capable of producing detailed computer forecasts which stretch for ten days into the future. These forecasts can be assessed early enough to make general travel plans which can then be refined as eclipse day approaches.

Long range (5-10 day) computer forecasts are notoriously inaccurate and must be used with care. They should be used only for long range planning, to pick a general area for travel, as they are likely to show many changes before August 11 arrives. The best sites will lie beneath an upper level ridge ahead of, or in, a surface high pressure region, if these structures are forecast. This does not guarantee good weather, but it does greatly enhance one's prospects. If the numerical models don't predict these conditions, then more sophisticated decisions must be made which will probably require the services of a meteorologist.

Computer forecasts become much more reliable within five days, and still more accurate within 48 hours. The amount of detail they reveal also increases, with fields of relative humidity, precipitation, cloud cover, winds and temperature becoming available as the critical date approaches. A logical approach would be to select two or three specific sites within a convenient travel distance at this time, and then make a decision between them at the last possible moment.

Long range forecasts for Europe from the National Weather Service (NWS) in the United States and the European Centre for Medium Range Weather Forecasting (ECMWF) are readily available on the World Wide Web from Purdue University and a number of other locations. At Purdue the U.S. medium range forecast (MRF) model is usually available each morning after 12:30 UT. The ECMWF forecast out to six days is also available at the same site.

MRF charts show the 500 millibar flow (about 5000 m above the surface) and the surface pressure pattern. Look for an upper level ridge on the 500 mb chart, and the position of highs and lows on the other. These two charts will allow an early evaluation of the weather over Europe (most models are global, though not all parts of the globe are shown in the Web sites). The really hard part is whether or not the models are trustworthy, especially eight or ten days into the future. What clues are available to indicate that the numerical weather patterns will match the real ones on eclipse day?

Consistency is one clue – is the same general pattern forecast from one day to the next? Another is whether or not two different models forecast the same general patterns. Most likely there will be small differences between the two, but if highs and lows are hundreds of kilometers apart, or upper features don't line up, then use the charts with caution. Try to find a spot which looks good on both charts, and wait another day to see if the agreement improves. Don't finalize long range plans at this stage.

Once the eclipse is within five days, the greater reliability and detail in the models allows serious planning, for now we know not only the location of weather systems, but also how the cloud is draped around them. But be careful! Predictions are not gospel, and thunderstorms especially are difficult to predict with accuracy when more than a day or two away. Low level cloud is often unforecast, or indicated only in a subtle variation in the humidity pattern. Keep looking for the best spot within your travel range according to the position of the low and the upper ridge. By now of course, normal meteorological forecasts will be available and you may simply have to watch the television.

Long range forecasts over Turkey and the Middle East are not likely to be as informative as those over Europe, where weather systems are more active and changeable. Individuals and groups who are not mobile may wish to consider sites where climate statistics work in their favor, from Hungary eastward through Turkey, rather than take their chances with the westerlies in western Europe or England.

THE PROBABILITY OF SEEING THE ECLIPSE

When the Sun is high, the probability of seeing an eclipse depends only on the proportion of the sky which is covered by cloud, a factor which is represented by the mean cloudiness of a site. As the altitude of the Sun declines however, the depth of the cloud becomes more important as a restriction to visibility. This is easily understood by considering the effects of a towering thunderstorm – if overhead, the blocking capacity depends on the area of the base of the storm. If the thunderstorm is on the horizon, the blocking effect depends on the width of the storm perpendicular to the line of sight and its height.

If the average height of clouds in an area and the mean cloud amount are known, then the probability of seeing the eclipse can be calculated for a given solar altitude. Such a calculation is shown in Figure 23, using climatological values for these parameters. Adjustments are made for the time of day, since a large part of the cloud cover along the track is convective in nature, and thus dependent on the hour.

As with any calculation, the results of Figure 23 should be used cautiously. The climatological data used in the modeling are smoothed out from their true values, and small scale variations are lost or muted. Actual forecasts on and ahead of eclipse day will provide much more information than this figure, though of course they do not permit planning years ahead of time.

SUMMARY

Weather prospects for the eclipse begin in the dismal cloudiness of the North Atlantic and improve steadily along the eclipse track as far as the Middle East. Beyond Iran, cloud cover thickens again and the eclipse ends as it began with diminished promises. The best prospects in Europe are found along the shores of the Black Sea. Iran offers the most favorable weather along the entire track, though cautious observers may prefer the skies of central Turkey.

WEATHER WEB SITES

1. <http://www.tvweather.com>

A good starting point. This site has many links to current and past weather around the world.

2. http://shark1.esrin.esa.it:80/q_interface.html

A source of past satellite pictures around the world from 1992 onward (with gaps). These color pictures will give a feel for the weather over Europe and the Middle East in mid August.

3. <http://wxp.atms.purdue.edu/>

Purdue University, where the various model forecasts can be obtained. The MRF model is available over Europe for 10 days into the future. The ECMWF forecast can be examined for 6 days ahead. The MRF is available as a 9 panel display of 500 mb and surface features. Both models go as far east as Iraq. Another model, the AVN is available for 72 hours into the future and includes a chart of relative humidity (where 70% RH or higher implies cloud).

4. <http://www.meteo.fr/tpsreel/images/satt0.jpg>

A current European satellite image is available at Meteo France, as well as other sites.

OBSERVING THE ECLIPSE

EYE SAFETY AND SOLAR ECLIPSES

B. Ralph Chou, MSc, OD
Associate Professor, School of Optometry, University of Waterloo
Waterloo, Ontario, Canada N2L 3G1

A total solar eclipse is probably the most spectacular astronomical event that most people will experience in their lives. There is a great deal of interest in watching eclipses, and thousands of astronomers (both amateur and professional) travel around the world to observe and photograph them.

A solar eclipse offers students a unique opportunity to see a natural phenomenon that illustrates the basic principles of mathematics and science that are taught through elementary and secondary school. Indeed, many scientists (including astronomers!) have been inspired to study science as a result of seeing a total solar eclipse. Teachers can use eclipses to show how the laws of motion and the mathematics of orbital motion can predict the occurrence of eclipses. The use of pinhole cameras and telescopes or binoculars to observe an eclipse leads to an understanding of the optics of these devices. The rise and fall of environmental light levels during an eclipse illustrate the principles of radiometry and photometry, while biology classes can observe the associated behavior of plants and animals. It is also an opportunity for children of school age to contribute actively to scientific research - observations of contact timings at different locations along the eclipse path are useful in refining our knowledge of the orbital motions of the Moon and earth, and sketches and photographs of the solar corona can be used to build a three-dimensional picture of the Sun's extended atmosphere during the eclipse.

However, observing the Sun can be dangerous if you do not take the proper precautions. The solar radiation that reaches the surface of Earth ranges from ultraviolet (UV) radiation at wavelengths longer than 290 nm to radio waves in the meter range. The tissues in the eye transmit a substantial part of the radiation between 380 and 1400 nm to the light-sensitive retina at the back of the eye. While environmental exposure to UV radiation is known to contribute to the accelerated aging of the outer layers of the eye and the development of cataracts, the concern over improper viewing of the Sun during an eclipse is for the development of "eclipse blindness" or retinal burns.

Exposure of the retina to intense visible light causes damage to its light-sensitive rod and cone cells. The light triggers a series of complex chemical reactions within the cells which damages their ability to respond to a visual stimulus, and in extreme cases, can destroy them. The result is a loss of visual function which may be either temporary or permanent, depending on the severity of the damage. When a person looks repeatedly or for a long time at the Sun without proper protection for the eyes, this photochemical retinal damage may be accompanied by a thermal injury - the high level of visible and near-infrared radiation causes heating that literally cooks the exposed tissue. This thermal injury or photocoagulation destroys the rods and cones, creating a small blind area. The danger to vision is significant because photic retinal injuries occur without any feeling of pain (there are no pain receptors in the retina), and the visual effects do not occur for at least several hours after the damage is done [Pitts, 1993].

The only time that the Sun can be viewed safely with the naked eye is during a total eclipse, when the Moon completely covers the disk of the Sun. ***It is never safe to look at a partial or annular eclipse, or the partial phases of a total solar eclipse, without the proper equipment and techniques.*** Even when 99% of the Sun's surface (the photosphere) is obscured during the partial phases of a solar eclipse, the remaining crescent Sun is still intense enough to cause a retinal burn, even though illumination levels are comparable to twilight [Chou, 1981, 1996; Marsh, 1982]. Failure to use proper observing methods may result in permanent eye damage or severe visual loss. This can have important adverse effects on career choices and earning potential, since it has been shown that most individuals who sustain eclipse-related eye injuries are children and young adults [Penner and McNair, 1966; Chou and Krailo, 1981].

The same techniques for observing the Sun outside of eclipses are used to view and photograph annular solar eclipses and the partly eclipsed Sun [Sherrod, 1981; Pasachoff & Menzel 1992; Pasachoff & Covington, 1993; Reynolds & Sweetsir, 1995]. The safest and most inexpensive method is by projection. A pinhole or small opening is used to form an image of the Sun on a screen placed about a meter behind the opening. Multiple openings in perforboard, in a loosely woven straw hat, or even between interlaced fingers can be used to cast a pattern of solar images on a screen. A similar effect is seen on the ground below a broad-leaved tree: the many "pinholes" formed by overlapping leaves creates hundreds of crescent-shaped images.

Binoculars or a small telescope mounted on a tripod can also be used to project a magnified image of the Sun onto a white card. All of these methods can be used to provide a safe view of the partial phases of an eclipse to a group of observers, but care must be taken to ensure that no one looks through the device. The main advantage of the projection methods is that nobody is looking directly at the Sun. The disadvantage of the pinhole method is that the screen must be placed at least a meter behind the opening to get a solar image that is large enough to see easily.

The Sun can only be viewed directly when filters specially designed to protect the eyes are used. Most such filters have a thin layer of chromium alloy or aluminum deposited on their surfaces that attenuates both visible and near-infrared radiation. A safe solar filter should transmit less than 0.003% (density~4.5)¹⁰ of visible light (380 to 780 nm) and no more than 0.5% (density~2.3) of the near-infrared radiation (780 to 1400 nm). Figure 24 shows the spectral response for a selection of safe solar filters.

One of the most widely available filters for safe solar viewing is shade number 14 welder's glass, which can be obtained from welding supply outlets. A popular inexpensive alternative is aluminized mylar manufactured specifically for solar observation. ("Space blankets" and aluminized mylar used in gardening are *not* suitable for this purpose!) Unlike the welding glass, mylar can be cut to fit any viewing device, and doesn't break when dropped. Many experienced solar observers use one or two layers of black-and-white film that has been fully exposed to light and developed to maximum density. The metallic silver contained in the film emulsion is the protective filter. Some of the newer black and white films use dyes instead of silver and these are *unsafe*. Black-and-white negatives with images on it (e.g., medical x-rays) are also *not* suitable. More recently, solar observers have used floppy disks and compact disks (both CDs and CD-ROMs) as protective filters by covering the central openings and looking through the disk media. However, the optical quality of the solar image formed by a floppy disk or CD is relatively poor compared to mylar or welder's glass. Some CDs are made with very thin aluminum coatings which are not safe - if you can see through the CD in normal room lighting, don't use it!! No filter should be used with an optical device (e.g. binoculars, telescope, camera) unless it has been specifically designed for that purpose and is mounted at the front end (i.e., end towards the Sun). Some sources of solar filters are listed in the following section.

Unsafe filters include all color film, black-and-white film that contains no silver, photographic negatives with images on them (x-rays and snapshots), smoked glass, sunglasses (single or multiple pairs), photographic neutral density filters and polarizing filters. Most of these transmit high levels of invisible infrared radiation which can cause a thermal retinal burn (see Figure 24). The fact that the Sun appears dim, or that you feel no discomfort when looking at the Sun through the filter, is no guarantee that your eyes are safe. Solar filters designed to thread into eyepieces that are often provided with inexpensive telescopes are also unsafe. These glass filters can crack unexpectedly from overheating when the telescope is pointed at the Sun, and retinal damage can occur faster than the observer can move the eye from the eyepiece. Avoid unnecessary risks. Your local planetarium, science center, or amateur astronomy club can provide additional information on how to observe the eclipse safely.

There has been concern expressed about the possibility that UVA radiation (wavelengths between 315 and 380 nm) in sunlight may also adversely affect the retina [Del Priore, 1991]. While there is some experimental evidence for this, it only applies to the special case of aphakia, where the natural lens of the eye has been removed because of cataract or injury, and no UV-blocking spectacle, contact or intraocular lens has been fitted. In an intact normal human eye, UVA radiation does not reach the retina because it is absorbed by the crystalline lens. In aphakia, normal environmental exposure to solar UV radiation may indeed cause chronic retinal damage. However, the solar filter materials discussed in this article attenuate solar UV radiation to a level well below the minimum permissible occupational exposure for UVA (ACGIH, 1994), so an aphakic observer is at no additional risk of retinal damage when looking at the Sun through a proper solar filter.

In the days and weeks preceding a solar eclipse, there are often news stories and announcements in the media, warning about the dangers of looking at the eclipse. Unfortunately, despite the good intentions behind these messages, they frequently contain misinformation, and may be designed to scare people from seeing the eclipse at all. However, this tactic may backfire, particularly when the messages are intended for students. A student who heeds warnings from teachers and other authorities not to view the eclipse because of the danger to vision, and learns later that other students did see it safely, may feel cheated out of the experience. Having now learned that the authority figure was wrong on one occasion, how is this student

¹⁰ In addition to the term transmittance (in percent), the energy transmission of a filter can also be described by the term density (unitless) where density 'd' is the common logarithm of the reciprocal of transmittance 't' or $d = \log_{10}[1/t]$. A density of '0' corresponds to a transmittance of 100%; a density of '1' corresponds to a transmittance of 10%; a density of '2' corresponds to a transmittance of 1%, etc....

going to react when other health-related advice about drugs, alcohol, AIDS, or smoking is given [Pasachoff, 1997]? Misinformation may be just as bad, if not worse than no information at all.

In spite of these precautions, the *total* phase of an eclipse can and should be viewed without any filters whatsoever. The naked eye view of totality is not only completely safe, it is truly and overwhelmingly awe-inspiring!

SOURCES FOR SOLAR FILTERS

The following is a brief list of sources for mylar and/or glass filters specifically designed for safe solar viewing with or without a telescope. The list is not meant to be exhaustive, but is simply a representative sample of sources for solar filters currently available in North America and Europe. For additional sources, see advertisements in *Astronomy* and/or *Sky & Telescope* magazines. The inclusion of any source on this list does not imply an endorsement of that source by the authors or NASA.

- ABELexpress - Astronomy Division, 230-Y E. Main St., Carnegie, PA 15106. (412) 279-0672
- Celestron International, 2835 Columbia St., Torrance, CA 90503. (310) 328-9560
- Edwin Hirsch, 29 Lakeview Dr., Tomkins Cove, NY 10986. (914) 786-3738
- Meade Instruments Corporation, 16542 Millikan Ave., Irvine, CA 92714. (714) 756-2291
- Orion Telescope Center, 2450 17th Ave., PO Box 1158-S, Santa Cruz, CA 95061. (408) 464-0446
- Pocono Mountain Optics, 104 NP 502 Plaza, Moscow, PA 18444. (717) 842-1500
- Rainbow Symphony, Inc., 6860 Canby Ave., #120, Reseda, CA 91335 (800) 821-5122
- Roger W. Tuthill, Inc., 11 Tanglewood Lane, Mountainside, NJ 07092. (908) 232-1786
- Telescope and Binocular Center, P.O. Box 1815, Santa Cruz, CA 95061-1815. (408) 763-7030
- Thousand Oaks Optical, Box 5044-289, Thousand Oaks, CA 91359. (805) 491-3642
- Khan Scope Centre, 3243 Dufferin Street, Toronto, Ontario, Canada M6A 2T2(416) 783-4140
- Perceptor Telescopes TransCanada, Brownsville Junction Plaza, Box 38,
Schomberg, Ontario, Canada L0G 1T0 (905) 939-2313
- Eclipse 99 Ltd., Belle Etoile, Rue du Hamel, Guernsey GY5 7QJ. 001 44 1481 64847

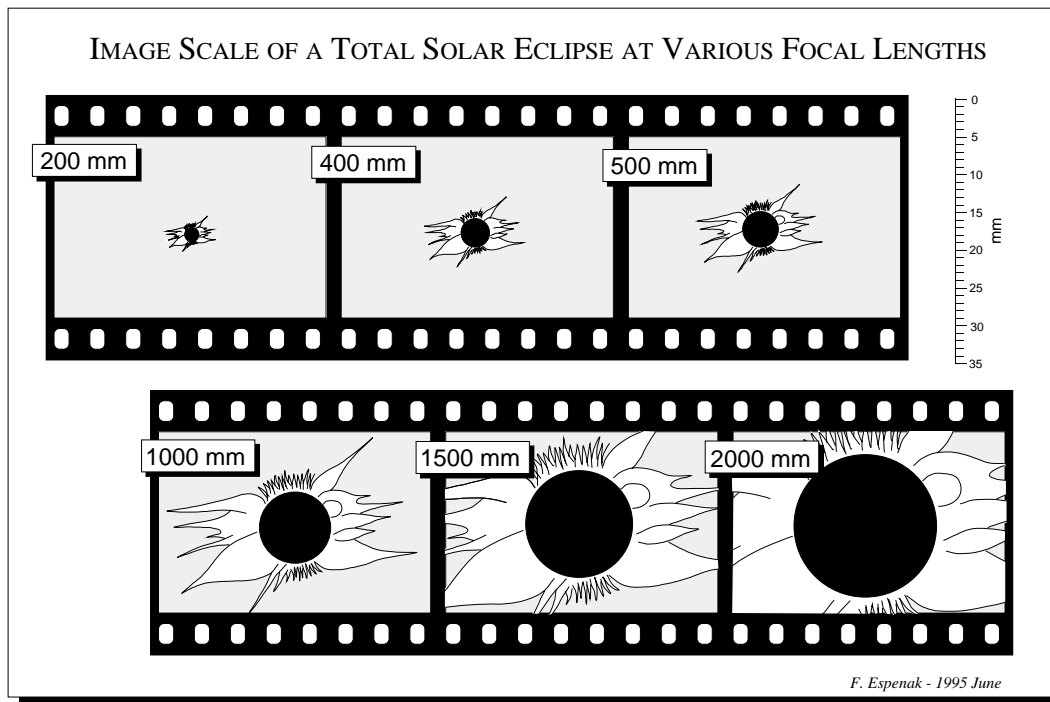
IAU SOLAR ECLIPSE EDUCATION COMMITTEE

In order to ensure that astronomers and public health authorities have access to information related to safe viewing practices, the International Astronomical Union, the international organization for professional astronomers, set up a Solar Eclipse Education Committee. Under Prof. Jay M. Pasachoff of Williams College, the Committee has assembled information on safe methods of observing the Sun and solar eclipses, eclipse-related eye injuries, and samples of educational materials on solar eclipses.

For more information, contact Prof. Jay M. Pasachoff, Hopkins Observatory, Williams College, Williamstown, MA 01267, USA (e-mail: jay.m.pasachoff@williams.edu). Information on safe solar filters can be obtained by contacting Dr. B. Ralph Chou (e-mail: bchou@sciborg.uwaterloo.ca).

ECLIPSE PHOTOGRAPHY

The eclipse may be safely photographed provided that the above precautions are followed. Almost any kind of camera with manual controls can be used to capture this rare event. However, a lens with a fairly long focal length is recommended to produce as large an image of the Sun as possible. A standard 50 mm lens yields a minuscule 0.5 mm image, while a 200 mm telephoto or zoom produces a 1.9 mm image. A better choice would be one of the small, compact catadioptric or mirror lenses that have become widely available in the past ten years. The focal length of 500 mm is most common among such mirror lenses and yields a solar image of 4.6 mm. With one solar radius of corona on either side, an eclipse view during totality will cover 9.2 mm. Adding a 2x tele-converter will produce a 1000 mm focal length, which doubles the Sun's size to 9.2 mm. Focal lengths in excess of 1000 mm usually fall within the realm of amateur telescopes. If full disk photography of partial phases on 35 mm format is planned, the focal length of the optics must not exceed 2600 mm. However, since most cameras don't show the full extent of the image in their viewfinders, a more practical limit is about 2000 mm. Longer focal lengths permit photography of only a magnified portion of the Sun's disk. In order to photograph the Sun's corona during totality, the focal length should be no longer than 1500 mm to 1800 mm (for 35 mm equipment). However, a focal length of 1000 mm requires less critical framing and can capture some of the longer coronal streamers. For any particular focal length, the diameter of the Sun's image is approximately equal to the focal length divided by 109 (Table 25).



A mylar or glass solar filter must be used on the lens throughout the partial phases for both photography and safe viewing. Such filters are most easily obtained through manufacturers and dealers listed in *Sky & Telescope* and *Astronomy* magazines (see: SOURCES FOR SOLAR FILTERS). These filters typically attenuate the Sun's visible and infrared energy by a factor of 100,000. However, the actual filter factor and choice of ISO film speed will play critical roles in determining the correct photographic exposure. A low to medium speed film is recommended (ISO 50 to 100) since the Sun gives off abundant light. The easiest method for determining the correct exposure is accomplished by running a calibration test on the uneclipsed Sun. Shoot a roll of film of the mid-day Sun at a fixed aperture (f/8 to f/16) using every shutter speed between 1/1000 and 1/4 second. After the film is developed, note the best exposures and use them to photograph all the partial phases. The Sun's surface brightness remains constant throughout the eclipse, so no exposure compensation is needed except for the crescent phases which require two more stops due to solar limb darkening. Bracketing by several stops is also necessary if haze or clouds interfere on eclipse day.

Certainly the most spectacular and awe inspiring phase of the eclipse is totality. For a few brief minutes or seconds, the Sun's pearly white corona, red prominences and chromosphere are visible. The great

challenge is to obtain a set of photographs which captures some aspect of these fleeting phenomena. The most important point to remember is that during the total phase, all solar filters *must be removed!* The corona has a surface brightness a million times fainter than the photosphere, so photographs of the corona are made without a filter. Furthermore, it is completely safe to view the totally eclipsed Sun directly with the naked eye. No filters are needed and they will only hinder your view. The average brightness of the corona varies inversely with the distance from the Sun's limb. The inner corona is far brighter than the outer corona. Thus, no single exposure can capture its full dynamic range. The best strategy is to choose one aperture or f/number and bracket the exposures over a range of shutter speeds (i.e., 1/1000 down to 1 second). Rehearsing this sequence is highly recommended since great excitement accompanies totality and there is little time to think.

Exposure times for various combinations of film speeds (ISO), apertures (f/number) and solar features (chromosphere, prominences, inner, middle and outer corona) are summarized in Table 26. The table was developed from eclipse photographs made by Espenak as well as from photographs published in *Sky and Telescope*. To use the table, first select the ISO film speed in the upper left column. Next, move to the right to the desired aperture or f/number for the chosen ISO. The shutter speeds in that column may be used as starting points for photographing various features and phenomena tabulated in the 'Subject' column at the far left. For example, to photograph prominences using ISO 100 at f/11, the table recommends an exposure of 1/500. Alternatively, you can calculate the recommended shutter speed using the 'Q' factors tabulated along with the exposure formula at the bottom of Table 26. Keep in mind that these exposures are based on a clear sky and a corona of average brightness. You should bracket your exposures one or more stops to take into account the actual sky conditions and the variable nature of these phenomena.

Another interesting way to photograph the eclipse is to record its phases all on one frame. This is accomplished by using a stationary camera capable of making multiple exposures (check the camera instruction manual). Since the Sun moves through the sky at the rate of 15 degrees per hour, it slowly drifts through the field of view of any camera equipped with a normal focal length lens (i.e., 35 to 50 mm). If the camera is oriented so that the Sun drifts along the frame's diagonal, it will take over three hours for the Sun to cross the field of a 50 mm lens. The proper camera orientation can be determined through trial and error several days before the eclipse. This will also insure that no trees or buildings obscure the view during the eclipse. The Sun should be positioned along the eastern (left in the northern hemisphere) edge or corner of the viewfinder shortly before the eclipse begins. Exposures are then made throughout the eclipse at ~five minute intervals. The camera must remain perfectly rigid during this period and may be clamped to a wall or post since tripods are easily bumped. If you're in the path of totality, remove the solar filter during the total phase and take a long exposure (~1 second) in order to record the corona in your sequence. The final photograph will consist of a string of Suns, each showing a different phase of the eclipse.

Finally, an eclipse effect that is easily captured with point-and-shoot or automatic cameras should not be overlooked. Use a kitchen sieve or colander and allow its shadow to fall on a piece of white cardboard placed several feet away. The holes in the utensil act like pinhole cameras and each one projects its own image of the Sun. The effect can also be duplicated by forming a small aperture with one's hands and watching the ground below. The pinhole camera effect becomes more prominent with increasing eclipse magnitude. Virtually any camera can be used to photograph the phenomenon, but automatic cameras must have their flashes turned off since this would otherwise obliterate the pinhole images.

For those who choose to photograph this eclipse from one of the many cruise ships in the path, some special comments are in order. Shipboard photography puts certain limits on the focal length and shutter speeds that can be used. It's difficult to make specific recommendations since it depends on the stability of the ship as well as wave heights encountered on eclipse day. Certainly telescopes with focal lengths of 1000 mm or more can be ruled out since their small fields of view would require the ship to remain virtually motionless during totality, and this is rather unlikely even given calm seas. A 500 mm lens might be a safe upper limit in focal length. Film choice could be determined on eclipse day by viewing the Sun through the camera lens and noting the image motion due to the rolling sea. If it's a calm day, you might try an ISO 100 film. For rougher seas, ISO 400 or more might be a better choice. Shutter speeds as slow as 1/8 or 1/4 may be tried if the conditions warrant it. Otherwise, stick with a 1/15 or 1/30 and shoot a sequence through 1/1000 second. It might be good insurance to bring a wider 200 mm lens just in case the seas are rougher than expected. As worst case scenario, Espenak photographed the 1984 total eclipse aboard a 95 foot yacht in seas of 3 feet. He had to hold on with one hand and point his 350 mm lens with the other! Even at that short focal length, it was difficult to keep the Sun in the field. However, any large cruise ship will offer a far more stable platform than this.

For more information on eclipse photography, observations and eye safety, see FURTHER READING in the BIBLIOGRAPHY.

SKY AT TOTALITY

The total phase of an eclipse is accompanied by the onset of a rapidly darkening sky whose appearance resembles evening twilight about 30 to 40 minutes after sunset. The effect presents an excellent opportunity to view planets and bright stars in the daytime sky. Aside from the sheer novelty of it, such observations are useful in gauging the apparent sky brightness and transparency during totality. The Sun is in Pisces and all five naked eye planets as well as a number of bright stars will be above the horizon for observers within the umbral path. Figure 25 depicts the appearance of the sky during totality as seen from the center line at 11:00 UT. This corresponds to western Romania near the point of greatest eclipse.

Mercury ($m_v=+0.7$) and Venus ($m_v=-3.5$) are located 18° west and 15° east of the Sun, respectively, and both will be easily visible during totality. Venus is two months past inferior conjunction while Mercury is one month shy of superior conjunction. As the brightest planet in the sky, Venus can actually be observed in broad daylight provided that the sky is cloud free and of high transparency (i.e., no dust or particulates). Look for the planet during the partial phases by first covering the crescent Sun with an extended hand. Venus will be shining so brightly, it will be impossible to miss during totality. Mercury will prove much more challenging, but not too difficult if the sky transparency is good. Under the right circumstances, it should be possible to view all five classical planets, the Moon and the Sun (or at least its corona) as one's eyes sweep across the darkened sky during totality.

A number of the brightest winter/spring stars may also be visible during totality. Regulus ($m_v=+1.35$) is 10° east of the Sun while Castor ($m_v=+1.94$) and Pollux ($m_v=+1.14$) stand 31° and 28° to the northwest. Procyon ($m_v=+0.38$) and Sirius ($m_v=-1.46$) are located 30° and 52° to the southwest, respectively. Betelgeuse ($m_v=+0.5v$) and Rigel ($m_v=+0.12$) are low in the southwest at 26° and 38° , while Aldebaran ($m_v=+0.85$) is 20° above the western horizon. Capella ($m_v=+0.08$) lies 63° to the northwest. Finally, Arcturus ($m_v=+1.94$) lies due east 30° above the horizon.

The following ephemeris [using Bretagnon and Simon, 1986] gives the positions of the naked eye planets during the eclipse. *Delta* is the distance of the planet from Earth (A.U.'s), *V* is the apparent visual magnitude of the planet, and *Elong* gives the solar elongation or angle between the Sun and planet.

Ephemeris: 1999 Aug 11 11:00:00 UT				Equinox = Mean Date			
Planet	RA	Dec	Delta	V	Size "	Phase	Elong °
Sun	09h23m08s	+15°19'42"	1.01358	-26.7	1893.6	-	-
Mercury	08h07m35s	+18°08'56"	0.82062	0.7	8.2	0.30	18.3W
Venus	10h06m46s	+04°18'35"	0.30047	-3.5	55.5	0.04	15.4E
Mars	14h55m32s	-18°28'03"	1.06814	0.3	8.8	0.86	88.5E
Jupiter	02h11m31s	+11°47'36"	4.61411	-2.1	42.7	0.99	103.7W
Saturn	03h00m28s	+14°34'31"	9.13383	0.1	18.2	1.00	91.5W

For sky maps from other locations along the path of totality, see the special web site for the total solar eclipse of 1999: <http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>

CONTACT TIMINGS FROM THE PATH LIMITS

Precise timings of beading phenomena made near the northern and southern limits of the umbral path (i.e., the graze zones), are of value in determining the diameter of the Sun relative to the Moon at the time of the eclipse. Such measurements are essential to an ongoing project to monitor changes in the solar diameter. Due to the conspicuous nature of the eclipse phenomena and their strong dependence on geographical location, scientifically useful observations can be made with relatively modest equipment. A small telescope, short wave radio and portable camcorder are usually used to make such measurements. Time signals are broadcast via short wave stations WWV and CHU, and are recorded simultaneously as the eclipse is videotaped. If a video camera is not available, a tape recorder can be used to record time signals with verbal timings of each event. Inexperienced observers are cautioned to use great care in making such observations. The safest timing technique consists of observing a projection of the Sun rather than directly imaging the solar disk itself. The observer's geodetic coordinates are required and can be measured from USGS or other large scale maps. If a map is unavailable, then a detailed description of the observing site should be included which provides information such as distance and directions of the nearest towns/settlements, nearby landmarks, identifiable buildings and road intersections. The method of contact timing should be described in detail, along with an estimate of the error. The precisional requirements of these observations are ± 0.5 seconds in time, 1" (~ 30 meters) in latitude and longitude, and ± 20 meters (~ 60 feet) in elevation. Although GPS's (Global Positioning Satellite receivers) are commercially available ($\sim \$150$ US), their positional accuracy of ± 100 meters is about three times larger than the minimum accuracy required by grazing eclipse measurements. GPS receivers are also a useful source for accurate UT. The International Occultation Timing Association (IOTA) coordinates observers world-wide during each eclipse. For more information, contact:

Dr. David W. Dunham, IOTA
7006 Megan Lane
Greenbelt, MD 20770-3012, USA

E-mail: David_Dunham@jhuapl.edu
Phone: (301) 474-4722

Send reports containing graze observations, eclipse contact and Baily's bead timings, including those made anywhere near or in the path of totality or annularity to:

Dr. Alan D. Fiala
Orbital Mechanics Dept.
U. S. Naval Observatory
3450 Massachusetts Ave., NW
Washington, DC 20392-5420, USA

PLOTTING THE PATH ON MAPS

If high resolution maps of the umbral path are needed, the coordinates listed in Tables 7 and 8 are conveniently provided in longitude increments of 1° and $30'$ respectively to assist plotting by hand. The path coordinates in Table 3 define a line of maximum eclipse at five minute increments in Universal Time. If observations are to be made near the limits, then the grazing eclipse zones tabulated in Table 8 should be used. A higher resolution table of graze zone coordinates at longitude increments of $7.5'$ is available via a special web site for the 1999 total eclipse (<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>). Global Navigation Charts (1:5,000,000), Operational Navigation Charts (scale 1:1,000,000) and Tactical Pilotage Charts (1:500,000) of many parts of the world are published by the National Imagery and Mapping Agency (formerly known as Defense Mapping Agency). Sales and distribution of these maps are through the National Ocean Service (NOS). For specific information about map availability, purchase prices, and ordering instructions, contact the NOS at:

NOAA Distribution Division, N/ACC3
National Ocean Service
Riverdale, MD 20737-1199, USA

phone: 301-436-8301
FAX: 301-436-6829

It is also advisable to check the telephone directory for any map specialty stores in your city or metropolitan area. They often have large inventories of many maps available for immediate delivery.

ONC (Operational Navigation Charts) series maps have a larger scale (1:1,000,000) than GNC's appearing in this publication. However, their use here would serve to increase an already record size eclipse bulletin. Instead, we offer a list of ONC maps for plotting the path using data from tables 7 and 8. In particular, the path of totality crosses the following ONC charts:

ONC E-1	England
ONC E-2	France, Germany, Austria
ONC F-2	Austria, Hungary
ONC F-3	Romania, Turkey
ONC G-4	Turkey, Syria, Iraq
ONC G-6,H-7	Iran
ONC H-8	Pakistan, India
ONC J8, J-9	India

IAU WORKING GROUP ON ECLIPSES

Professional scientists are asked to send descriptions of their eclipse plans to the Working Group on Eclipses of the International Astronomical Union, so that they can keep a list of observations planned. Send such descriptions, even in preliminary form, to:

International Astronomical Union/Working Group on Eclipses
Prof. Jay M. Pasachoff, Chair
Williams College–Hopkins Observatory email: jay.m.pasachoff@williams.edu
Williamstown, MA 01267, USA FAX: (413) 597-3200

The members of the Working Group on Eclipses of Commissions 10 and 12 of the International Astronomical Union are: Jay M. Pasachoff (USA), Chair; F. Clette (Belgium), F. Espenak (USA); Iraida Kim (Russia); V. Rusin (Slovakia); Jagdev Singh (India); M. Stavinschi (Romania); Yoshinori Suematsu (Japan); consultant: J. Anderson (Canada).

NATO WORKSHOP PROCEEDINGS FOR THE 1999 TOTAL SOLAR ECLIPSE

In 1996 June 1-5, a special NATO Advanced Research Workshop for “Theoretical and Observational Problems Related to Solar Eclipses” was held in Sinaia, Romania. For the first time in the history of the eclipse observations, observers and theorists were brought together to present and discuss their projects for a future eclipse (1999). Scientific sessions during the meeting covered the following areas:

- Principal scientific results from the past eclipse observation.
- Small and large scale theoretical models of coronal structures.
- Low temperature structures in coronal environment.
- Specific problems of solar eclipse observations.
- Instrumental improvement for future observations.
- Tasks for Total Solar Eclipse of 11 August 1999.
- Public education at eclipses and eye safety.

The proceedings from the meeting will be published in 1997 June by Kluwer Academic Publishing (Netherlands) as part of the *NATO ASI Series* (Z. Mouradian and M. Stavinschi, eds.). For ordering information, please contact:

Dr. Zadig Mouradian email: mouradian@obspm.fr
Observatoire de Paris-Meudon FAX: +33.1.4507.7959
DASOP
92195 Meudon Principal
FRANCE

ROMANIAN PREPARATIONS FOR 1999 ECLIPSE

Romanian astronomers have established the International Association ECLIPSA'99 for the purpose of assisting both the scientific community and the general public. In addition to carrying out a series of scientific eclipse observations, ECLIPSA'99 will also play an important role in public education so that everyone can enjoy this extraordinary astronomical event.

We plan to set up a new telescope outside the Capital, to complete the observation and data bases of the three observatories of the Astronomical Institute of the Romanian Academy, as well as to build a great Planetarium at Bucharest Observatory, in the immediate vicinity of the Park "Charles the 1st".

Through a system of scholarships and awards, ECLIPSA'99 aims to specially train the staff necessary in the eclipse observation. Naturally, we will not leave out the amateur astronomers in view of the important contribution they have brought to the development of astronomy. Throughout the preparations, ECLIPSA'99 will carry out an ample program of national and international conferences and symposia, the publication of specialized and advertising materials, as well as of its own journal "Eclipsa"; it will also conduct an ample publicity campaign both at home and abroad.

To accomplish its aims ECLIPSA'99 is carrying on collaborations with similar institutions at home and abroad, with specialists in related fields, as well as with agencies and other institutions in the fields of culture, tourism, transport, trade, etc.. The funds of ECLIPSA'99 come from subscriptions, as well as from legacies and donations from home and abroad.

ECLIPSA'99 will not stop its activity after the eclipse. At that time, it will become the International Association ASTRONOMIA 21, whose goal will be to keep alive the interest in this old and, at the same time, modern science. For more information please contact:

Dr. Magdalena Stavinschi
Astr. Inst. of the Romanian Academy
Str. Cutitul de Argint 5
RO-75212 Bucharest
ROMANIA

email: magda@roastro.astro.ro
Phone: +40.1.336 36 87
FAX: +40.1.337 33 89

JOSO WORKING GROUP FOR 1999 ECLIPSE

In October 1995, the Joint Organization for Solar Observations (JOSO), a European consortium of observing solar physicists, created a new working group dedicated to the preparation of the 1999 eclipse. This Working Group (WG7) will prompt collaborations between scientific teams coming from European countries and other parts of the world to observe the eclipse, and the local scientific and academic organizations in the path of totality, by gathering and distributing information about existing resources and requirements for the practical organisation of scientific expeditions. Another WG7 project is to disseminate to the general public basic but reliable information about the 1999 event and safe eclipse viewing .

All groups who are planning to set up a scientific eclipse program are invited to join this new community. For more information, contact:

Dr. Frederic Clette - JOSO WG 7
Observatoire Royal de Belgique
Avenue Circulaire, 3
B-1180 Bruxelles
BELGIUM

email: fred@oma.be
FAX: +32.2.373.02.24

JOSO Web site : <http://joso.oat.ts.astro.it/>

ECLIPSE DATA ON INTERNET

NASA ECLIPSE BULLETINS ON INTERNET

To make the NASA solar eclipse bulletins accessible to as large an audience as possible, these publications are also available via the Internet. This was made possible through the efforts and expertise of Dr. Joe Gurman (GSFC/Solar Physics Branch). All future eclipse bulletins will be available via Internet.

NASA eclipse bulletins can be read or downloaded via the World-Wide Web using a Web browser (e.g.: Netscape, Microsoft Explorer, etc.) from the GSFC SDAC (Solar Data Analysis Center) Eclipse Information home page, or from top-level URL's for the currently available eclipse bulletins themselves:

http://umbra.nascom.nasa.gov/eclipse/	(SDAC Eclipse Information)
http://umbra.nascom.nasa.gov/eclipse/941103/rp.html	(1994 Nov 3)
http://umbra.nascom.nasa.gov/eclipse/951024/rp.html	(1995 Oct 24)
http://umbra.nascom.nasa.gov/eclipse/970309/rp.html	(1997 Mar 9)
http://umbra.nascom.nasa.gov/eclipse/980226/rp.html	(1998 Feb 26)
http://umbra.nascom.nasa.gov/eclipse/990811/rp.html	(1999 Aug 11)

The original Microsoft Word text files and PICT figures (Macintosh format) are also available via anonymous ftp. They are stored as BinHex-encoded, StuffIt-compressed Mac folders with .hqx suffixes. For PC's, the text is available in a zip-compressed format in files with the .zip suffix. There are three sub directories for figures (GIF format), maps (JPEG format), and tables (html tables, easily readable as plain text). For example, NASA RP 1344 (Total Solar Eclipse of 1995 October 24 [=951024]) has a directory for these files is as follows:

file://umbra.nascom.nasa.gov/pub/eclipse/951024/RP1344text.hqx	
file://umbra.nascom.nasa.gov/pub/eclipse/951024/RP1344PICTs.hqx	
file://umbra.nascom.nasa.gov/pub/eclipse/951024/ec951024.zip	
file://umbra.nascom.nasa.gov/pub/eclipse/951024/figures	(directory with GIF's)
file://umbra.nascom.nasa.gov/pub/eclipse/951024/maps	(directory with JPEG's)
file://umbra.nascom.nasa.gov/pub/eclipse/951024/tables	(directory with html's)

Other eclipse bulletins have a similar directory format.

Current plans call for making all future NASA eclipse bulletins available over the Internet, at or before publication of each. The primary goal is to make the bulletins available to as large an audience as possible. Thus, some figures or maps may not be at their optimum resolution or format. Comments and suggestions are actively solicited to fix problems and improve on compatibility and formats.

FUTURE ECLIPSE PATHS ON INTERNET

Presently, the NASA eclipse bulletins are published 24 to 36 months before each eclipse. However, there have been a growing number of requests for eclipse path data with an even greater lead time. To accommodate the demand, predictions have been generated for all central solar eclipses from 1995 through 2005 using the JPL DE/LE 200 ephemerides. All predictions use the Moon's center of mass; no corrections have been made to adjust for center of figure. The value used for the Moon's mean radius is $k=0.272281$. The umbral path characteristics have been predicted at 2 minute intervals of time compared to the 6 minute interval used in *Fifty Year Canon of Solar Eclipses: 1986-2035* [Espanak, 1987]. This should provide enough detail for making preliminary plots of the path on larger scale maps. Note that positive latitudes are north and positive longitudes are west. A list of currently available eclipse paths includes:

1998 February 26	– Total Solar Eclipse
1998 August 22	– Annular Solar Eclipse
1999 February 16	– Annular Solar Eclipse
1999 August 11	– Total Solar Eclipse
2001 June 21	– Total Solar Eclipse
2001 December 14	– Annular Solar Eclipse

- 2002 June 10 – Annular Solar Eclipse
- 2002 December 04 – Total Solar Eclipse
- 2003 May 31 – Annular Solar Eclipse
- 2003 November 23 – Total Solar Eclipse
- 2005 April 08 – Annular/Total Solar Eclipse
- 2005 October 03 – Annular Solar Eclipse

URL: <http://umbra.nascom.nasa.gov/eclipse/predictions/year-month-day.html>

The tables can be accessed through the SDAC Eclipse Information home page, or directly from the above URL. For example, the eclipse path of 1999 August 11 would use the above address with the string “year-month-day” replaced by “1999-august-11”. Send comments, corrections, suggestions or requests for more detailed ‘ftp’ instructions, to Fred Espenak via e-mail (espenak@lepvax.gsfc.nasa.gov). For Internet related problems, please contact Joe Gurman (gurman@uvsp.nascom.nasa.gov).

DOWNLOADING BULLETINS AND PATH TABLES VIA ANONYMOUS FTP

The eclipse bulletins and path tables are also available via anonymous ftp for sites which do not have access to the World Wide Web. A user first ftp’s to umbra.nascom.nasa.gov (150.144.30.134), using the username “anonymous” and password “<username>@<host>”. Note that the password is your e-mail address where <username> is your name and <host> is the fully qualified Internet address of your machine (e.g.- gurman@uvsp.nascom.nasa.gov). Next, you change directory with the command “cd pub/eclipse”.

There are five directories 941103, 951024, 970309, 980226, and 990811; one for each of the last five eclipse bulletins (1318, 1344, 1369, 1383, and 1398 respectively). In each, there is a flat ASCII README file and two .hqx files: RPnnntext.hqx and RPnnnnPICTS.hqx, where “nnnn” is the Reference Publication number. All .hqx files are BinHex-encoded (ASCII), StuffIt-compressed files for the Macintosh. There’s also one .zip file: ecymmdd.zip, where “ymmdd” is the date of the eclipse. This is a zip-compressed and encoded file for PC’s. There are also three subdirectories, figures, maps, and tables, with (respectively), the GIF figures, the JPEG GNC charts, and the html tables (easily readable as plain text). For example, the total solar eclipse of 970309 (= 1997 Mar 9) and published as NASA RP 1369 has a directory for these files is as follows:

```
file://umbra.nascom.nasa.gov/pub/eclipse/970309/README
file://umbra.nascom.nasa.gov/pub/eclipse/970309/RP1369text.hqx
file://umbra.nascom.nasa.gov/pub/eclipse/970309/RP1369PICTS.hqx
file://umbra.nascom.nasa.gov/pub/eclipse/970309/ec970309.zip
file://umbra.nascom.nasa.gov/pub/eclipse/970309/figures (directory with GIF's)
file://umbra.nascom.nasa.gov/pub/eclipse/970309/maps (directory with JPEG's)
file://umbra.nascom.nasa.gov/pub/eclipse/970309/tables (directory with html's)
```

Directories for analogous files for other solar eclipses are arranged similarly.

The html files should be downloaded in ASCII mode and the other files in binary (IMAGE) mode. If you are not using a Web viewer to access the ftp documents, you must first type either “ascii” or “binary” to download an ASCII or a binary file, respectively. You then download the file using the ftp protocol for your particular machine.

SPECIAL WEB SITE FOR 1999 SOLAR ECLIPSE

A special web site has been set up to supplement this bulletin with additional predictions, tables and data for the total solar eclipse of 1999. Some of the data posted there include an expanded version of Table 8 (Mapping Coordinates for the Zones of Grazing Eclipse), and local circumstance tables with many more cities as well as for astronomical observatories. Also featured will be higher resolution maps of selected sections of the path of totality and limb profile figures for a range of locations/times along the path. The URL of this special site is:

<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>

TOTAL SOLAR ECLIPSE OF 2001 JUNE 21

The next total eclipse of the Sun is the first one of the twenty-first century. The path of the Moon's umbral shadow begins in the South Atlantic, off the east coast of Uruguay and continues across the Atlantic where it reaches the west coast of Africa. The shadow enters Angola in the early afternoon with a center line duration of 4 ¹/₂ minutes (Figure 26). Traveling eastward, the path sweeps through Zambia, Zimbabwe and Mozambique. At that point, the central duration drops to three minutes with the late afternoon Sun 23° above the horizon. Swiftly crossing the Mozambique Channel, the path intercepts southern Madagascar where the central duration lasts 2 ¹/₂ minutes with a Sun altitude of 11°. The path ends two minutes later in the Indian Ocean.

Complete details will be published in the next NASA bulletin scheduled for Fall-Winter 1997.

PREDICTIONS FOR ECLIPSE EXPERIMENTS

This publication has attempted to provide comprehensive information on the 1999 total solar eclipse to both the professional and amateur/lay communities. However, certain investigations and eclipse experiments may require additional information which lies beyond the scope of this work. We invite the international professional community to contact us for assistance with any aspect of eclipse prediction including predictions for locations not included in this publication, or for more detailed predictions for a specific location (e.g.: lunar limb profile and limb corrected contact times for an observing site).

This service is offered for the 1998 eclipse as well as for previous eclipses in which analysis is still in progress. To discuss your needs and requirements, please contact Fred Espenak (espenak@lepvax.gsfc.nasa.gov).

ALGORITHMS, EPHEMERIDES AND PARAMETERS

Algorithms for the eclipse predictions were developed by Espenak primarily from the *Explanatory Supplement* [1974] with additional algorithms from Meeus, Grosjean and Vanderleen [1966] and Meeus [1982]. The solar and lunar ephemerides were generated from the JPL DE200 and LE200, respectively. All eclipse calculations were made using a value for the Moon's radius of $k=0.2722810$ for umbral contacts, and $k=0.2725076$ (adopted IAU value) for penumbral contacts. Center of mass coordinates were used except where noted. Extrapolating from 1996 to 1998, a value for ΔT of 64.6 seconds was used to convert the predictions from Terrestrial Dynamical Time to Universal Time. The international convention of presenting date and time in descending order has been used throughout the bulletin (i.e., *year, month, day, hour, minute, second*).

The primary source for geographic coordinates used in the local circumstances tables is *The New International Atlas* (Rand McNally, 1991). Elevations for major cities were taken from *Climates of the World* (U. S. Dept. of Commerce, 1972).

All eclipse predictions presented in this publication were generated on a Macintosh PowerPC 8500 computer. Word processing and page layout for the publication were done using Microsoft Word v5.1. Figures were annotated with Claris MacDraw Pro 1.5. Meteorological diagrams were prepared using Corel Draw 5.0 and converted to Macintosh compatible files. Finally, the bulletin was printed on a 600 dpi laser printer (Apple LaserWriter Pro).

The names and spellings of countries, cities and other geopolitical regions are not authoritative, nor do they imply any official recognition in status. Corrections to names, geographic coordinates and elevations are actively solicited in order to update the data base for future eclipses. All calculations, diagrams and opinions presented in this publication are those of the authors and they assume full responsibility for their accuracy.

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TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

FIGURES

Total Solar Eclipse of 1999 August 11

FIGURE 1: ORTHOGRAPHIC PROJECTION MAP OF THE ECLIPSE PATH

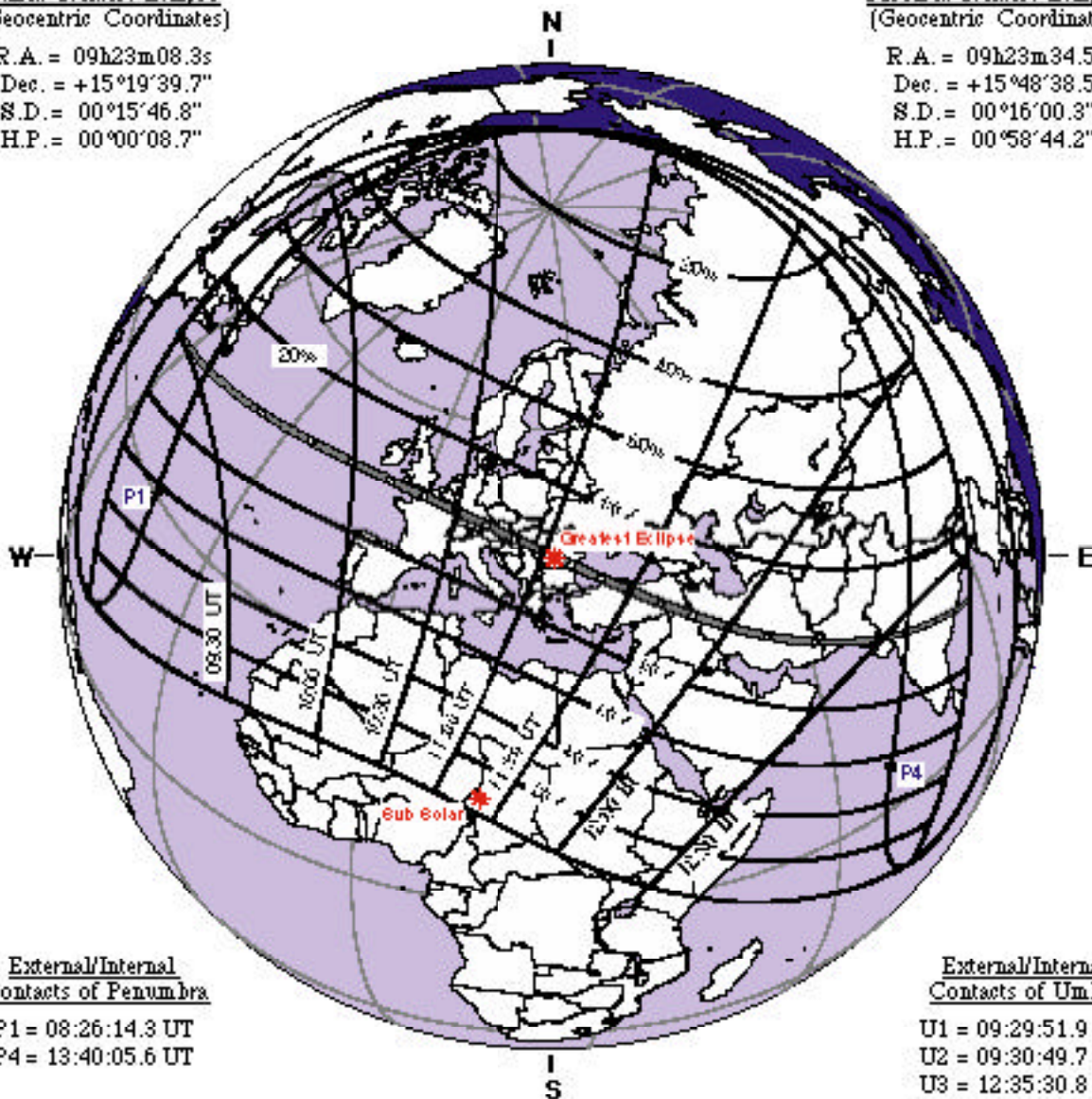
Geocentric Conjunction = 10:51:12.1 UT J.D. = 2451401.952223
 Greatest Eclipse = 11:03:04.4 UT J.D. = 2451401.960468
 Eclipse Magnitude = 1.02859 Gamma = 0.50623
 Saros Series = 145 Member = 21 of 77

Sun at Greatest Eclipse
 (Geocentric Coordinates)

R.A. = 09h23m08.3s
 Dec. = +15°19'39.7"
 S.D. = 00°15'46.8"
 H.P. = 00°00'08.7"

Moon at Greatest Eclipse
 (Geocentric Coordinates)

R.A. = 09h23m34.5s
 Dec. = +15°48'38.5"
 S.D. = 00°16'00.3"
 H.P. = 00°58'44.2"



External/Internal
Contacts of Penumbra

P1 = 08:26:14.3 UT
 P4 = 13:40:05.6 UT

External/Internal
Contacts of Umbra

U1 = 09:29:51.9 UT
 U2 = 09:30:49.7 UT
 U3 = 12:35:30.8 UT
 U4 = 12:36:23.2 UT

Local Circumstances at Greatest Eclipse

Lat. = 45°04.5'N Sun Alt. = 59.4°
 Long. = 024°18.0'E Sun Azm. = 196.2°
 Path Width = 112.3 km Duration = 02m22.4s

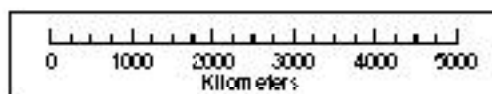
Ephemeris & Constants

Eph. = DE200/LE200
 $\Delta T = 64.6$ s
 $k_1 = 0.2725076$
 $k_2 = 0.2722810$
 $\Delta b = 0.0''$ $\Delta l = 0.0''$

Geocentric Libration
 (Optical + Physical)

$l = 4.73^\circ$
 $b = -0.68^\circ$
 $c = 19.50^\circ$

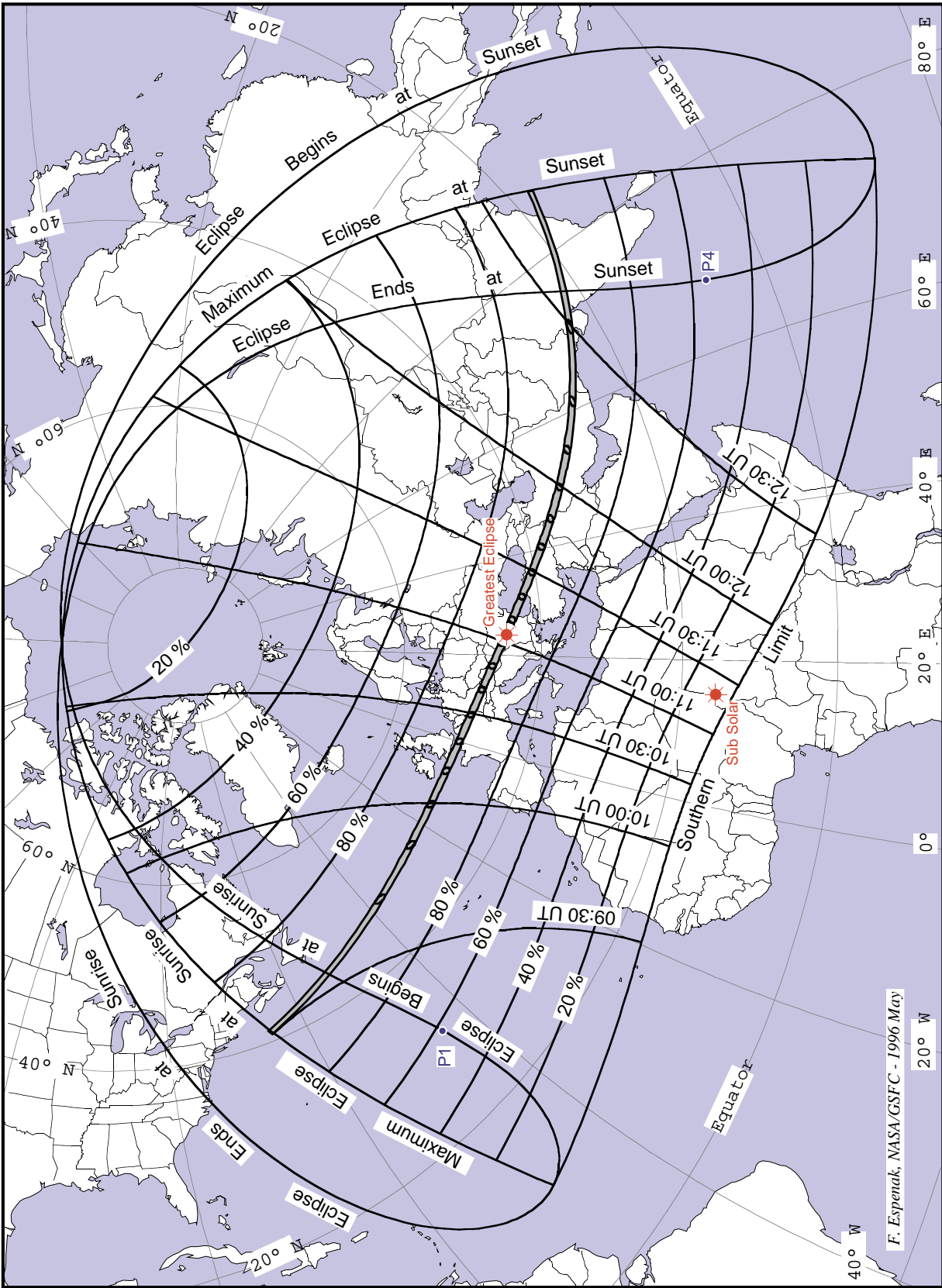
Brown Lun. No. = 1232



F. Espenak, NASA/GSFC - 1996 Apr 16

Total Solar Eclipse of 1999 August 11

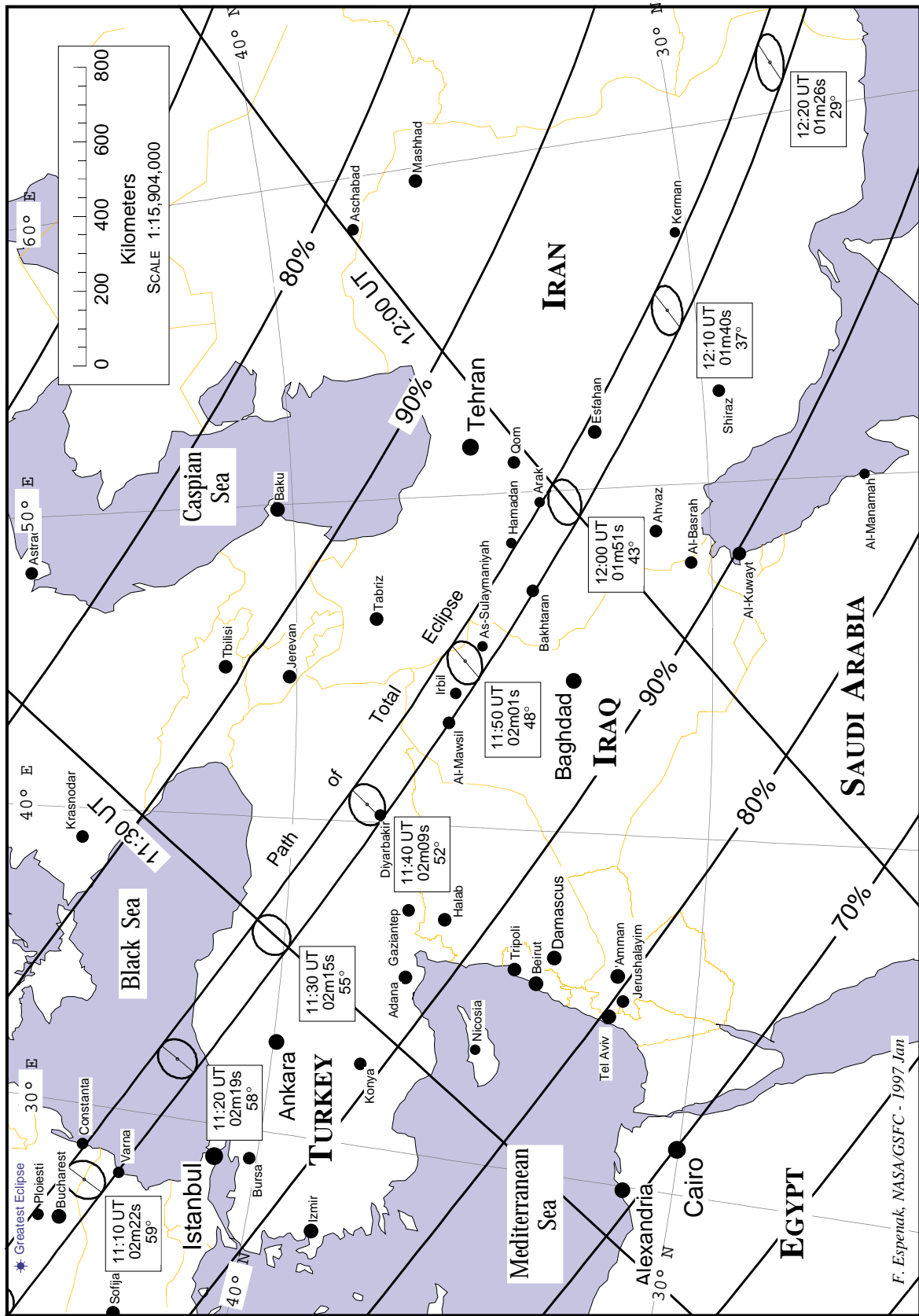
FIGURE 2: STEREOGRAPHIC PROJECTION MAP OF THE ECLIPSE PATH



F. Espenak, NASA/GSFC - 1996 May

Total Solar Eclipse of 1999 August 11

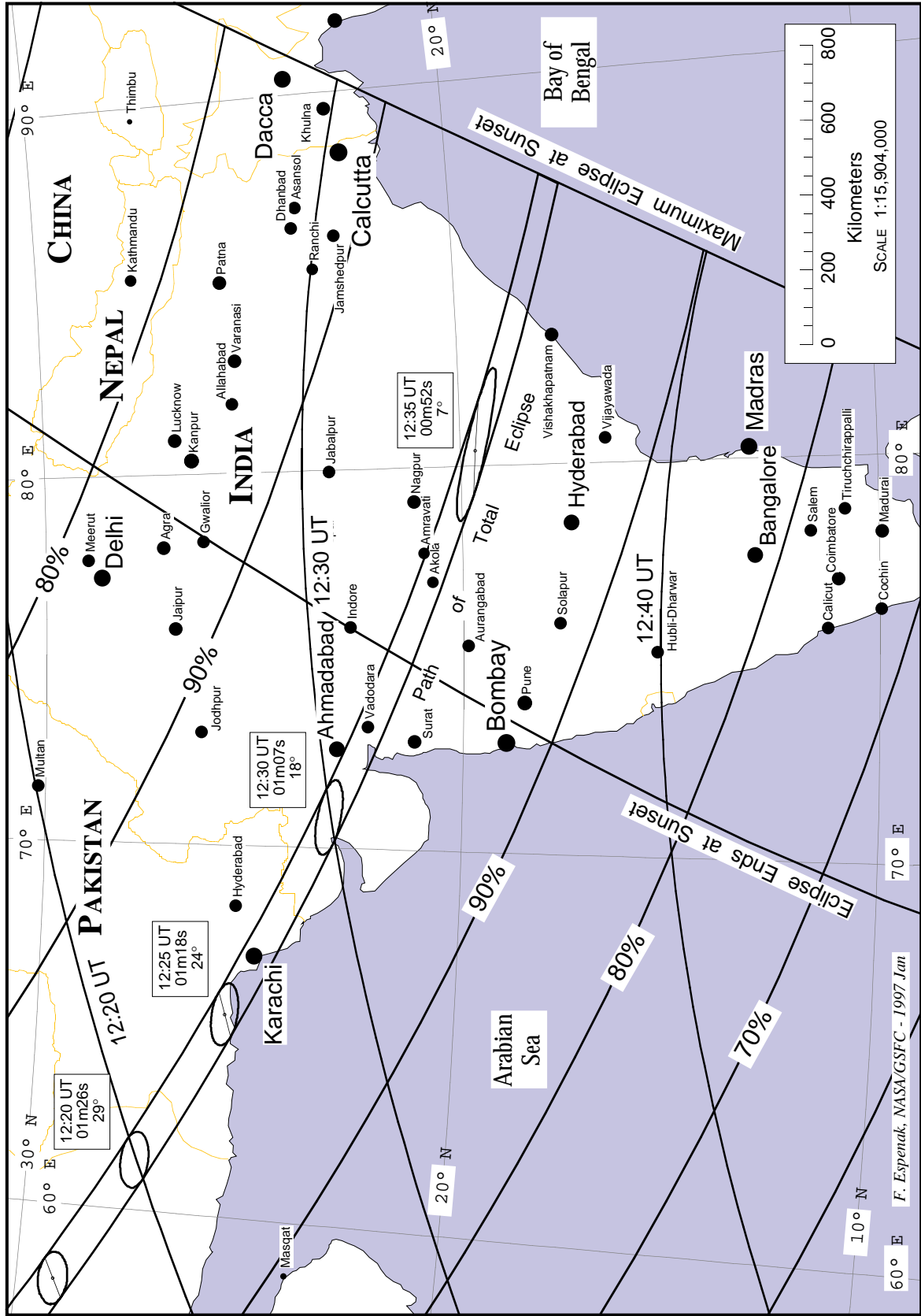
FIGURE 4: THE ECLIPSE PATH THROUGH THE MIDDLE EAST



F. Espenak, NASA/GSFC - 1997 Jan

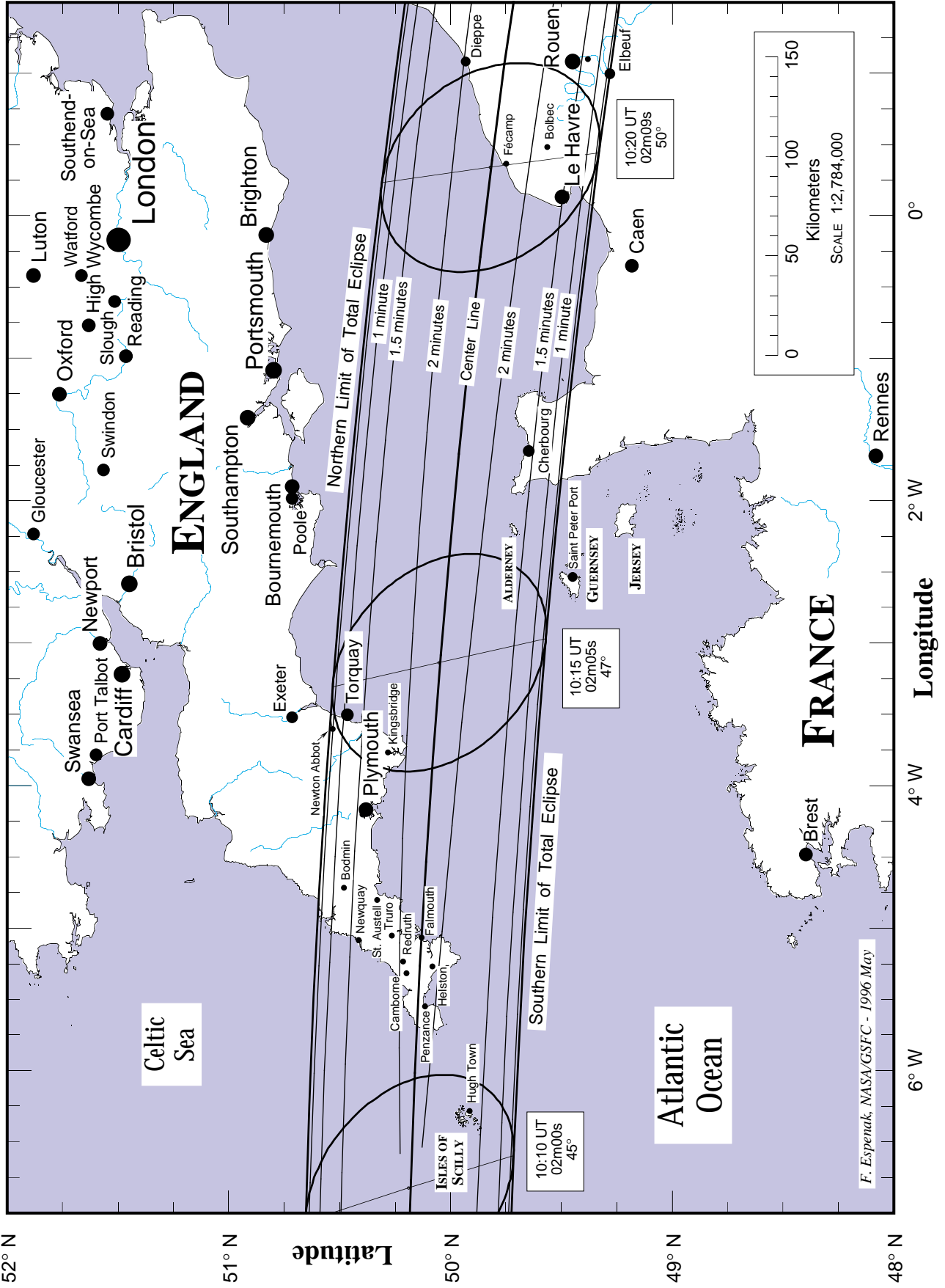
Total Solar Eclipse of 1999 August 11

FIGURE 5: THE ECLIPSE PATH THROUGH SOUTH ASIA



Total Solar Eclipse of 1999 August 11

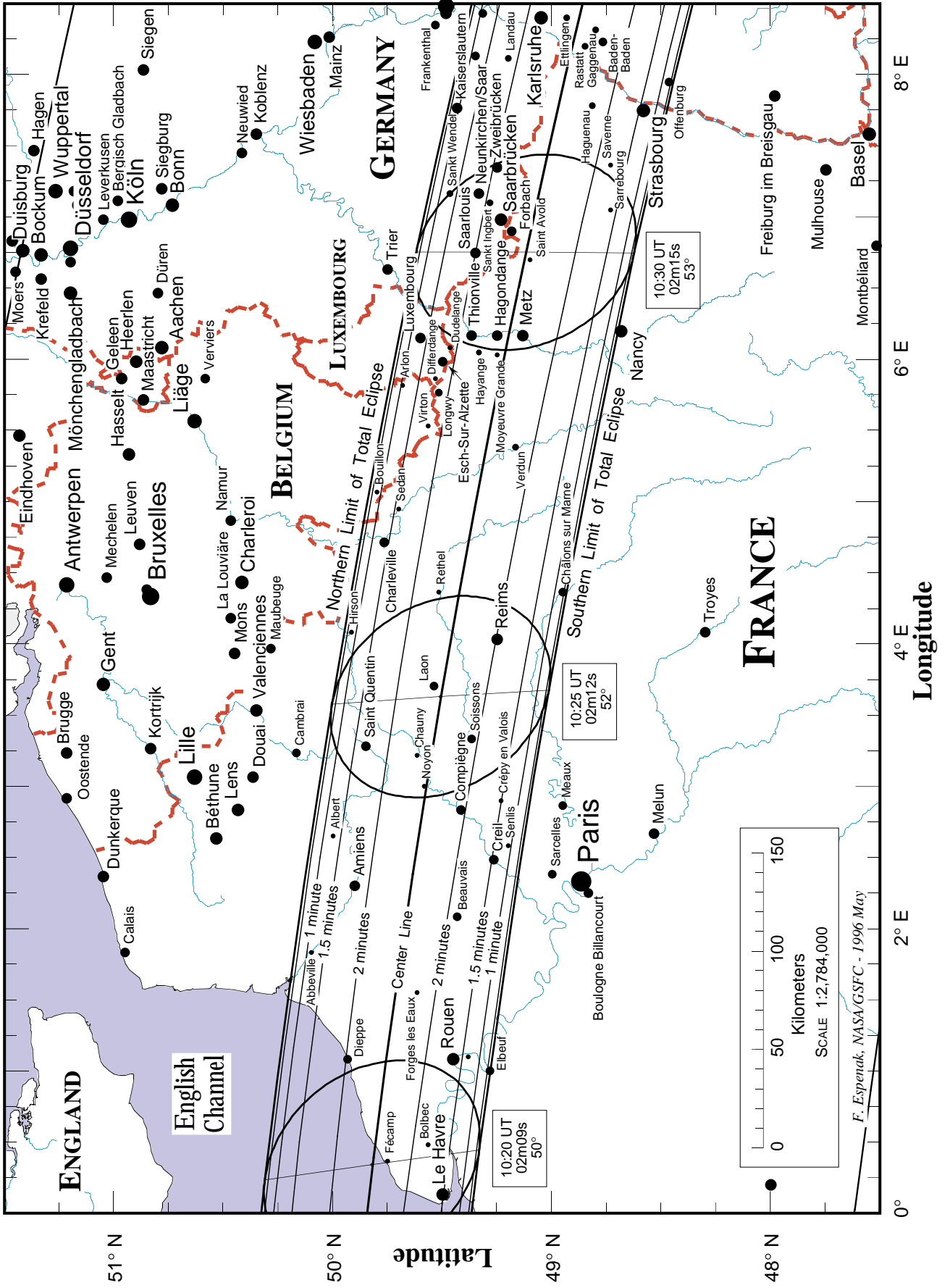
FIGURE 6: THE ECLIPSE PATH THROUGH ENGLAND AND FRANCE



F. Espenak, NASA/GSFC - 1996 May

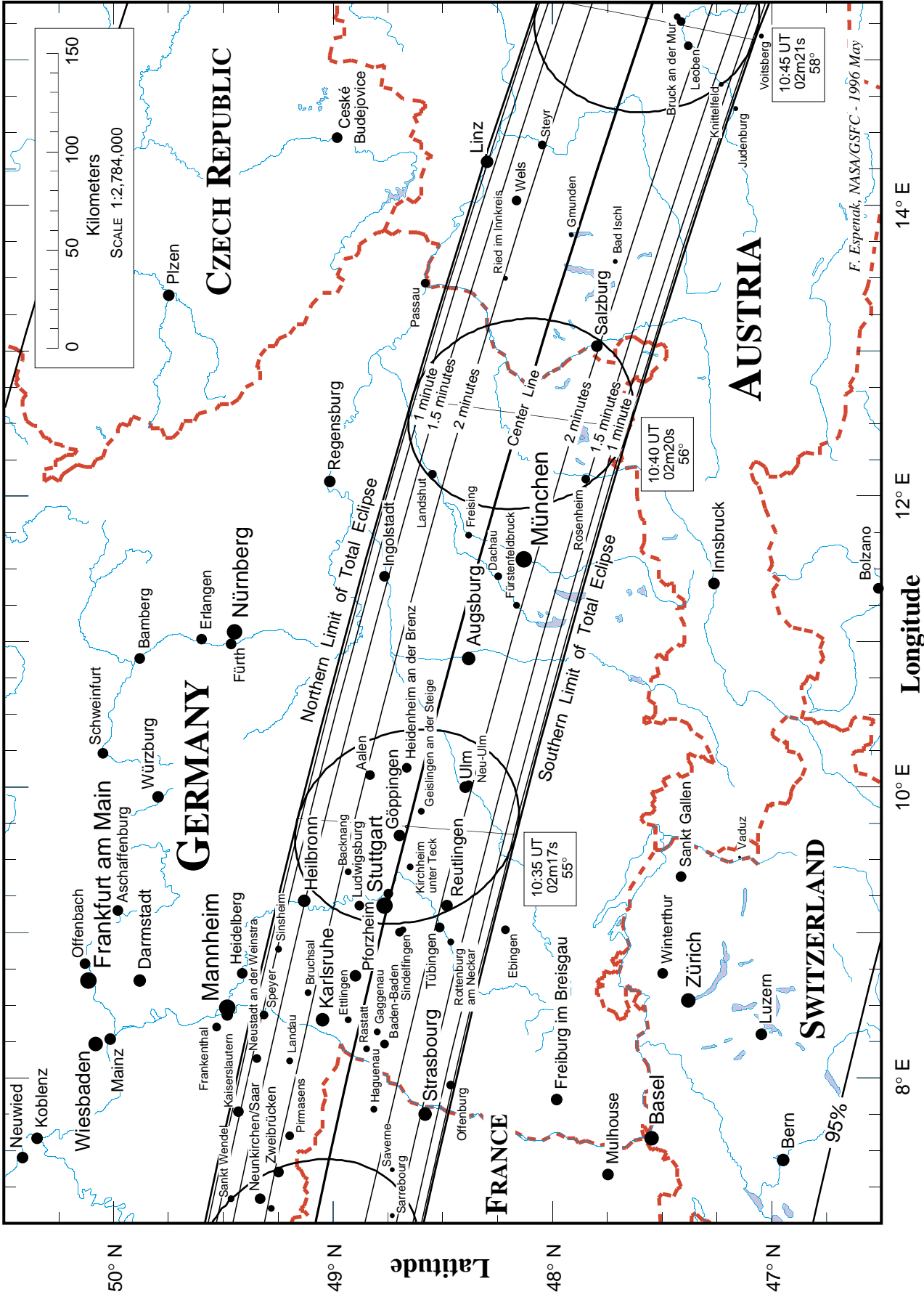
Total Solar Eclipse of 1999 August 11

FIGURE 7: THE ECLIPSE PATH THROUGH FRANCE, BELGIUM, LUXEMBOURG AND GERMANY



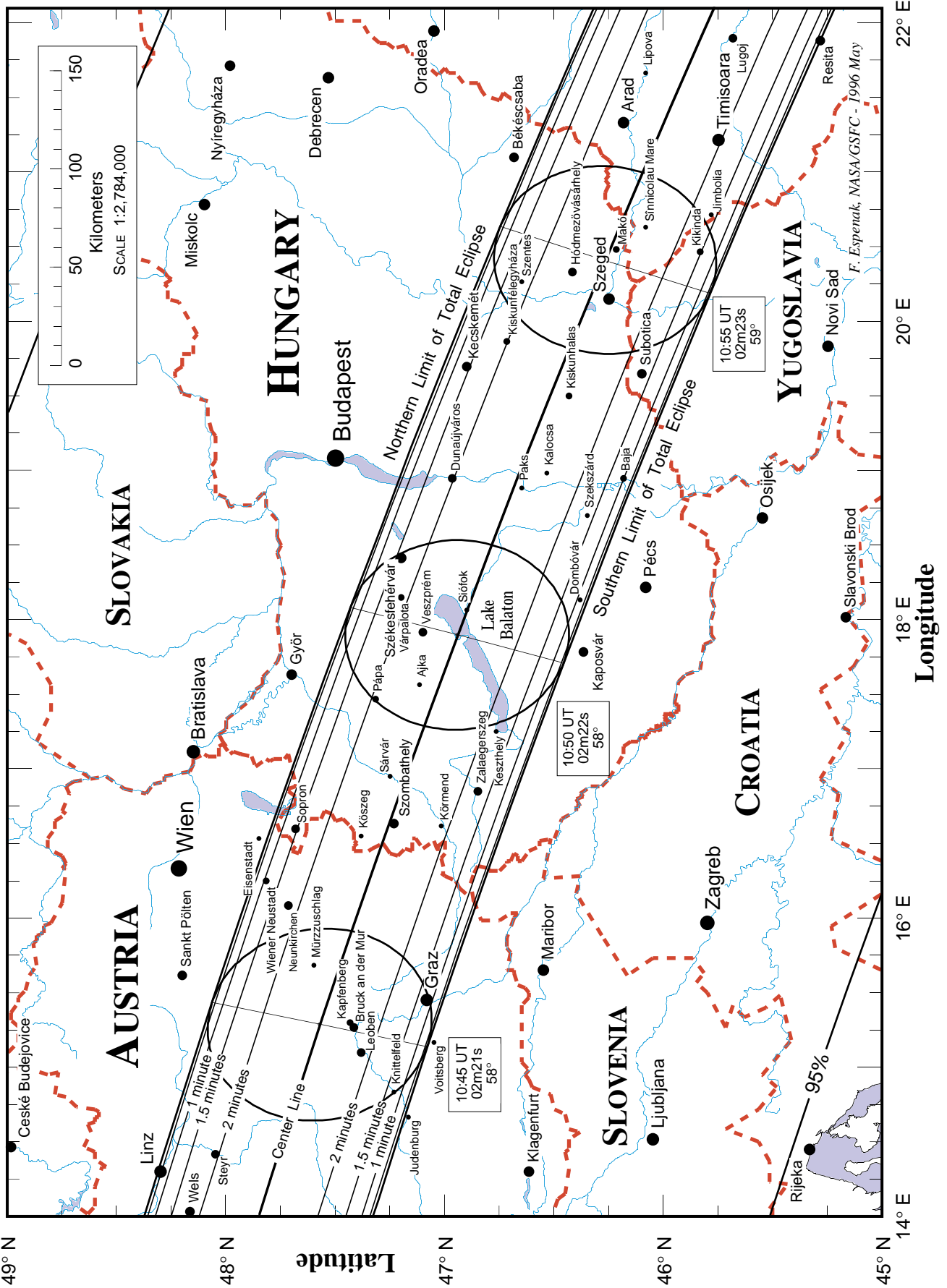
Total Solar Eclipse of 1999 August 11

FIGURE 8: THE ECLIPSE PATH THROUGH GERMANY AND AUSTRIA



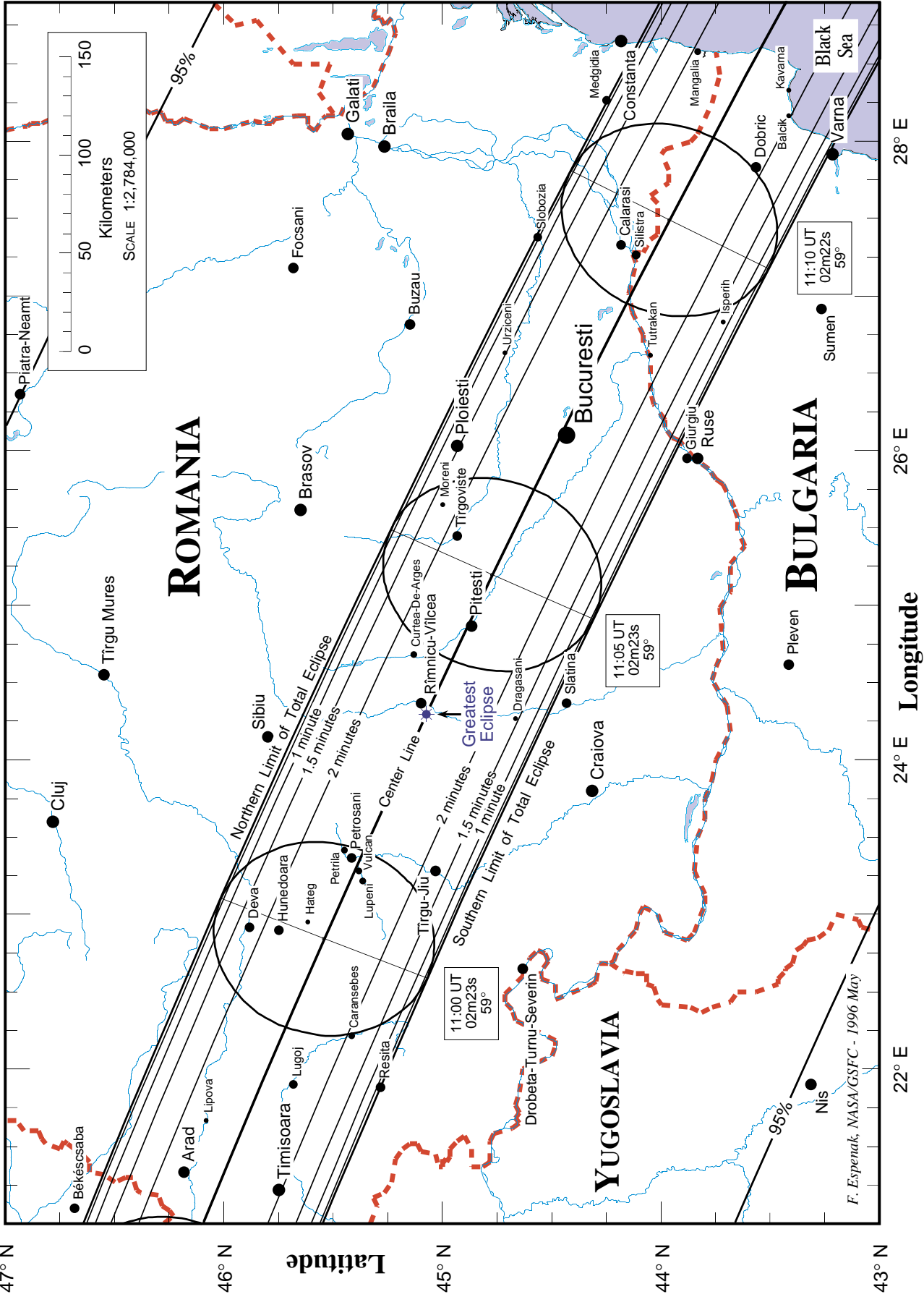
Total Solar Eclipse of 1999 August 11

FIGURE 9: THE ECLIPSE PATH THROUGH AUSTRIA, HUNGARY AND ROMANIA



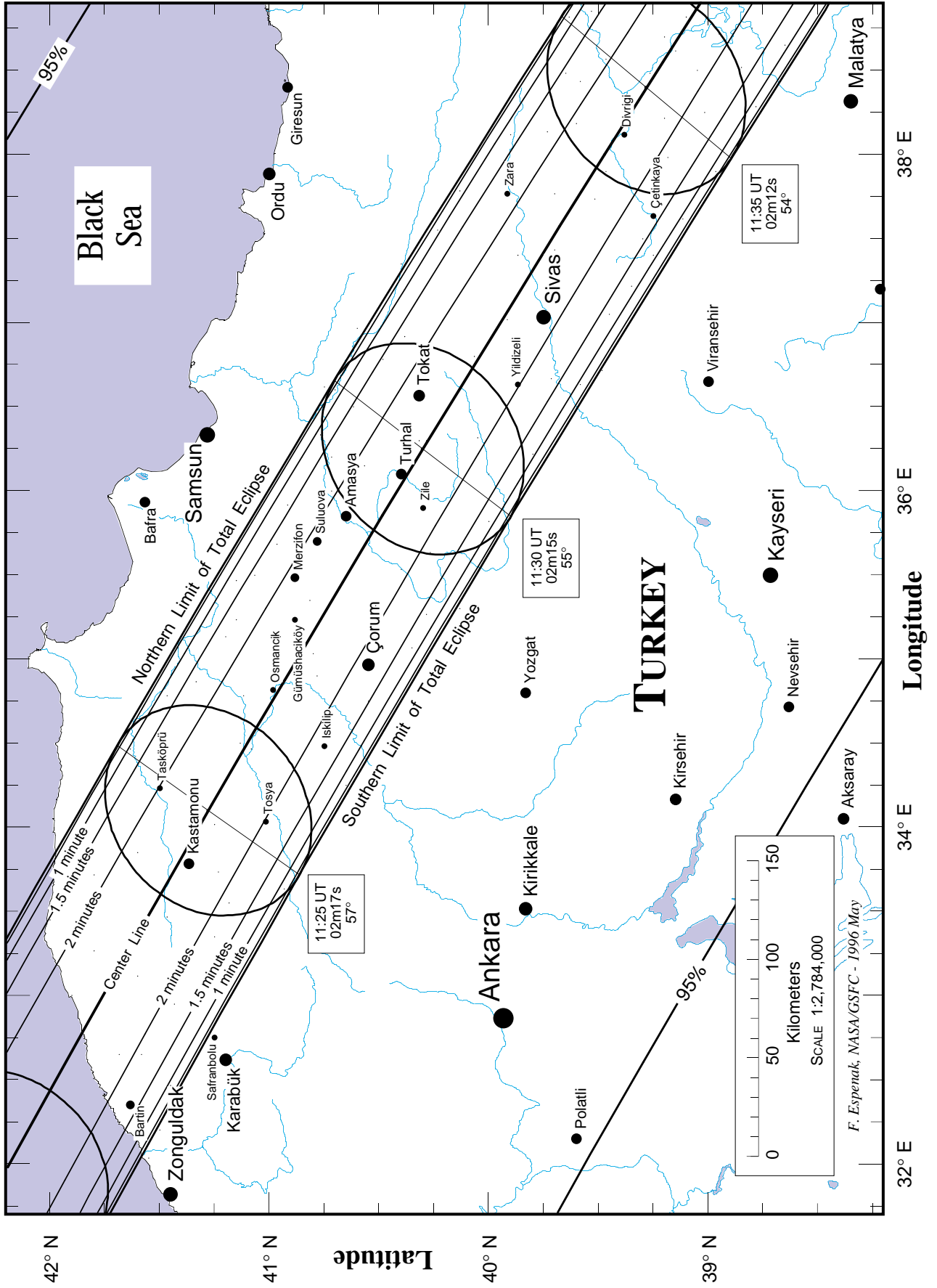
Total Solar Eclipse of 1999 August 11

FIGURE 10: THE ECLIPSE PATH THROUGH ROMANIA AND BULGARIA

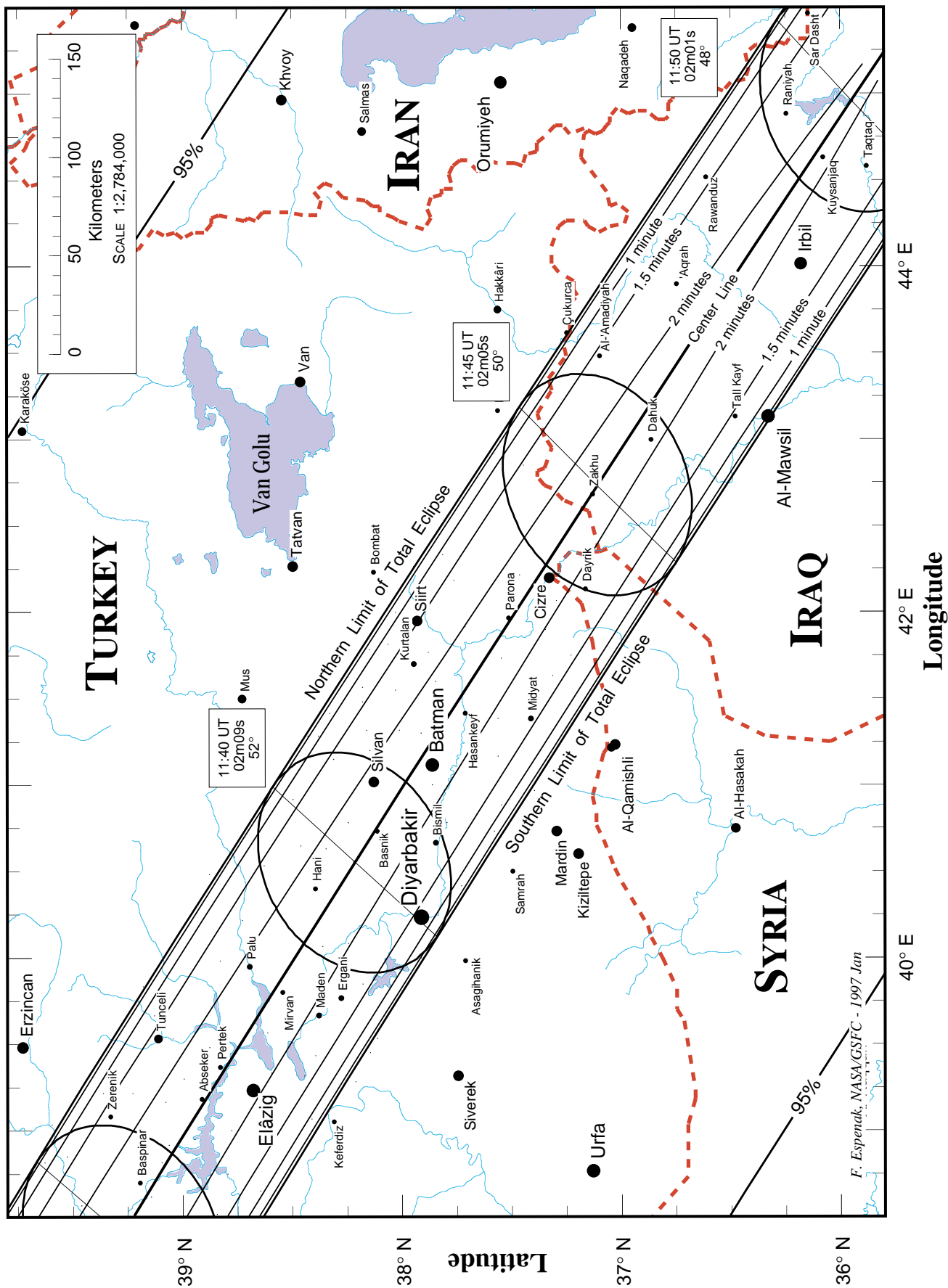


Total Solar Eclipse of 1999 August 11

FIGURE 11: THE ECLIPSE PATH THROUGH TURKEY

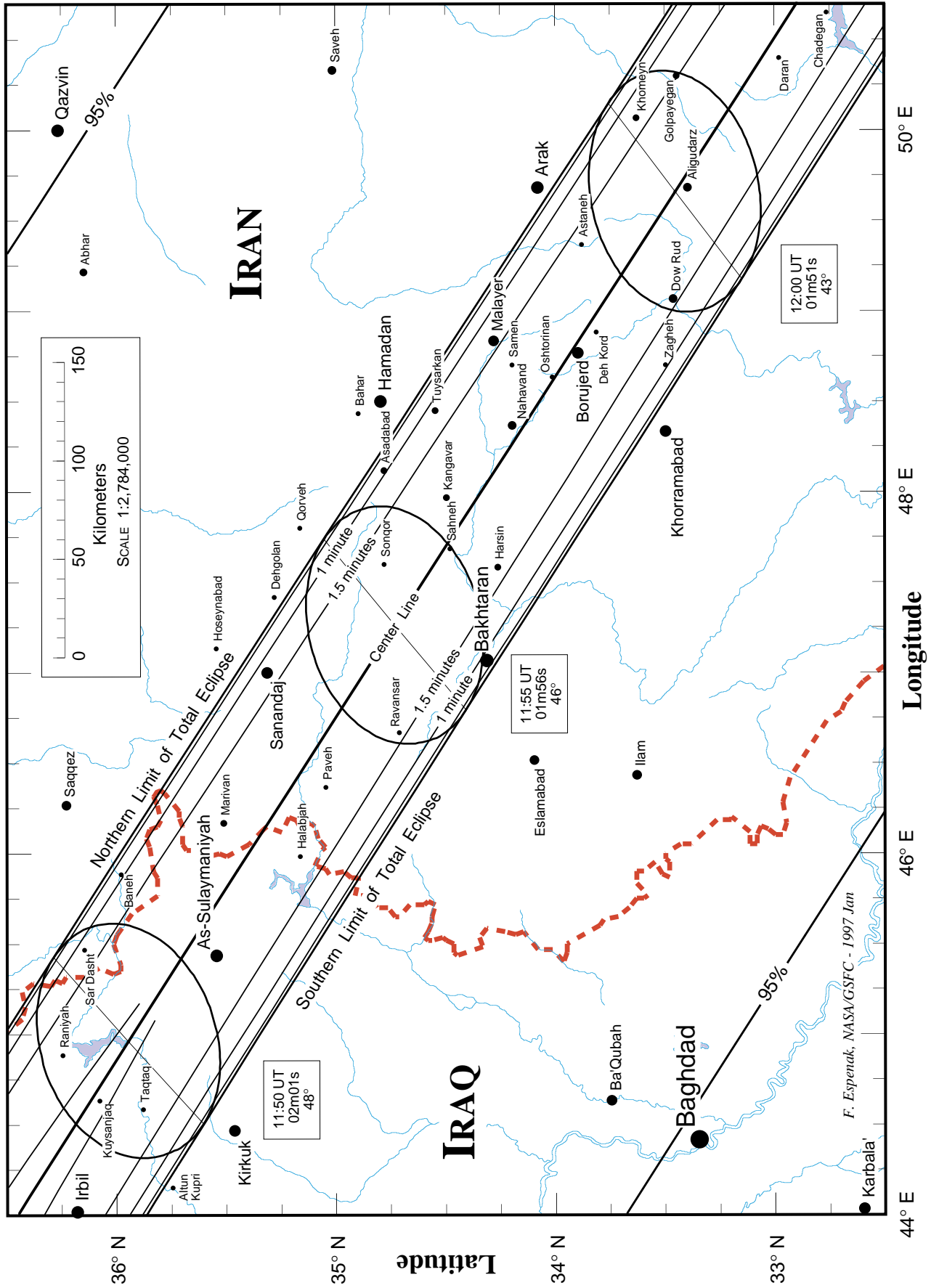


Total Solar Eclipse of 1999 August 11
FIGURE 12: THE ECLIPSE PATH THROUGH TURKEY, SYRIA AND IRAQ



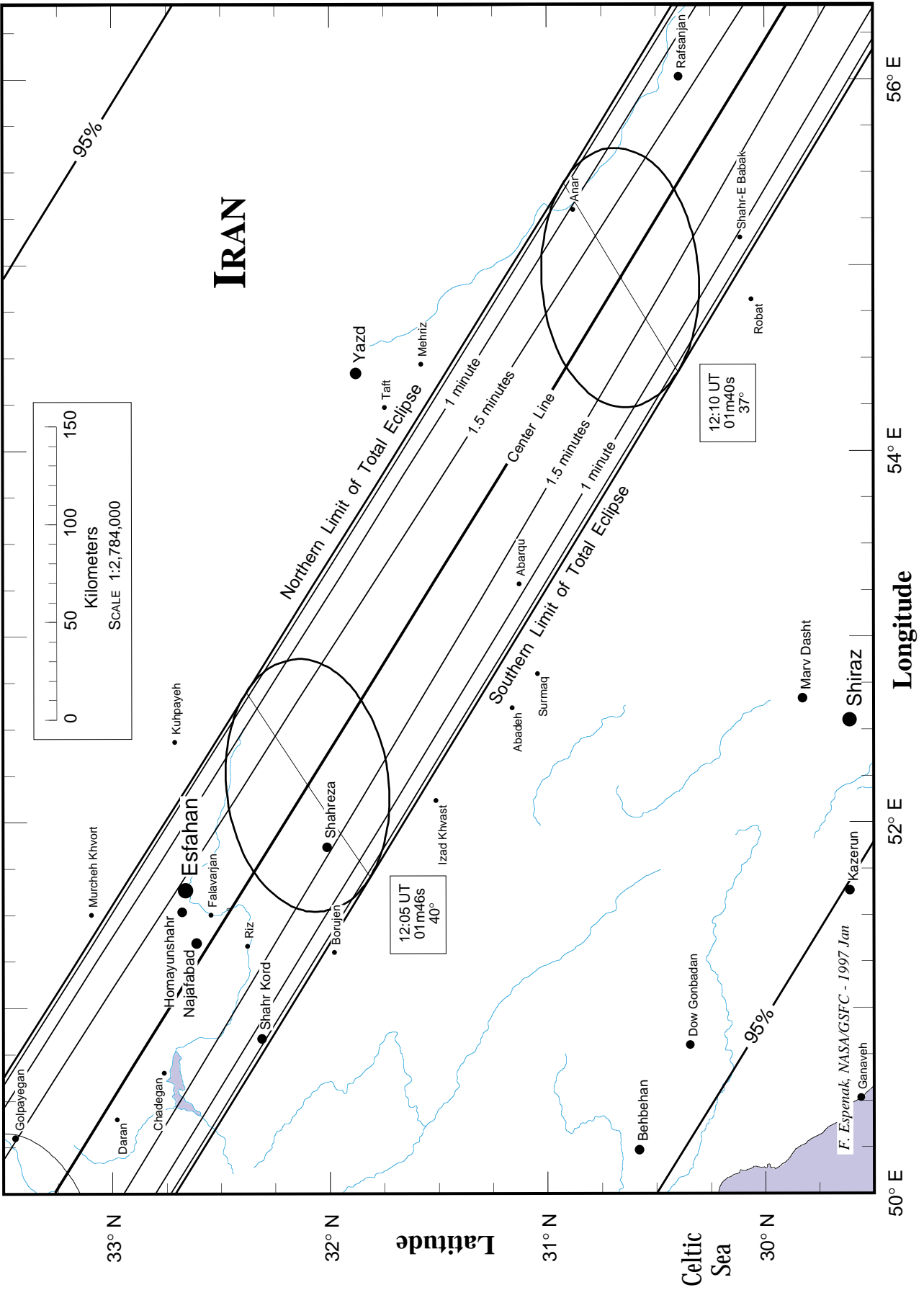
Total Solar Eclipse of 1999 August 11

FIGURE 13: THE ECLIPSE PATH THROUGH IRAQ AND IRAN



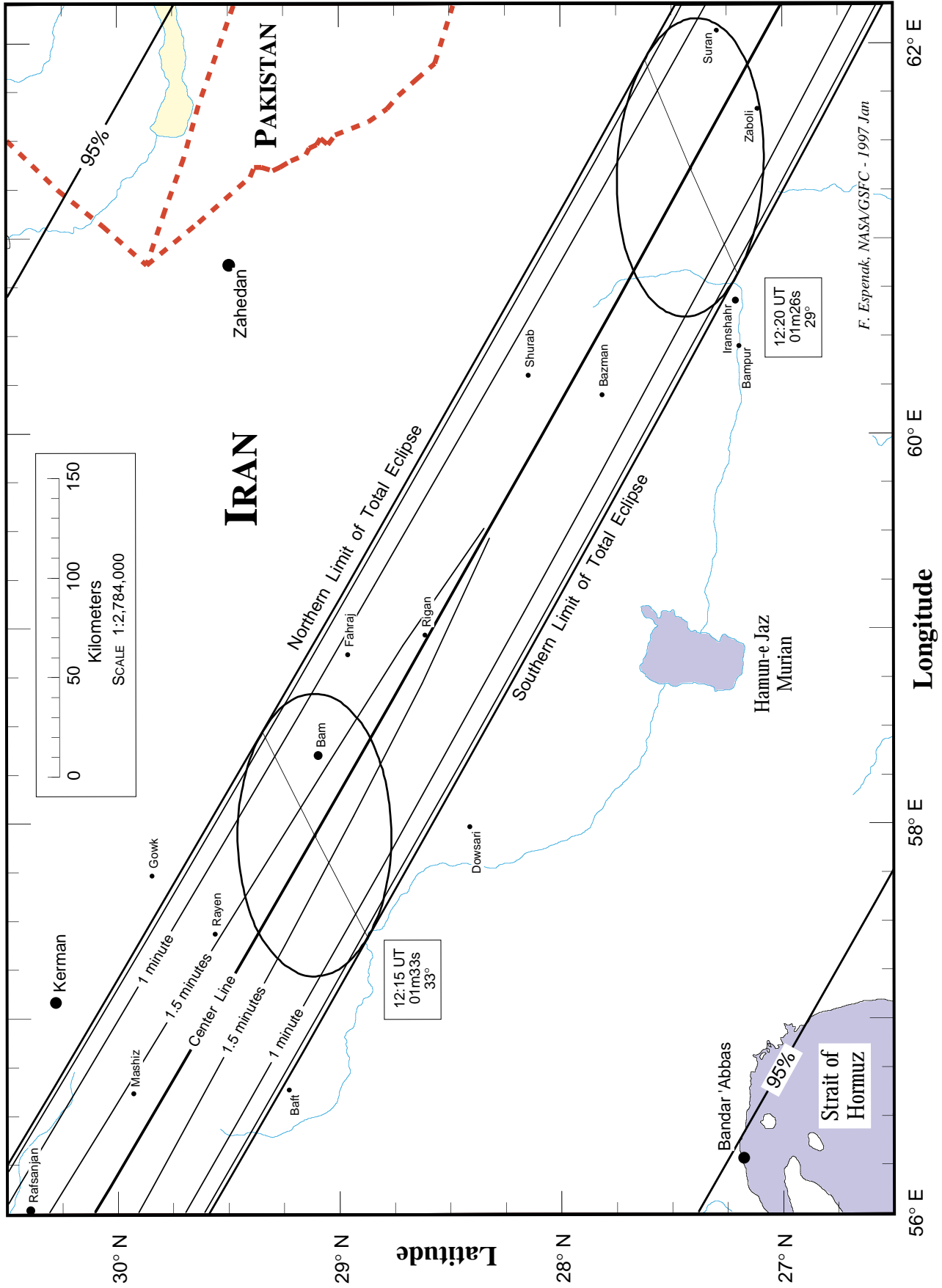
Total Solar Eclipse of 1999 August 11

FIGURE 14: THE ECLIPSE PATH THROUGH IRAN



Total Solar Eclipse of 1999 August 11

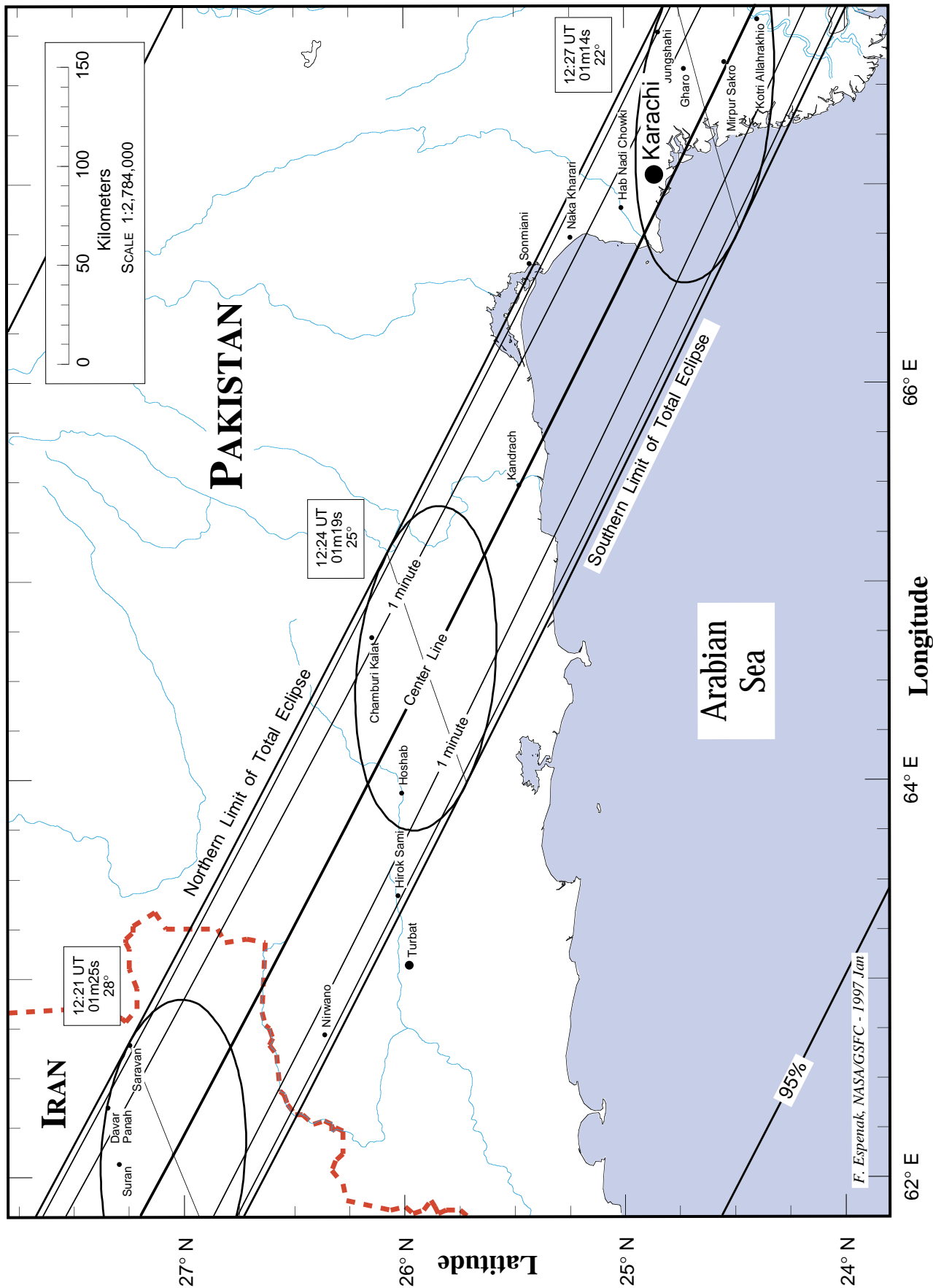
FIGURE 15: THE ECLIPSE PATH THROUGH SOUTHERN IRAN



F. Espenak, NASA/GSFC - 1997 Jan

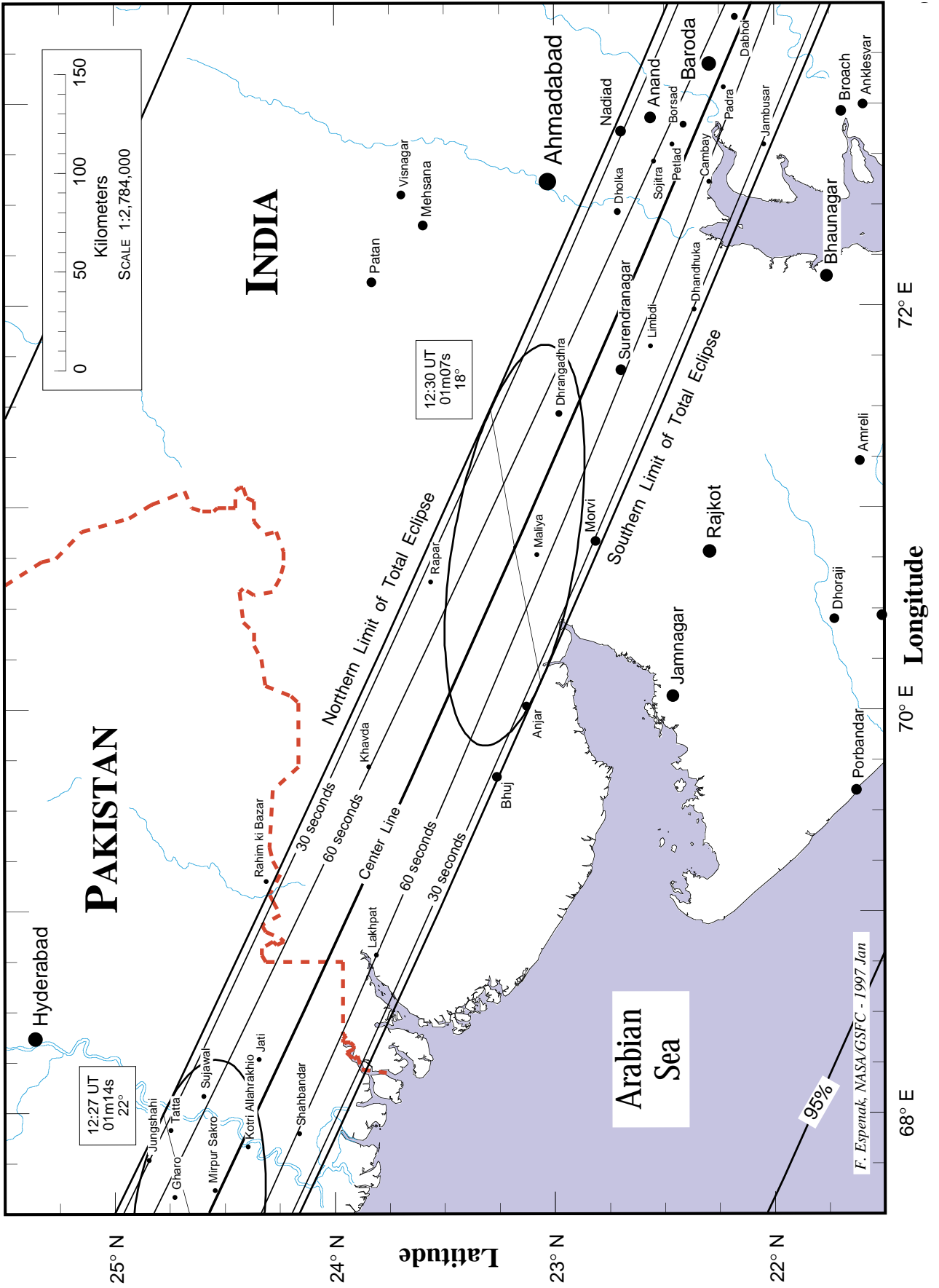
Total Solar Eclipse of 1999 August 11

FIGURE 16: THE ECLIPSE PATH THROUGH PAKISTAN



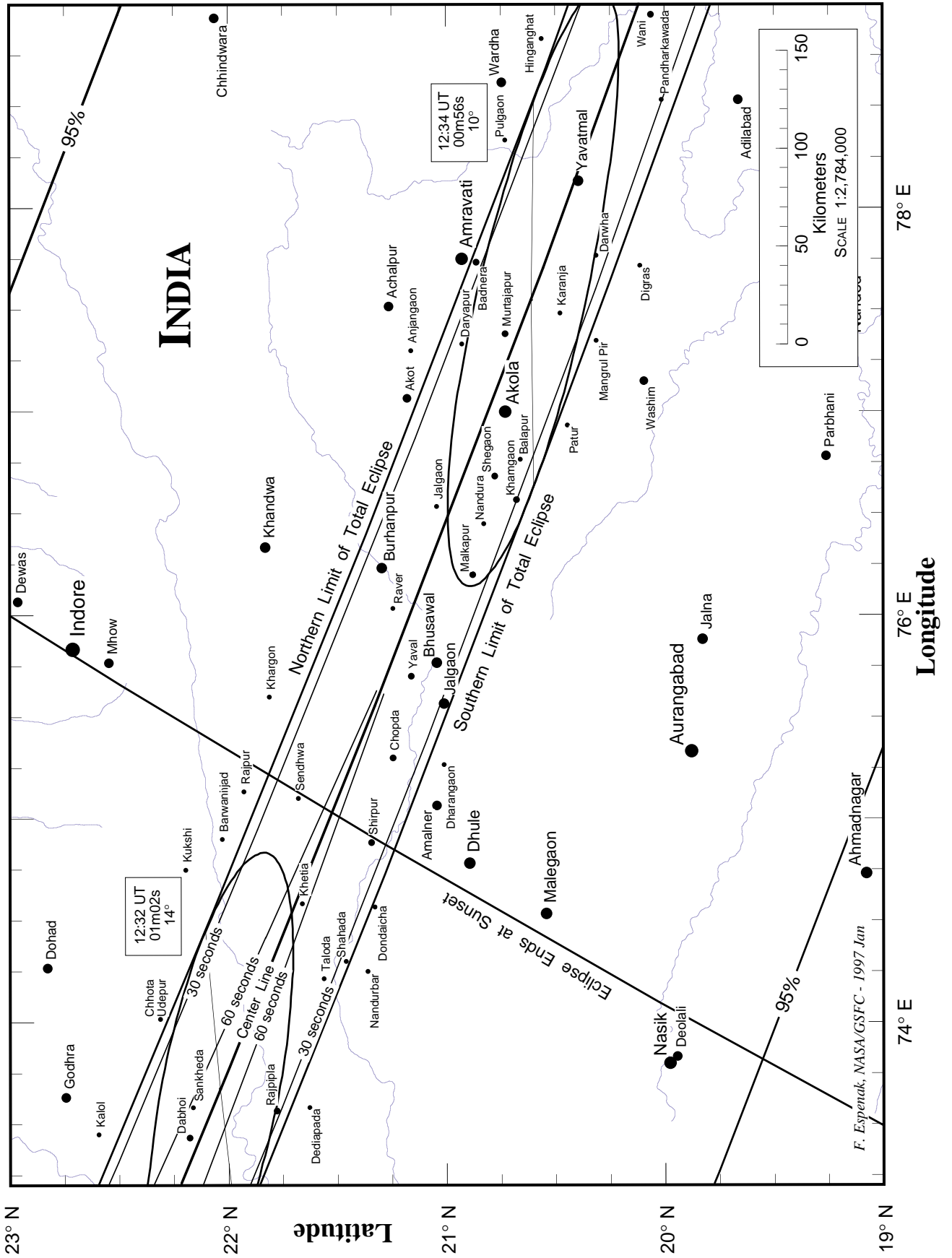
Total Solar Eclipse of 1999 August 11

FIGURE 17: THE ECLIPSE PATH THROUGH PAKISTAN AND INDIA



Total Solar Eclipse of 1999 August 11

FIGURE 18: THE ECLIPSE PATH THROUGH CENTRAL INDIA



Total Solar Eclipse of 1999 August 11

FIGURE 19: THE ECLIPSE PATH THROUGH EASTERN INDIA

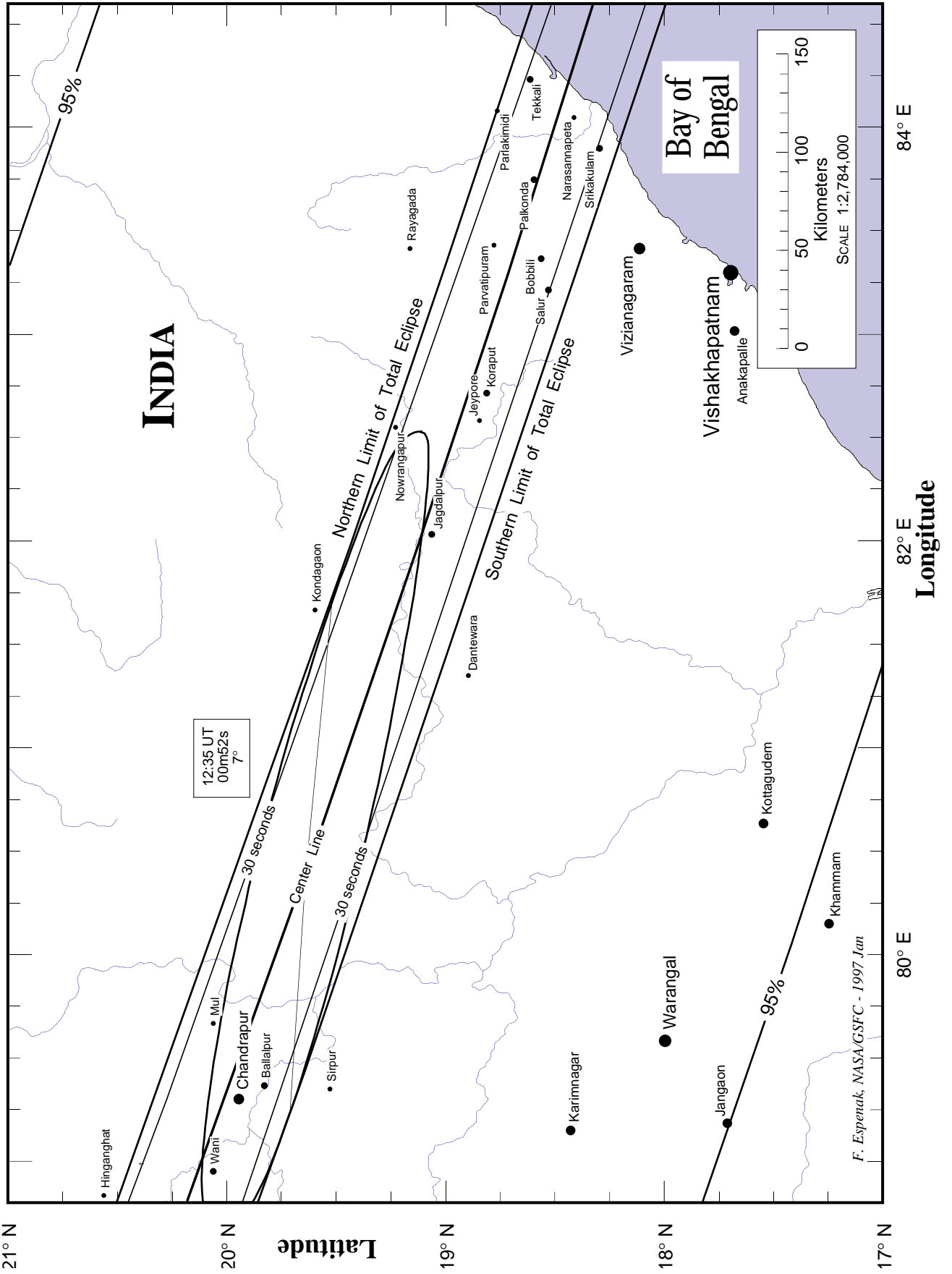
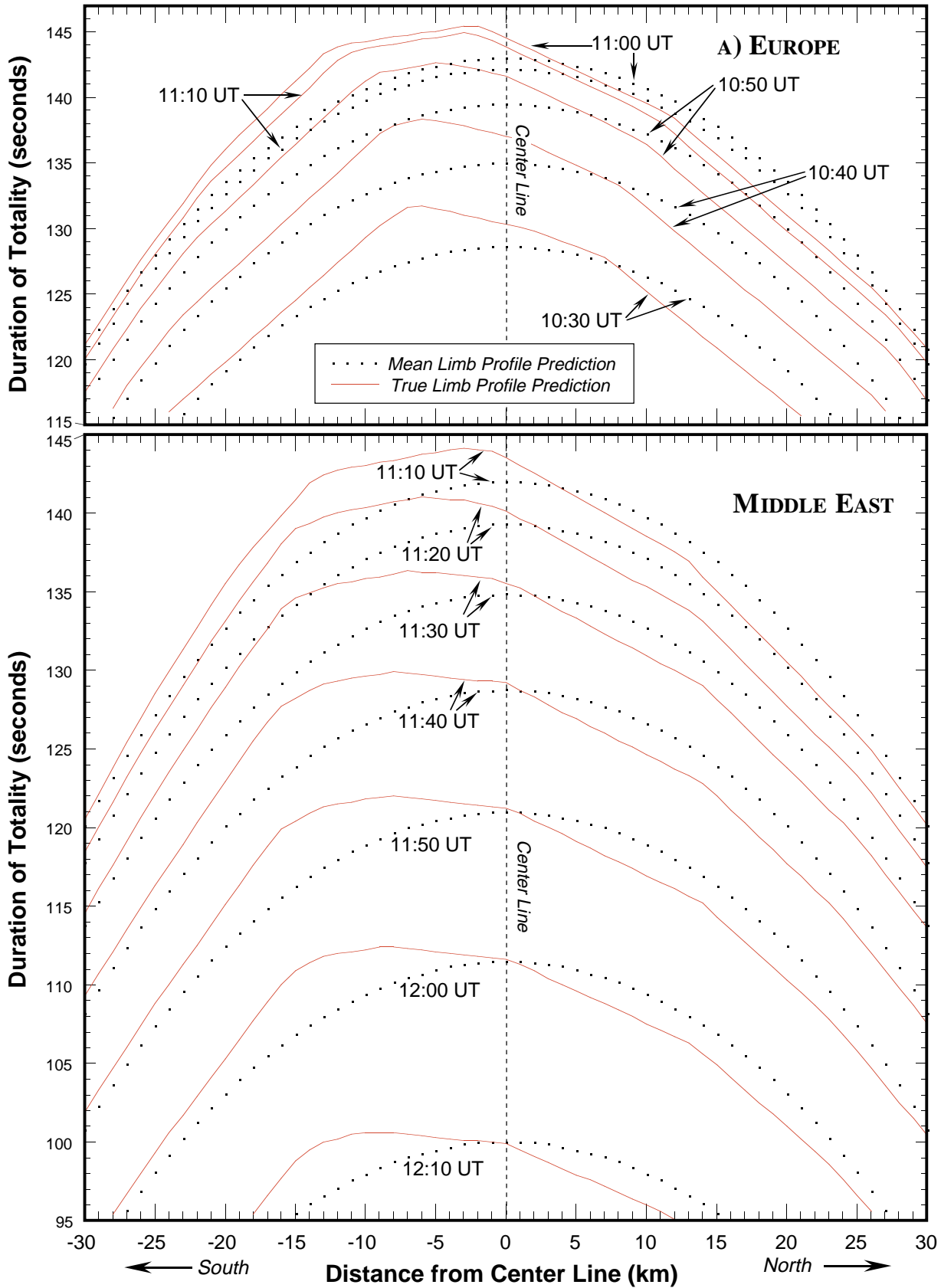


Figure 20 would not convert
to the PDF format.

Total Solar Eclipse of 1999 August 11

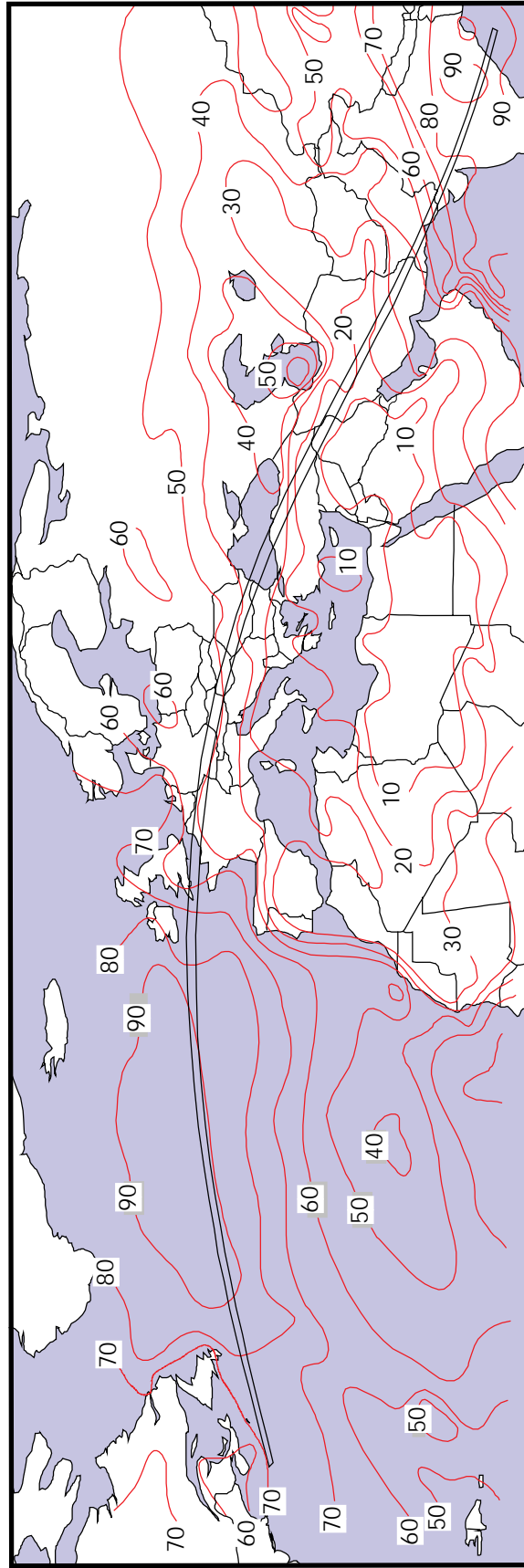
FIGURE 21: LIMB PROFILE EFFECTS ON THE DURATION OF TOTALITY



F. Esdenak - 1997 Jan 31

Total Solar Eclipse of 1999 August 11

FIGURE 22: MEAN CLOUD COVER IN AUGUST ALONG THE ECLIPSE PATH

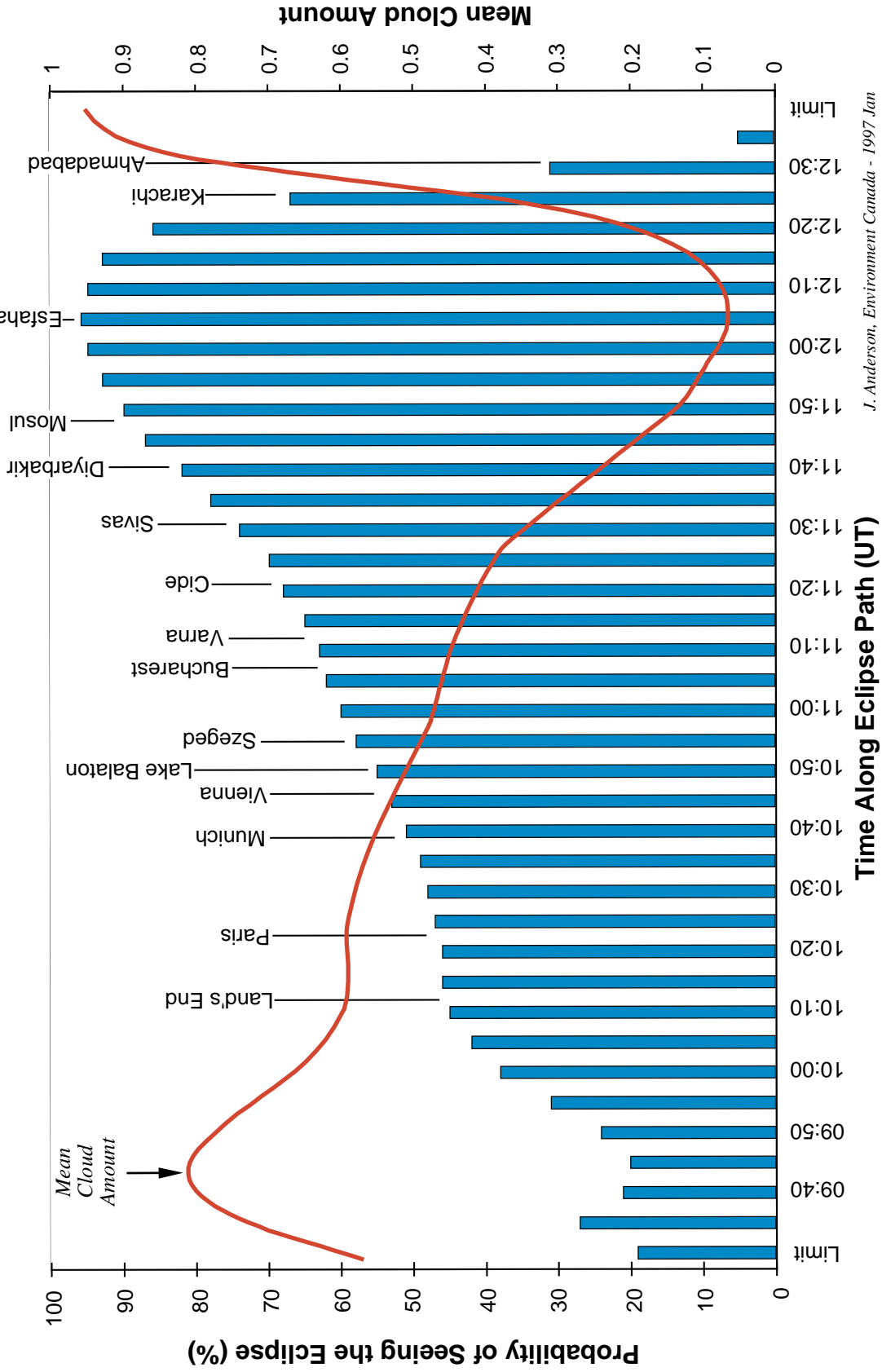


J. Anderson, Environment Canada - 1997 Jan

Figure 22: Mean cloud cover (in percent) along the eclipse track as determined by satellite measurements from eight years of analysis (1983-1990). These data are collected and analyzed globally from a number of satellites (International Satellite Cloud Climatology Project), and processed by computer. It provides an excellent comparative database for different locations around the globe. Statistics are collected over a 5° by 5° latitude/longitude area, and represent the large scale cloud characteristics of an area. Small scale variations are smoothed out.

Total Solar Eclipse of 1999 August 11

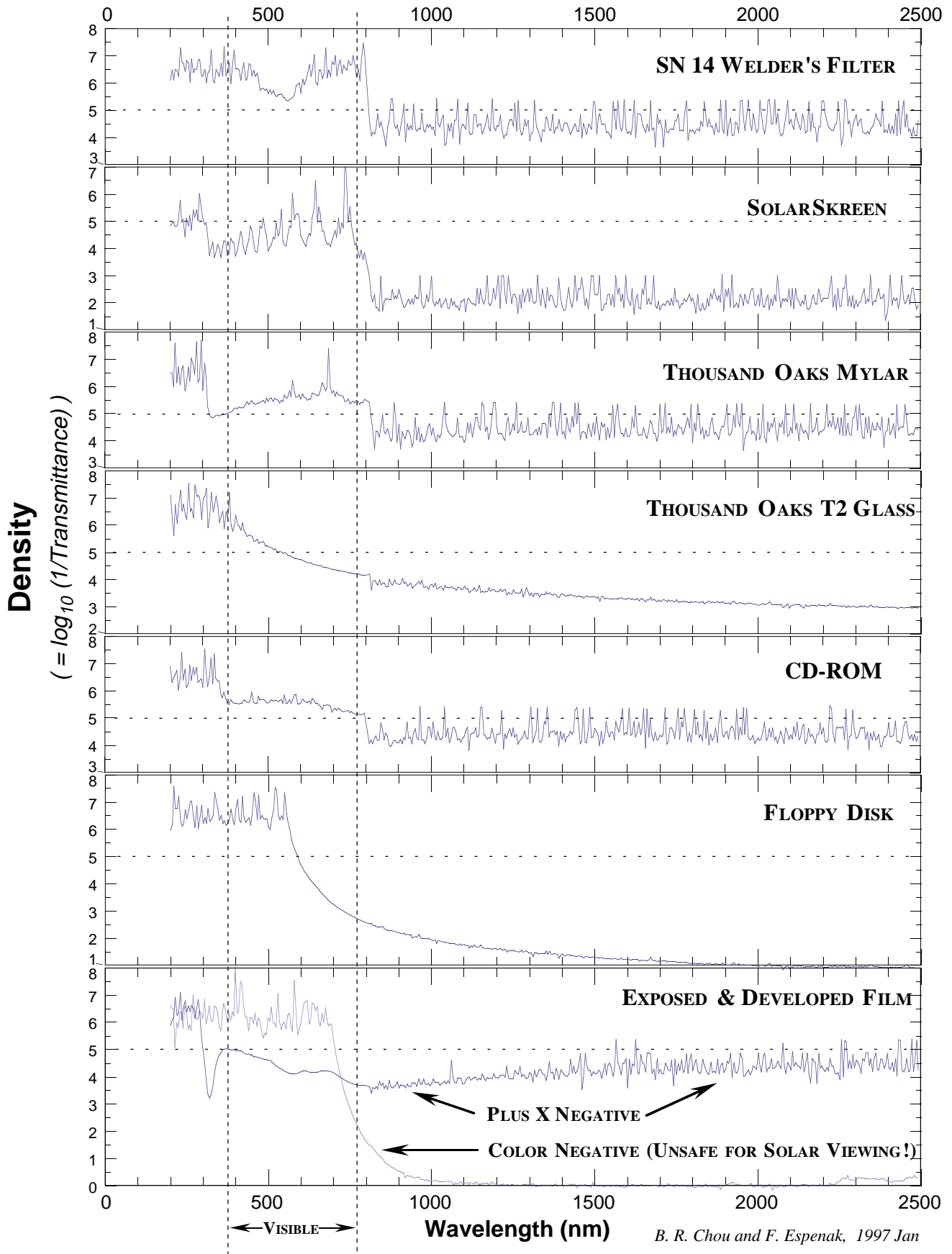
FIGURE 23: PROBABILITY OF SEEING THE ECLIPSE ALONG THE PATH



J. Anderson, Environment Canada - 1997 Jan

Figure 23: Probability of seeing the eclipse, adjusted for Sun altitude, time of day, and the actual date of the eclipse. Values are smoothed by the computer modeling so that fine details of the cloud cover are lost. Also included is a graph of the mean cloud cover along the track. The data are presented at 5 minute intervals along the eclipse track, with the locations of important cities and landmarks.

FIGURE 24: SPECTRAL RESPONSE OF SOME COMMONLY AVAILABLE SOLAR FILTERS



B. R. Chou and F. Espenak, 1997 Jan

Total Solar Eclipse of 1999 August 11

FIGURE 25: THE SKY DURING TOTALITY AS SEEN FROM CENTER LINE AT 11:00 UT

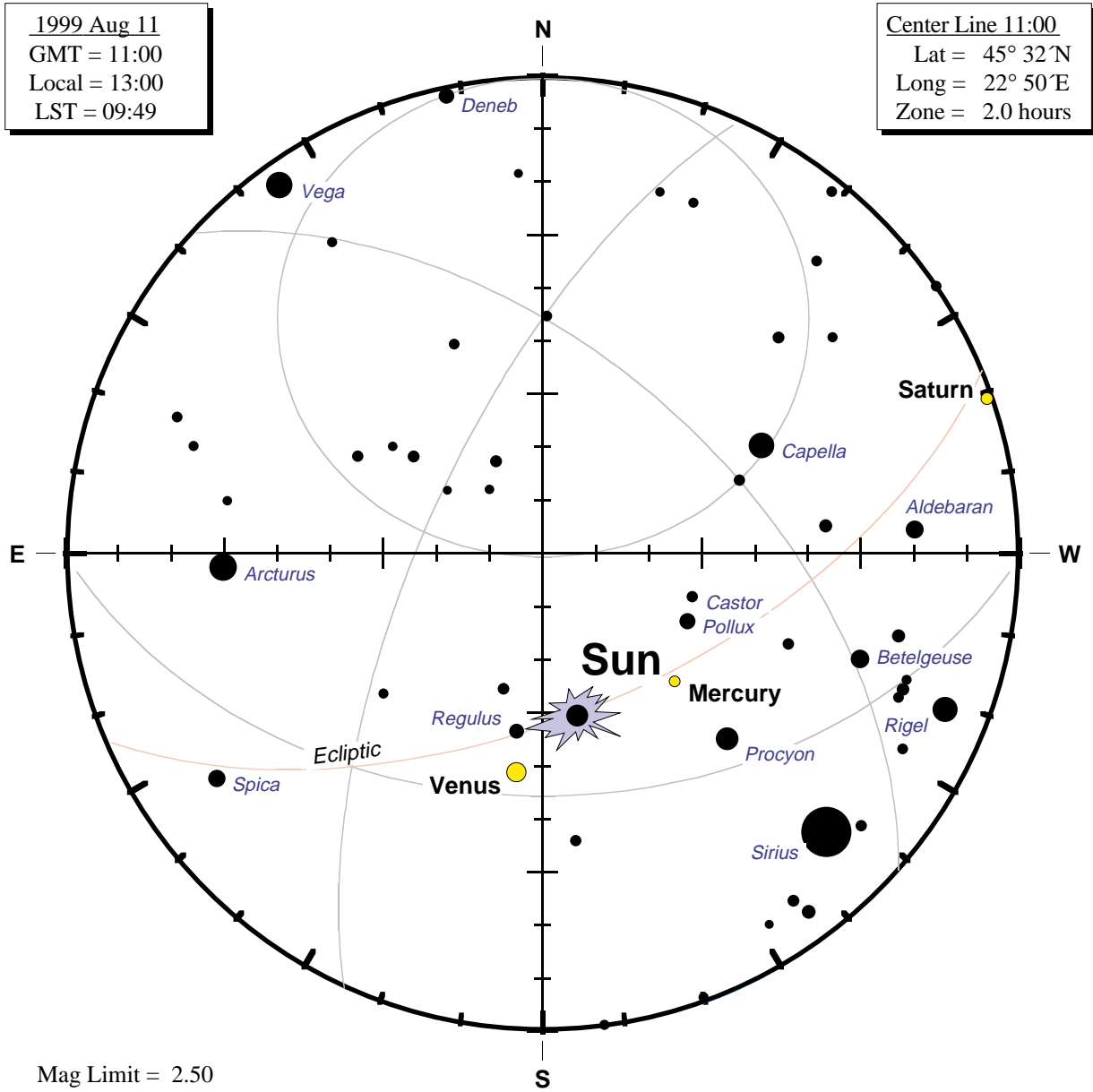
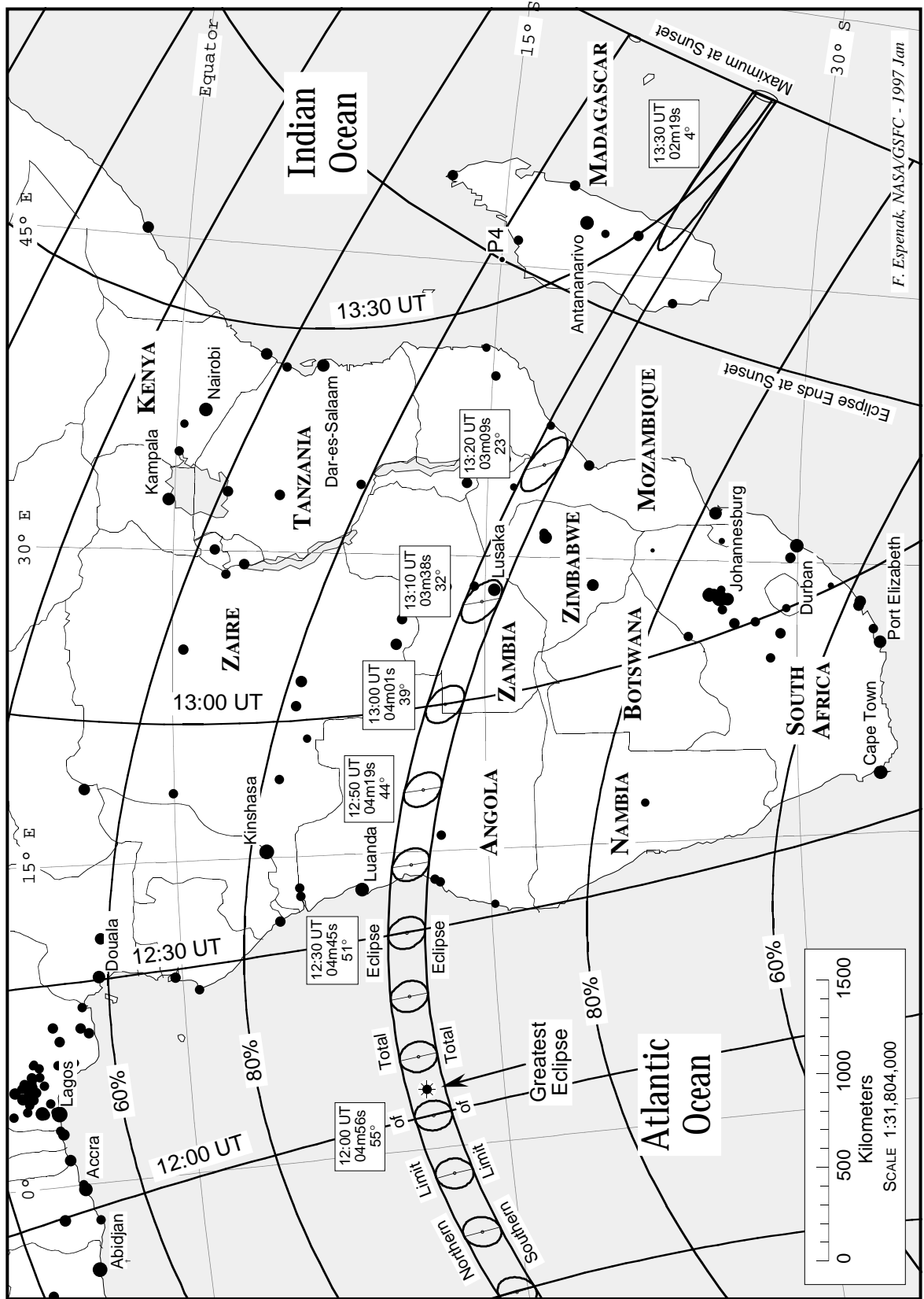


Figure 25: The sky during totality as seen from the center line in Romania at 11:00 UT. Venus ($m=-3.5$) will be the most conspicuous planet located 15° east of the Sun. Mercury ($m=+0.7$) should also be visible 18° west of the Sun. The southwestern sky will be dominated by the bright stars of winter, including Capella ($m=+0.08$), Aldebaran ($m=+0.85v$), Procyon ($+0.38$), Betelgeuse ($+0.5v$), and Sirius ($m=-1.46$). Other bright stars which may also be visible include Spica ($m=+1.0v$), Arcturus ($m=-0.04$), and Regulus ($m=+1.35$).

For sky maps from other locations along the path of totality, see the special 1999 eclipse web site:
<http://planets.gsfc.nasa.gov/eclipse/TSE1999/TSE1999.html>

Total Solar Eclipse of 2001 June 21

FIGURE 26: THE ECLIPSE PATH THROUGH AFRICA



TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

TABLES

TABLE 1

ELEMENTS OF THE TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Geocentric Conjunction 10:52:16.66 TDT J.D. = 2451401.952971
of Sun & Moon in R.A.: (=10:51:12.06 UT)

Instant of 11:04:09.01 TDT J.D. = 2451401.961215
Greatest Eclipse: (=11:03:04.41 UT)

Geocentric Coordinates of Sun & Moon at Greatest Eclipse (DE200/LE200):

<u>Sun:</u>	R.A. = 09h23m08.297s	<u>Moon:</u>	R.A. = 09h23m34.531s
	Dec. = +15°19'39.72"		Dec. = +15°48'38.51"
Semi-Diameter =	15'46.77"	Semi-Diameter =	16'00.34"
Eq. Hor. Par. =	8.68"	Eq. Hor. Par. =	0°58'44.24"
Δ R.A. =	9.467s/h	Δ R.A. =	142.037s/h
Δ Dec. =	-44.35"/h	Δ Dec. =	-462.21"/h

<u>Lunar Radius</u>	k1 = 0.2725076 (Penumbra)	<u>Shift in</u>	Δb = 0.00"
<u>Constants:</u>	k2 = 0.2722810 (Umbra)	<u>Lunar Position:</u>	Δl = 0.00"

<u>Geocentric Libration:</u>	l = 4.8°	Brown Lun. No. =	1232
(Optical + Physical)	b = -0.8°	Saros Series =	145 (21/77)
	c = 19.7°	Ephemeris =	(DE200/LE200)

Eclipse Magnitude = 1.02859 Gamma = 0.50623 ΔT = 64.6 s

Polynomial Besselian Elements for: 1999 Aug 11 11:00:00.0 TDT (=t₀)

n	x	y	d	l ₁	l ₂	μ
0	0.0700559	0.5028388	15.3273401	0.5424893	-0.0036496	343.690308
1	0.5443045	-0.1184931	-0.0120348	0.0001168	0.0001163	15.002983
2	-0.0000406	-0.0001158	-0.0000033	-0.0000117	-0.0000116	0.000002
3	-0.0000081	0.0000017	0.0000000	0.0000000	0.0000000	0.000000

Tan f₁ = 0.0046129 Tan f₂ = 0.0045900

At time 't₁' (decimal hours), each Besselian element is evaluated by:

$$a = a_0 + a_1t + a_2t^2 + a_3t^3 \quad (\text{or } a = \sum [a_n t^n]; n = 0 \text{ to } 3)$$

where: a = x, y, d, l₁, l₂, or μ
 t = t₁ - t₀ (decimal hours) and t₀ = 11.000 TDT

The Besselian elements were derived from a least-squares fit to elements calculated at five separate times over a six hour period centered at t₀. Thus the Besselian elements are valid over the period 8.00 ≤ t₀ ≤ 14.00 TDT.

Saros Series 145: Member 21 of 77 eclipses in series.

TABLE 2

SHADOW CONTACTS AND CIRCUMSTANCES
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

$\Delta T = 64.6 \text{ s}$
 $= 000^{\circ}16'09.6''\text{E}$

		Terrestrial Dynamical Time h m s	Latitude	Ephemeris Longitude†	True Longitude*
External/Internal Contacts of Penumbra:	P1	08:27:18.9	30°20.1'N	044°45.6'W	044°29.4'W
	P4	13:41:10.2	06°38.3'N	067°50.0'E	068°06.2'E
Extreme South Limits of Penumbral Path:	S1	09:09:07.2	10°07.7'N	048°46.6'W	048°30.4'W
	S2	12:59:11.4	13°43.1'S	072°40.6'E	072°56.8'E
External/Internal Contacts of Umbra:	U1	09:30:56.5	40°54.7'N	065°10.7'W	064°54.5'W
	U2	09:31:54.3	41°09.4'N	065°32.4'W	065°16.2'W
	U3	12:36:35.4	17°40.4'N	087°09.6'E	087°25.8'E
	U4	12:37:27.8	17°26.7'N	086°52.4'E	087°08.6'E
Extreme North/South Limits of Umbral Path:	N1	09:31:42.6	41°16.4'N	065°33.0'W	065°16.8'W
	S1	09:31:08.3	40°47.7'N	065°10.2'W	064°54.0'W
	N2	12:36:46.0	17°46.9'N	087°09.0'E	087°25.2'E
	S2	12:37:17.0	17°20.1'N	086°53.1'E	087°09.3'E
Extreme Limits of Center Line:	C1	09:31:25.4	41°02.0'N	065°21.5'W	065°05.4'W
	C2	12:37:01.6	17°33.5'N	087°01.0'E	087°17.2'E
Instant of Greatest Eclipse:	G0	11:04:09.0	45°04.5'N	024°01.8'E	024°18.0'E
Circumstances at Greatest Eclipse:	Sun's Altitude = 59.3°		Path Width = 112.3 km		
	Sun's Azimuth = 196.7°		Central Duration = 02m22.9s		

† Ephemeris Longitude is the terrestrial dynamical longitude assuming a uniformly rotating Earth.

* True Longitude is calculated by correcting the Ephemeris Longitude for the non-uniform rotation of Earth.

(T.L. = E.L. - 1.002738* ΔT /240, where ΔT (in seconds) = TDT - UT)

Note: Longitude is measured positive to the East.

Since ΔT is not known in advance, the value used in the predictions is an extrapolation based on pre-1997 measurements. Nevertheless, the actual value is expected to fall within ± 0.3 seconds of the estimated ΔT used here.

TABLE 3

**PATH OF THE UMBRAL SHADOW
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Universal Time	Northern Limit		Southern Limit		Center Line		Sun Alt °	Path Width km	Central Durat. s
	Latitude	Longitude	Latitude	Longitude	Latitude	Longitude			
Limits	41°16.4'N	065°16.8'W	40°47.7'N	064°54.0'W	41°02.0'N	065°05.4'W	0	61	00m46.5s
09:35	46°08.0'N	046°39.6'W	45°48.4'N	044°45.1'W	45°58.6'N	045°41.4'W	16	79	01m09.1s
09:40	47°57.4'N	037°45.4'W	47°25.2'N	036°16.6'W	47°41.6'N	037°00.3'W	22	86	01m20.4s
09:45	49°04.2'N	030°56.5'W	48°24.6'N	029°42.0'W	48°44.6'N	030°18.7'W	28	91	01m29.4s
09:50	49°48.1'N	025°09.5'W	49°03.2'N	024°06.1'W	49°25.8'N	024°37.3'W	32	94	01m37.0s
09:55	50°16.4'N	020°01.7'W	49°27.5'N	019°08.1'W	49°52.0'N	019°34.4'W	36	97	01m43.8s
10:00	50°32.9'N	015°22.0'W	49°40.9'N	014°37.4'W	50°07.0'N	014°59.3'W	39	100	01m49.9s
10:05	50°40.1'N	011°04.0'W	49°45.6'N	010°27.9'W	50°12.9'N	010°45.6'W	42	102	01m55.4s
10:10	50°39.4'N	007°03.7'W	49°43.0'N	006°35.7'W	50°11.3'N	006°49.4'W	45	103	02m00.3s
10:15	50°32.0'N	003°18.4'W	49°34.2'N	002°58.2'W	50°03.1'N	003°08.0'W	47	105	02m04.7s
10:20	50°18.8'N	000°13.9'E	49°20.0'N	000°26.8'E	49°49.4'N	000°20.6'E	50	106	02m08.6s
10:25	50°00.3'N	003°34.8'E	49°00.9'N	003°40.8'E	49°30.6'N	003°37.9'E	52	107	02m12.1s
10:30	49°37.1'N	006°45.4'E	48°37.5'N	006°44.9'E	49°07.3'N	006°45.3'E	53	108	02m15.0s
10:35	49°09.8'N	009°46.9'E	48°10.1'N	009°40.3'E	48°39.9'N	009°43.7'E	55	109	02m17.5s
10:40	48°38.6'N	012°40.2'E	47°39.1'N	012°27.9'E	48°08.9'N	012°34.1'E	56	110	02m19.5s
10:45	48°03.9'N	015°26.0'E	47°04.9'N	015°08.4'E	47°34.4'N	015°17.2'E	58	111	02m21.1s
10:50	47°25.9'N	018°05.1'E	46°27.5'N	017°42.7'E	46°56.7'N	017°53.9'E	58	111	02m22.2s
10:55	46°44.8'N	020°38.3'E	45°47.3'N	020°11.5'E	46°16.1'N	020°24.8'E	59	112	02m22.8s
11:00	46°00.9'N	023°06.2'E	45°04.3'N	022°35.3'E	45°32.6'N	022°50.6'E	59	112	02m23.0s
11:05	45°14.2'N	025°29.4'E	44°18.7'N	024°54.8'E	44°46.5'N	025°12.0'E	59	112	02m22.7s
11:10	44°24.8'N	027°48.7'E	43°30.6'N	027°10.7'E	43°57.8'N	027°29.6'E	59	112	02m22.0s
11:15	43°32.9'N	030°04.6'E	42°40.1'N	029°23.6'E	43°06.5'N	029°44.0'E	59	113	02m20.9s
11:20	42°38.5'N	032°17.9'E	41°47.1'N	031°34.2'E	42°12.9'N	031°55.9'E	58	112	02m19.3s
11:25	41°41.6'N	034°29.3'E	40°51.7'N	033°43.1'E	41°16.7'N	034°06.0'E	57	112	02m17.3s
11:30	40°42.2'N	036°39.4'E	39°53.9'N	035°51.1'E	40°18.1'N	036°15.1'E	55	112	02m14.9s
11:35	39°40.2'N	038°49.2'E	38°53.6'N	037°58.8'E	39°17.0'N	038°23.8'E	54	111	02m12.0s
11:40	38°35.5'N	040°59.4'E	37°50.7'N	040°07.3'E	38°13.2'N	040°33.2'E	52	111	02m08.8s
11:45	37°28.0'N	043°11.2'E	36°45.1'N	042°17.4'E	37°06.6'N	042°44.1'E	50	110	02m05.1s
11:50	36°17.4'N	045°25.7'E	35°36.4'N	044°30.3'E	35°57.0'N	044°57.8'E	48	109	02m01.0s
11:55	35°03.4'N	047°44.3'E	34°24.5'N	046°47.5'E	34°44.0'N	047°15.7'E	46	107	01m56.5s
12:00	33°45.5'N	050°08.9'E	33°08.9'N	049°10.7'E	33°27.3'N	049°39.6'E	43	105	01m51.5s
12:05	32°23.2'N	052°41.9'E	31°49.0'N	051°42.3'E	32°06.2'N	052°11.9'E	40	103	01m46.0s
12:10	30°55.5'N	055°26.7'E	30°24.0'N	054°25.4'E	30°39.9'N	054°55.9'E	37	100	01m40.0s
12:15	29°21.0'N	058°28.3'E	28°52.4'N	057°25.1'E	29°06.9'N	057°56.5'E	33	97	01m33.4s
12:20	27°37.3'N	061°55.1'E	27°12.2'N	060°49.0'E	27°24.9'N	061°21.8'E	29	93	01m26.1s
12:25	25°39.8'N	066°02.6'E	25°19.3'N	064°51.8'E	25°29.7'N	065°27.0'E	24	87	01m17.7s
12:30	23°17.4'N	071°29.4'E	23°03.9'N	070°08.0'E	23°10.9'N	070°48.3'E	18	79	01m07.4s
12:35	19°31.1'N	081°43.2'E	19°42.9'N	079°13.7'E	19°38.5'N	080°24.0'E	7	65	00m51.6s
Limits	17°46.9'N	087°25.2'E	17°20.1'N	087°09.3'E	17°33.5'N	087°17.2'E	0	55	00m42.3s

TABLE 4

**PHYSICAL EPHEMERIS OF THE UMBRAL SHADOW
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Universal Time	Center Line		Diameter Ratio	Eclipse Obscur.	Sun Alt °	Sun Azim °	Path Width km	Major Axis km	Minor Axis km	Umbral Veloc. km/s	Central Durat.
	Latitude	Longitude									
09:30.3	41°02.0'N	065°05.4'W	1.0143	1.0287	0.0	69.5	60.8	-	49.0	-	00m46.5s
09:35	45°58.6'N	045°41.4'W	1.0189	1.0382	15.6	83.9	78.9	241.0	64.6	2.856	01m09.1s
09:40	47°41.6'N	037°00.3'W	1.0208	1.0421	22.5	91.6	85.8	186.4	71.1	1.922	01m20.4s
09:45	48°44.6'N	030°18.7'W	1.0222	1.0449	27.6	98.3	90.5	163.4	75.7	1.531	01m29.4s
09:50	49°25.8'N	024°37.3'W	1.0233	1.0472	32.0	104.4	94.1	150.2	79.4	1.305	01m37.0s
09:55	49°52.0'N	019°34.4'W	1.0243	1.0491	35.7	110.3	97.1	141.3	82.5	1.156	01m43.8s
10:00	50°07.0'N	014°59.3'W	1.0250	1.0507	39.1	116.1	99.5	135.0	85.0	1.049	01m49.9s
10:05	50°12.9'N	010°45.6'W	1.0257	1.0520	42.1	121.8	101.6	130.2	87.3	0.969	01m55.4s
10:10	50°11.3'N	006°49.4'W	1.0263	1.0532	44.8	127.6	103.3	126.4	89.1	0.906	02m00.3s
10:15	50°03.1'N	003°08.0'W	1.0268	1.0542	47.3	133.4	104.9	123.4	90.8	0.857	02m04.7s
10:20	49°49.4'N	000°20.6'E	1.0272	1.0551	49.6	139.3	106.2	121.0	92.2	0.816	02m08.6s
10:25	49°30.6'N	003°37.9'E	1.0275	1.0558	51.6	145.4	107.4	119.0	93.4	0.784	02m12.1s
10:30	49°07.3'N	006°45.3'E	1.0278	1.0565	53.4	151.6	108.4	117.4	94.4	0.757	02m15.0s
10:35	48°39.9'N	009°43.7'E	1.0281	1.0570	55.0	158.0	109.3	116.1	95.2	0.735	02m17.5s
10:40	48°08.9'N	012°34.1'E	1.0283	1.0574	56.4	164.6	110.1	115.0	95.8	0.718	02m19.5s
10:45	47°34.4'N	015°17.2'E	1.0284	1.0577	57.5	171.4	110.7	114.1	96.3	0.704	02m21.1s
10:50	46°56.7'N	017°53.9'E	1.0285	1.0579	58.4	178.3	111.3	113.5	96.6	0.693	02m22.2s
10:55	46°16.1'N	020°24.8'E	1.0286	1.0580	59.0	185.3	111.7	113.0	96.8	0.686	02m22.8s
11:00	45°32.6'N	022°50.6'E	1.0286	1.0580	59.3	192.4	112.1	112.7	96.9	0.681	02m23.0s
11:05	44°46.5'N	025°12.0'E	1.0286	1.0580	59.3	199.4	112.3	112.5	96.8	0.679	02m22.7s
11:10	43°57.8'N	027°29.6'E	1.0285	1.0578	59.1	206.4	112.5	112.5	96.5	0.680	02m22.0s
11:15	43°06.5'N	029°44.0'E	1.0284	1.0576	58.6	213.1	112.5	112.6	96.1	0.683	02m20.9s
11:20	42°12.9'N	031°55.9'E	1.0282	1.0572	57.8	219.6	112.5	113.0	95.6	0.690	02m19.3s
11:25	41°16.7'N	034°06.0'E	1.0280	1.0568	56.7	225.8	112.2	113.5	94.9	0.699	02m17.3s
11:30	40°18.1'N	036°15.1'E	1.0278	1.0563	55.5	231.6	111.9	114.2	94.1	0.712	02m14.9s
11:35	39°17.0'N	038°23.8'E	1.0274	1.0556	53.9	237.1	111.4	115.1	93.0	0.728	02m12.0s
11:40	38°13.2'N	040°33.2'E	1.0271	1.0549	52.2	242.2	110.7	116.3	91.8	0.750	02m08.8s
11:45	37°06.6'N	042°44.1'E	1.0267	1.0540	50.2	247.0	109.8	117.8	90.5	0.776	02m05.1s
11:50	35°57.0'N	044°57.8'E	1.0262	1.0531	48.0	251.5	108.6	119.7	88.9	0.810	02m01.0s
11:55	34°44.0'N	047°15.7'E	1.0256	1.0519	45.5	255.7	107.2	122.0	87.1	0.852	01m56.5s
12:00	33°27.3'N	049°39.6'E	1.0250	1.0507	42.9	259.5	105.3	125.0	85.0	0.905	01m51.5s
12:05	32°06.2'N	052°11.9'E	1.0243	1.0492	39.9	263.2	103.1	128.9	82.7	0.974	01m46.0s
12:10	30°39.9'N	054°55.9'E	1.0235	1.0475	36.7	266.6	100.3	134.0	79.9	1.067	01m40.0s
12:15	29°06.9'N	057°56.5'E	1.0225	1.0456	33.0	269.9	96.9	141.0	76.8	1.196	01m33.4s
12:20	27°24.9'N	061°21.8'E	1.0214	1.0433	28.9	273.1	92.5	151.5	73.1	1.390	01m26.1s
12:25	25°29.7'N	065°27.0'E	1.0200	1.0405	23.9	276.2	86.9	168.9	68.5	1.714	01m17.7s
12:30	23°10.9'N	070°48.3'E	1.0182	1.0368	17.7	279.5	79.0	206.0	62.4	2.412	01m07.4s
12:35	19°38.5'N	080°24.0'E	1.0151	1.0304	7.1	283.8	64.8	422.7	51.7	6.535	00m51.6s
12:35.9	17°33.5'N	087°17.2'E	1.0130	1.0261	0.0	286.1	55.2	-	44.6	-	00m42.3s

TABLE 5

**LOCAL CIRCUMSTANCES ON THE CENTER LINE
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Center Line Maximum Eclipse			First Contact				Second Contact			Third Contact			Fourth Contact			
U.T.	Durat.	Alt	U.T.	P	V	Alt	U.T.	P	V	U.T.	P	V	U.T.	P	V	Alt
		°		°	°	°		°	°		°	°		°	°	°
09:35	01m09.1s	16	08:36:00	277	321	6	09:34:26	98	144	09:35:35	278	324	10:39:27	98	144	27
09:40	01m20.4s	22	08:37:49	278	322	12	09:39:20	98	143	09:40:40	278	323	10:48:07	99	142	34
09:45	01m29.4s	28	08:40:13	279	322	17	09:44:15	99	142	09:45:45	279	322	10:55:56	100	139	39
09:50	01m37.0s	32	08:42:56	279	322	21	09:49:12	100	141	09:50:49	280	321	11:03:17	101	136	43
09:55	01m43.8s	36	08:45:52	280	322	25	09:54:08	101	140	09:55:52	281	320	11:10:16	102	133	46
10:00	01m49.9s	39	08:48:58	280	321	28	09:59:05	101	138	10:00:55	281	318	11:16:58	103	129	49
10:05	01m55.4s	42	08:52:11	281	321	31	10:04:02	102	137	10:05:58	282	317	11:23:24	103	125	51
10:10	02m00.3s	45	08:55:32	282	321	34	10:09:00	103	135	10:11:00	283	315	11:29:38	104	121	53
10:15	02m04.7s	47	08:59:00	282	320	37	10:13:58	103	133	10:16:02	283	312	11:35:38	105	117	54
10:20	02m08.6s	50	09:02:34	283	320	40	10:18:56	104	130	10:21:04	284	310	11:41:28	106	112	55
10:25	02m12.1s	52	09:06:14	283	319	42	10:23:54	105	128	10:26:06	285	307	11:47:06	106	107	56
10:30	02m15.0s	53	09:10:00	284	318	45	10:28:53	105	125	10:31:08	285	304	11:52:34	107	103	56
10:35	02m17.5s	55	09:13:53	285	317	47	10:33:51	106	121	10:36:09	286	301	11:57:53	107	98	56
10:40	02m19.5s	56	09:17:52	285	315	49	10:38:50	107	118	10:41:10	287	297	12:03:03	108	94	56
10:45	02m21.1s	58	09:21:58	286	314	51	10:43:49	107	114	10:46:11	287	293	12:08:05	108	89	55
10:50	02m22.2s	58	09:26:11	286	312	54	10:48:49	108	109	10:51:11	288	289	12:12:58	109	85	54
10:55	02m22.8s	59	09:30:31	287	310	56	10:53:49	108	105	10:56:11	288	284	12:17:44	109	82	53
11:00	02m23.0s	59	09:34:59	288	307	57	10:58:48	109	100	11:01:11	289	280	12:22:24	110	78	52
11:05	02m22.7s	59	09:39:35	288	304	59	11:03:49	109	96	11:06:11	289	275	12:26:56	110	75	51
11:10	02m22.0s	59	09:44:18	289	300	61	11:08:49	110	91	11:11:11	290	270	12:31:22	110	72	49
11:15	02m20.9s	59	09:49:10	289	296	62	11:13:49	110	86	11:16:10	290	265	12:35:43	111	69	48
11:20	02m19.3s	58	09:54:11	290	291	63	11:18:50	111	81	11:21:10	291	261	12:39:57	111	67	46
11:25	02m17.3s	57	09:59:21	290	286	64	11:23:51	111	77	11:26:09	291	257	12:44:07	111	64	44
11:30	02m14.9s	55	10:04:40	291	280	64	11:28:52	111	73	11:31:07	291	253	12:48:11	111	62	42
11:35	02m12.0s	54	10:10:08	291	273	65	11:33:54	112	69	11:36:06	292	249	12:52:10	111	60	40
11:40	02m08.8s	52	10:15:46	291	267	64	11:38:55	112	66	11:41:04	292	245	12:56:04	112	58	38
11:45	02m05.1s	50	10:21:35	292	260	64	11:43:57	112	62	11:46:02	292	242	12:59:53	112	56	36
11:50	02m01.0s	48	10:27:34	292	254	62	11:48:59	112	59	11:51:00	292	239	13:03:37	112	55	33
11:55	01m56.5s	46	10:33:44	292	248	61	11:54:02	112	56	11:55:58	292	236	13:07:17	112	53	31
12:00	01m51.5s	43	10:40:05	292	243	59	11:59:04	112	54	12:00:56	292	234	13:10:50	111	51	28
12:05	01m46.0s	40	10:46:39	293	238	56	12:04:07	112	51	12:05:53	292	231	13:14:18	111	50	25
12:10	01m40.0s	37	10:53:26	293	233	53	12:09:10	112	49	12:10:50	292	229	13:17:39	111	48	22
12:15	01m33.4s	33	11:00:28	292	229	49	12:14:13	112	47	12:15:47	292	227	13:20:52	111	47	19
12:20	01m26.1s	29	11:07:48	292	226	45	12:19:17	111	44	12:20:43	291	224	13:23:54	110	45	15
12:25	01m17.7s	24	11:15:32	292	222	40	12:24:21	111	42	12:25:39	291	222	13:26:41	110	44	10
12:30	01m07.4s	18	11:23:57	291	219	33	12:29:26	110	40	12:30:34	290	220	13:29:01	109	42	5
12:35	00m51.6s	7	11:34:19	290	215	21	12:34:34	109	37	12:35:26	289	217	-	-	-	-
12:35	00m42.3s	7	11:34:19	290	215	21	12:34:34	109	37	12:35:26	289	217	-	-	-	-
12:35	00m42.3s	7	11:34:19	290	215	21	12:34:34	109	37	12:35:26	289	217	-	-	-	-

TABLE 6

TOPOCENTRIC DATA AND PATH CORRECTIONS DUE TO LUNAR LIMB PROFILE
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Universal Time	Moon Topo H.P.	Moon Topo S.D.	Moon Rel. Ang. ∇ "/s	Topo Lib. Long °	Sun Alt. °	Sun Az. °	Path Az. °	North Limit P.A. °	North Limit		South Limit		Central Durat. Cor. s
									Int.	Ext.	Int.	Ext.	
09:35	3542.7	964.7	0.518	5.54	15.6	83.9	72.6	7.7	-0.5	0.6	0.7	-2.4	1.3
09:40	3549.5	966.5	0.491	5.49	22.5	91.6	75.4	8.5	-0.5	0.8	0.7	-1.7	1.7
09:45	3554.4	967.8	0.471	5.45	27.6	98.3	78.2	9.2	-0.5	0.7	0.7	-0.8	1.7
09:50	3558.2	968.9	0.455	5.41	32.0	104.4	81.0	10.0	-0.4	0.4	0.7	-0.9	1.7
09:55	3561.5	969.7	0.442	5.37	35.7	110.3	83.8	10.7	-0.4	0.3	0.7	-1.7	1.6
10:00	3564.2	970.5	0.431	5.32	39.1	116.1	86.6	11.4	-0.4	0.6	0.7	-2.2	1.6
10:05	3566.5	971.1	0.422	5.28	42.1	121.8	89.3	12.1	-0.4	0.8	0.8	-2.6	1.7
10:10	3568.6	971.7	0.413	5.24	44.8	127.6	92.0	12.8	-0.3	0.8	0.8	-2.8	1.7
10:15	3570.3	972.1	0.406	5.20	47.3	133.4	94.6	13.5	-0.3	0.7	0.8	-2.9	1.7
10:20	3571.8	972.5	0.400	5.16	49.6	139.3	97.1	14.2	-0.3	0.6	0.8	-2.8	1.7
10:25	3573.1	972.9	0.395	5.11	51.6	145.4	99.6	14.8	-0.2	0.5	0.8	-2.6	1.7
10:30	3574.1	973.2	0.390	5.07	53.4	151.6	102.0	15.5	-0.2	0.4	0.8	-2.2	2.0
10:35	3575.0	973.4	0.387	5.03	55.0	158.0	104.3	16.1	-0.2	0.3	0.8	-1.7	2.0
10:40	3575.7	973.6	0.384	4.99	56.4	164.6	106.5	16.7	-0.2	0.2	0.9	-1.5	2.1
10:45	3576.2	973.7	0.382	4.94	57.5	171.4	108.5	17.3	-0.2	0.2	0.9	-1.8	2.0
10:50	3576.6	973.8	0.380	4.90	58.4	178.3	110.5	17.9	-0.2	0.3	0.8	-2.2	1.7
10:55	3576.8	973.9	0.379	4.86	59.0	185.3	112.3	18.4	-0.2	0.5	0.9	-2.4	1.6
11:00	3576.9	973.9	0.379	4.81	59.3	192.4	114.0	18.9	-0.2	0.7	0.9	-2.6	1.6
11:05	3576.8	973.9	0.379	4.77	59.3	199.4	115.6	19.4	-0.2	0.7	0.9	-2.6	1.5
11:10	3576.5	973.8	0.380	4.73	59.1	206.4	117.1	19.9	-0.1	0.6	0.9	-2.6	1.5
11:15	3576.1	973.7	0.382	4.69	58.6	213.1	118.4	20.3	-0.1	0.6	0.8	-2.7	0.9
11:20	3575.5	973.5	0.384	4.64	57.8	219.6	119.5	20.7	-0.1	0.7	0.8	-2.9	0.9
11:25	3574.8	973.3	0.386	4.60	56.7	225.8	120.5	21.0	-0.1	0.8	0.8	-3.0	0.8
11:30	3573.9	973.1	0.390	4.56	55.5	231.6	121.4	21.3	-0.1	0.9	0.8	-3.2	0.7
11:35	3572.8	972.8	0.394	4.52	53.9	237.1	122.0	21.6	-0.1	0.9	0.8	-3.3	0.6
11:40	3571.5	972.5	0.398	4.47	52.2	242.2	122.5	21.8	-0.0	1.0	0.8	-3.5	0.4
11:45	3570.1	972.1	0.404	4.43	50.2	247.0	122.9	21.9	-0.0	1.0	0.8	-3.5	0.3
11:50	3568.4	971.6	0.410	4.39	48.0	251.5	123.0	22.1	-0.0	1.0	0.8	-3.6	0.3
11:55	3566.5	971.1	0.417	4.35	45.5	255.7	122.9	22.1	-0.0	1.0	0.8	-3.6	0.2
12:00	3564.3	970.5	0.425	4.30	42.9	259.5	122.6	22.1	-0.0	1.0	0.8	-3.6	0.2
12:05	3561.8	969.8	0.434	4.26	39.9	263.2	122.1	22.0	-0.0	1.0	0.8	-3.7	-0.1
12:10	3558.9	969.0	0.445	4.22	36.7	266.6	121.2	21.9	-0.0	1.0	0.8	-3.6	-0.1
12:15	3555.6	968.1	0.457	4.18	33.0	269.9	120.1	21.7	-0.0	1.0	0.8	-3.5	-0.2
12:20	3551.6	967.1	0.471	4.13	28.9	273.1	118.7	21.4	-0.1	1.0	0.8	-3.4	-0.3
12:25	3546.8	965.8	0.489	4.09	23.9	276.2	116.7	20.9	-0.1	0.9	0.8	-3.1	-0.5
12:30	3540.5	964.1	0.512	4.05	17.7	279.5	114.0	20.3	-0.1	0.7	0.8	-2.7	-0.7
12:35	3529.5	961.1	0.553	4.01	7.1	283.8	109.2	19.0	-0.1	0.7	0.8	-2.7	-0.8

TABLE 7
MAPPING COORDINATES FOR THE UMBRAL PATH
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Longitude	Latitude of:			Universal Time at:			Circumstances on the Center Line		
	Northern	Southern	Center	Northern	Southern	Center	Sun	Sun Path	Center
	Limit	Limit	Line	Limit	Limit	Line	Alt	Az. Width	Durat.
				h m s	h m s	h m s	°	° km	
065°00.0'W	41°21.17'N	-	41°03.56'N	09:30:20	-	09:30:14	0	-	-
064°00.0'W	41°38.43'N	41°02.72'N	41°20.47'N	09:30:39	09:30:04	09:30:22	1	70 62	00m47.6s
063°00.0'W	41°55.51'N	41°19.26'N	41°37.28'N	09:30:42	09:30:06	09:30:24	2	71 63	00m48.7s
062°00.0'W	42°12.41'N	41°35.66'N	41°53.93'N	09:30:47	09:30:09	09:30:28	2	72 64	00m49.8s
061°00.0'W	42°29.17'N	41°51.84'N	42°10.44'N	09:30:53	09:30:14	09:30:33	3	72 65	00m50.9s
060°00.0'W	42°45.79'N	42°07.92'N	42°26.78'N	09:31:00	09:30:20	09:30:40	4	73 66	00m52.0s
059°00.0'W	43°02.23'N	42°23.87'N	42°42.96'N	09:31:09	09:30:28	09:30:48	5	74 67	00m53.1s
058°00.0'W	43°18.49'N	42°39.60'N	42°58.96'N	09:31:20	09:30:37	09:30:58	6	74 68	00m54.3s
057°00.0'W	43°34.57'N	42°55.16'N	43°14.78'N	09:31:31	09:30:48	09:31:10	7	75 69	00m55.4s
056°00.0'W	43°50.44'N	43°10.53'N	43°30.40'N	09:31:45	09:31:00	09:31:22	7	76 70	00m56.6s
055°00.0'W	44°06.11'N	43°25.71'N	43°45.83'N	09:32:00	09:31:14	09:31:37	8	77 71	00m57.7s
054°00.0'W	44°21.57'N	43°40.69'N	44°01.05'N	09:32:16	09:31:29	09:31:52	9	77 72	00m58.9s
053°00.0'W	44°36.81'N	43°55.47'N	44°16.06'N	09:32:34	09:31:46	09:32:10	10	78 72	01m00.1s
052°00.0'W	44°51.81'N	44°10.08'N	44°30.85'N	09:32:53	09:32:04	09:32:28	11	79 73	01m01.3s
051°00.0'W	45°06.71'N	44°24.31'N	44°45.41'N	09:33:14	09:32:24	09:32:49	11	80 74	01m02.5s
050°00.0'W	45°21.22'N	44°38.40'N	44°59.73'N	09:33:36	09:32:45	09:33:10	12	80 75	01m03.7s
049°00.0'W	45°35.52'N	44°52.27'N	45°13.82'N	09:33:59	09:33:08	09:33:33	13	81 76	01m05.0s
048°00.0'W	45°49.57'N	45°05.89'N	45°27.65'N	09:34:24	09:33:32	09:33:58	14	82 77	01m06.2s
047°00.0'W	46°03.36'N	45°19.26'N	45°41.23'N	09:34:51	09:33:57	09:34:24	15	83 78	01m07.5s
046°00.0'W	46°16.89'N	45°32.37'N	45°54.55'N	09:35:19	09:34:24	09:34:51	15	84 79	01m08.7s
045°00.0'W	46°30.14'N	45°45.22'N	46°07.60'N	09:35:48	09:34:53	09:35:20	16	84 79	01m10.0s
044°00.0'W	46°43.11'N	45°57.80'N	46°20.38'N	09:36:18	09:35:23	09:35:50	17	85 80	01m11.3s
043°00.0'W	46°55.80'N	46°10.10'N	46°32.88'N	09:36:50	09:35:54	09:36:22	18	86 81	01m12.6s
042°00.0'W	47°08.20'N	46°22.12'N	46°45.09'N	09:37:24	09:36:26	09:36:55	19	87 82	01m13.9s
041°00.0'W	47°20.30'N	46°33.85'N	46°57.00'N	09:37:58	09:37:00	09:37:29	19	88 83	01m15.1s
040°00.0'W	47°32.10'N	46°45.28'N	47°08.62'N	09:38:34	09:37:36	09:38:05	20	89 83	01m16.5s
039°00.0'W	47°43.58'N	46°56.42'N	47°19.93'N	09:39:12	09:38:13	09:38:42	21	90 84	01m17.8s
038°00.0'W	47°54.76'N	47°07.25'N	47°30.94'N	09:39:50	09:38:51	09:39:20	22	91 85	01m19.1s
037°00.0'W	48°05.61'N	47°17.77'N	47°41.62'N	09:40:30	09:39:31	09:40:00	22	92 86	01m20.4s
036°00.0'W	48°16.14'N	47°27.97'N	47°51.99'N	09:41:12	09:40:11	09:40:41	23	93 87	01m21.7s
035°00.0'W	48°26.34'N	47°37.84'N	48°02.03'N	09:41:54	09:40:54	09:41:24	24	94 87	01m23.1s
034°00.0'W	48°36.20'N	47°47.39'N	48°11.74'N	09:42:38	09:41:37	09:42:08	25	95 88	01m24.4s
033°00.0'W	48°45.72'N	47°56.61'N	48°21.11'N	09:43:23	09:42:22	09:42:53	26	96 89	01m25.7s
032°00.0'W	48°54.90'N	48°05.49'N	48°30.14'N	09:44:10	09:43:09	09:43:39	26	97 89	01m27.1s
031°00.0'W	49°03.73'N	48°14.03'N	48°38.82'N	09:44:57	09:43:56	09:44:26	27	98 90	01m28.4s
030°00.0'W	49°12.20'N	48°22.22'N	48°47.16'N	09:45:46	09:44:45	09:45:15	28	99 91	01m29.8s
029°00.0'W	49°20.31'N	48°30.05'N	48°55.14'N	09:46:36	09:45:35	09:46:06	29	100 91	01m31.1s
028°00.0'W	49°28.07'N	48°37.53'N	49°02.75'N	09:47:28	09:46:27	09:46:57	29	101 92	01m32.5s
027°00.0'W	49°35.45'N	48°44.65'N	49°10.01'N	09:48:20	09:47:20	09:47:50	30	102 93	01m33.8s
026°00.0'W	49°42.46'N	48°51.40'N	49°16.89'N	09:49:14	09:48:14	09:48:44	31	103 93	01m35.2s
025°00.0'W	49°49.10'N	48°57.78'N	49°23.40'N	09:50:09	09:49:09	09:49:39	32	104 94	01m36.5s
024°00.0'W	49°55.36'N	49°03.78'N	49°29.53'N	09:51:05	09:50:06	09:50:35	32	105 95	01m37.9s
023°00.0'W	50°01.23'N	49°09.40'N	49°35.28'N	09:52:02	09:51:04	09:51:33	33	106 95	01m39.2s
022°00.0'W	50°06.71'N	49°14.63'N	49°40.64'N	09:53:01	09:52:03	09:52:32	34	107 96	01m40.6s
021°00.0'W	50°11.80'N	49°19.48'N	49°45.61'N	09:54:01	09:53:04	09:53:32	35	109 96	01m41.9s
020°00.0'W	50°16.50'N	49°23.93'N	49°50.18'N	09:55:02	09:54:05	09:54:33	35	110 97	01m43.3s

TABLE 7 - continued
MAPPING COORDINATES FOR THE UMBRAL PATH
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Longitude	Latitude of:			Universal Time at:			Circumstances on the Center Line					
	Northern	Southern	Center	Northern	Southern	Center	Sun	Sun	Path	Center		
	Limit	Limit	Line	Limit	Limit	Line	Alt	Az.	Width	Durat.		
							°	°	km			
				h	m	s	h	m	s	h	m	s
019°00.0'W	50°20.79'N	49°27.98'N	49°54.36'N	09:56:04	09:55:09	09:55:36	36	111	97	01m44.6s		
018°00.0'W	50°24.68'N	49°31.63'N	49°58.13'N	09:57:07	09:56:13	09:56:40	37	112	98	01m45.9s		
017°00.0'W	50°28.16'N	49°34.87'N	50°01.49'N	09:58:12	09:57:19	09:57:45	38	114	98	01m47.3s		
016°00.0'W	50°31.22'N	49°37.69'N	50°04.43'N	09:59:18	09:58:26	09:58:52	38	115	99	01m48.6s		
015°00.0'W	50°33.86'N	49°40.10'N	50°06.96'N	10:00:25	09:59:34	09:59:59	39	116	100	01m49.9s		
014°00.0'W	50°36.08'N	49°42.08'N	50°09.06'N	10:01:33	10:00:43	10:01:08	40	117	100	01m51.2s		
013°00.0'W	50°37.88'N	49°43.63'N	50°10.74'N	10:02:42	10:01:54	10:02:18	40	119	100	01m52.5s		
012°00.0'W	50°39.24'N	49°44.75'N	50°11.98'N	10:03:53	10:03:07	10:03:30	41	120	101	01m53.8s		
011°00.0'W	50°40.16'N	49°45.43'N	50°12.78'N	10:05:05	10:04:20	10:04:42	42	122	101	01m55.1s		
010°00.0'W	50°40.64'N	49°45.67'N	50°13.14'N	10:06:18	10:05:35	10:05:56	43	123	102	01m56.4s		
009°00.0'W	50°40.67'N	49°45.45'N	50°13.05'N	10:07:32	10:06:51	10:07:12	43	124	102	01m57.6s		
008°00.0'W	50°40.24'N	49°44.78'N	50°12.50'N	10:08:48	10:08:09	10:08:28	44	126	103	01m58.9s		
007°00.0'W	50°39.36'N	49°43.64'N	50°11.49'N	10:10:05	10:09:28	10:09:46	45	127	103	02m00.1s		
006°00.0'W	50°38.01'N	49°42.03'N	50°10.02'N	10:11:23	10:10:48	10:11:05	45	129	104	02m01.3s		
005°00.0'W	50°36.20'N	49°39.95'N	50°08.07'N	10:12:42	10:12:10	10:12:26	46	130	104	02m02.5s		
004°00.0'W	50°33.90'N	49°37.39'N	50°05.64'N	10:14:03	10:13:33	10:13:48	47	132	105	02m03.7s		
003°00.0'W	50°31.12'N	49°34.34'N	50°02.73'N	10:15:25	10:14:57	10:15:11	47	134	105	02m04.9s		
002°00.0'W	50°27.85'N	49°30.80'N	49°59.32'N	10:16:49	10:16:23	10:16:36	48	135	105	02m06.0s		
001°00.0'W	50°24.09'N	49°26.75'N	49°55.42'N	10:18:14	10:17:51	10:18:02	49	137	106	02m07.2s		
000°00.0'E	50°19.82'N	49°22.19'N	49°51.00'N	10:19:40	10:19:20	10:19:30	49	139	106	02m08.3s		
001°00.0'E	50°15.03'N	49°17.11'N	49°46.07'N	10:21:07	10:20:50	10:20:59	50	141	106	02m09.3s		
002°00.0'E	50°09.73'N	49°11.50'N	49°40.62'N	10:22:36	10:22:22	10:22:29	51	142	107	02m10.4s		
003°00.0'E	50°03.90'N	49°05.37'N	49°34.64'N	10:24:07	10:23:56	10:24:01	51	144	107	02m11.4s		
004°00.0'E	49°57.54'N	48°58.68'N	49°28.12'N	10:25:39	10:25:31	10:25:35	52	146	108	02m12.4s		
005°00.0'E	49°50.63'N	48°51.45'N	49°21.05'N	10:27:12	10:27:07	10:27:10	52	148	108	02m13.4s		
006°00.0'E	49°43.17'N	48°43.66'N	49°13.43'N	10:28:47	10:28:45	10:28:46	53	150	108	02m14.3s		
007°00.0'E	49°35.15'N	48°35.30'N	49°05.23'N	10:30:24	10:30:25	10:30:24	54	152	108	02m15.2s		
008°00.0'E	49°26.56'N	48°26.36'N	48°56.47'N	10:32:02	10:32:07	10:32:04	54	154	109	02m16.1s		
009°00.0'E	49°17.39'N	48°16.83'N	48°47.12'N	10:33:41	10:33:50	10:33:45	55	156	109	02m16.9s		
010°00.0'E	49°07.63'N	48°06.70'N	48°37.18'N	10:35:22	10:35:35	10:35:28	55	159	109	02m17.7s		
011°00.0'E	48°57.27'N	47°55.97'N	48°26.64'N	10:37:05	10:37:21	10:37:13	56	161	110	02m18.4s		
012°00.0'E	48°46.30'N	47°44.62'N	48°15.48'N	10:38:49	10:39:09	10:38:59	56	163	110	02m19.1s		
013°00.0'E	48°34.72'N	47°32.65'N	48°03.70'N	10:40:35	10:40:59	10:40:47	57	166	110	02m19.8s		
014°00.0'E	48°22.51'N	47°20.04'N	47°51.29'N	10:42:23	10:42:51	10:42:36	57	168	110	02m20.4s		
015°00.0'E	48°09.66'N	47°06.78'N	47°38.24'N	10:44:12	10:44:44	10:44:28	57	171	111	02m20.9s		
016°00.0'E	47°56.16'N	46°52.87'N	47°24.54'N	10:46:03	10:46:39	10:46:21	58	173	111	02m21.4s		
017°00.0'E	47°42.01'N	46°38.30'N	47°10.17'N	10:47:56	10:48:36	10:48:15	58	176	111	02m21.8s		
018°00.0'E	47°27.19'N	46°23.06'N	46°55.14'N	10:49:50	10:50:34	10:50:12	58	179	111	02m22.2s		
019°00.0'E	47°11.69'N	46°07.13'N	46°39.43'N	10:51:46	10:52:35	10:52:10	59	181	111	02m22.5s		
020°00.0'E	46°55.51'N	45°50.52'N	46°23.03'N	10:53:44	10:54:37	10:54:10	59	184	112	02m22.7s		

TABLE 7 - continued
MAPPING COORDINATES FOR THE UMBRAL PATH
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Longitude	Latitude of:			Universal Time at:			Circumstances on the Center Line			
	Northern	Southern	Center	Northern	Southern	Center	Sun	Sun	Path	Center
	Limit	Limit	Line	Limit	Limit	Line	Alt	Az.	Width	Durat.
							°	°	km	
	h	m	s	h	m	s	h	m	s	
021°00.0'E	46°38.64'N	45°33.22'N	46°05.94'N	10:55:43	10:56:40	10:56:12	59	187	112	02m22.9s
022°00.0'E	46°21.07'N	45°15.21'N	45°48.15'N	10:57:45	10:58:46	10:58:15	59	190	112	02m23.0s
023°00.0'E	46°02.79'N	44°56.50'N	45°29.66'N	10:59:47	11:00:53	11:00:20	59	193	112	02m23.0s
024°00.0'E	45°43.80'N	44°37.08'N	45°10.45'N	11:01:52	11:03:01	11:02:26	59	196	112	02m22.9s
025°00.0'E	45°24.09'N	44°16.95'N	44°50.53'N	11:03:58	11:05:11	11:04:34	59	199	112	02m22.8s
026°00.0'E	45°03.67'N	43°56.10'N	44°29.89'N	11:06:05	11:07:23	11:06:44	59	202	112	02m22.5s
027°00.0'E	44°42.52'N	43°34.54'N	44°08.54'N	11:08:14	11:09:36	11:08:55	59	205	112	02m22.2s
028°00.0'E	44°20.65'N	43°12.28'N	43°46.47'N	11:10:25	11:11:51	11:11:07	59	208	113	02m21.8s
029°00.0'E	43°58.07'N	42°49.31'N	43°23.69'N	11:12:37	11:14:06	11:13:21	59	211	113	02m21.3s
030°00.0'E	43°34.77'N	42°25.65'N	43°00.20'N	11:14:50	11:16:23	11:15:36	58	214	113	02m20.7s
031°00.0'E	43°10.76'N	42°01.29'N	42°36.02'N	11:17:04	11:18:41	11:17:52	58	217	112	02m20.0s
032°00.0'E	42°46.05'N	41°36.27'N	42°11.15'N	11:19:19	11:21:00	11:20:09	58	220	112	02m19.2s
033°00.0'E	42°20.66'N	41°10.58'N	41°45.60'N	11:21:36	11:23:19	11:22:27	57	223	112	02m18.4s
034°00.0'E	41°54.59'N	40°44.26'N	41°19.40'N	11:23:53	11:25:39	11:24:46	57	225	112	02m17.4s
035°00.0'E	41°27.87'N	40°17.32'N	40°52.57'N	11:26:11	11:28:00	11:27:05	56	228	112	02m16.3s
036°00.0'E	41°00.52'N	39°49.79'N	40°25.12'N	11:28:29	11:30:21	11:29:25	56	231	112	02m15.2s
037°00.0'E	40°32.55'N	39°21.69'N	39°57.08'N	11:30:48	11:32:42	11:31:45	55	234	112	02m13.9s
038°00.0'E	40°04.01'N	38°53.06'N	39°28.49'N	11:33:06	11:35:03	11:34:05	54	236	111	02m12.6s
039°00.0'E	39°34.91'N	38°23.93'N	38°59.37'N	11:35:25	11:37:23	11:36:24	53	239	111	02m11.1s
040°00.0'E	39°05.31'N	37°54.34'N	38°29.76'N	11:37:43	11:39:43	11:38:43	53	241	111	02m09.6s
041°00.0'E	38°35.22'N	37°24.32'N	37°59.70'N	11:40:01	11:42:02	11:41:02	52	243	111	02m08.0s
042°00.0'E	38°04.69'N	36°53.94'N	37°29.23'N	11:42:18	11:44:20	11:43:19	51	245	110	02m06.4s
043°00.0'E	37°33.77'N	36°23.21'N	36°58.40'N	11:44:35	11:46:37	11:45:36	50	248	110	02m04.6s
044°00.0'E	37°02.50'N	35°52.20'N	36°27.26'N	11:46:50	11:48:52	11:47:51	49	250	109	02m02.8s
045°00.0'E	36°30.93'N	35°20.96'N	35°55.84'N	11:49:03	11:51:06	11:50:05	48	252	109	02m00.9s
046°00.0'E	35°59.11'N	34°49.52'N	35°24.21'N	11:51:15	11:53:17	11:52:17	47	253	108	01m59.0s
047°00.0'E	35°27.08'N	34°17.95'N	34°52.40'N	11:53:25	11:55:27	11:54:26	46	255	107	01m57.0s
048°00.0'E	34°54.90'N	33°46.29'N	34°20.47'N	11:55:33	11:57:34	11:56:34	45	257	107	01m54.9s
049°00.0'E	34°22.62'N	33°14.59'N	33°48.48'N	11:57:39	11:59:38	11:58:39	44	259	106	01m52.9s
050°00.0'E	33°50.30'N	32°42.90'N	33°16.46'N	11:59:42	12:01:40	12:00:41	42	260	105	01m50.8s
051°00.0'E	33°17.97'N	32°11.26'N	32°44.48'N	12:01:42	12:03:38	12:02:41	41	262	104	01m48.6s
052°00.0'E	32°45.69'N	31°39.73'N	32°12.56'N	12:03:40	12:05:34	12:04:37	40	263	103	01m46.5s
053°00.0'E	32°13.51'N	31°08.34'N	31°40.78'N	12:05:34	12:07:26	12:06:31	39	264	102	01m44.3s
054°00.0'E	31°41.47'N	30°37.15'N	31°09.15'N	12:07:25	12:09:15	12:08:21	38	266	101	01m42.1s
055°00.0'E	31°09.61'N	30°06.17'N	30°37.74'N	12:09:13	12:11:00	12:10:07	37	267	100	01m39.9s
056°00.0'E	30°37.98'N	29°35.46'N	30°06.56'N	12:10:58	12:12:42	12:11:50	35	268	99	01m37.7s
057°00.0'E	30°06.62'N	29°05.05'N	29°35.67'N	12:12:38	12:14:20	12:13:30	34	269	98	01m35.5s
058°00.0'E	29°35.54'N	28°34.97'N	29°05.10'N	12:14:15	12:15:55	12:15:05	33	270	97	01m33.3s
059°00.0'E	29°04.81'N	28°05.24'N	28°34.86'N	12:15:49	12:17:25	12:16:38	32	271	96	01m31.1s
060°00.0'E	28°34.43'N	27°35.90'N	28°05.00'N	12:17:19	12:18:52	12:18:06	31	272	94	01m29.0s

TABLE 7 - continued
MAPPING COORDINATES FOR THE UMBRAL PATH
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Longitude	Latitude of:			Universal Time at:			Circumstances on the Center Line					
	Northern	Southern	Center	Northern	Southern	Center	Sun	Sun	Path	Center		
	Limit	Limit	Line	Limit	Limit	Line	Alt	Az.	Width	Durat.		
							°	°	km			
				h	m	s	h	m	s	h	m	s
061°00.0'E	28°04.43'N	27°06.95'N	27°35.53'N	12:18:45	12:20:15	12:19:30	29	273	93	01m26.9s		
062°00.0'E	27°34.85'N	26°38.43'N	27°06.49'N	12:20:07	12:21:34	12:20:51	28	274	92	01m24.7s		
063°00.0'E	27°05.70'N	26°10.36'N	26°37.87'N	12:21:25	12:22:50	12:22:08	27	274	90	01m22.7s		
064°00.0'E	26°36.99'N	25°42.74'N	26°09.71'N	12:22:39	12:24:01	12:23:21	26	275	89	01m20.6s		
065°00.0'E	26°08.75'N	25°15.58'N	25°42.01'N	12:23:50	12:25:09	12:24:30	24	276	88	01m18.6s		
066°00.0'E	25°40.98'N	24°48.91'N	25°14.79'N	12:24:57	12:26:13	12:25:36	23	277	86	01m16.6s		
067°00.0'E	25°13.70'N	24°22.72'N	24°48.07'N	12:26:00	12:27:14	12:26:37	22	277	85	01m14.6s		
068°00.0'E	24°46.92'N	23°57.02'N	24°21.83'N	12:27:00	12:28:11	12:27:36	21	278	83	01m12.7s		
069°00.0'E	24°20.65'N	23°31.83'N	23°56.10'N	12:27:56	12:29:04	12:28:30	20	278	82	01m10.8s		
070°00.0'E	23°54.88'N	23°07.13'N	23°30.87'N	12:28:48	12:29:54	12:29:21	19	279	80	01m08.9s		
071°00.0'E	23°29.62'N	22°42.94'N	23°06.15'N	12:29:37	12:30:40	12:30:09	17	280	79	01m07.1s		
072°00.0'E	23°04.88'N	22°19.26'N	22°41.94'N	12:30:23	12:31:23	12:30:53	16	280	77	01m05.3s		
073°00.0'E	22°40.65'N	21°56.08'N	22°18.24'N	12:31:05	12:32:03	12:31:34	15	281	76	01m03.5s		
074°00.0'E	22°16.94'N	21°33.40'N	21°55.05'N	12:31:44	12:32:39	12:32:12	14	281	74	01m01.8s		
075°00.0'E	21°53.74'N	21°11.22'N	21°32.36'N	12:32:19	12:33:12	12:32:46	13	282	73	01m00.1s		
076°00.0'E	21°31.06'N	20°49.54'N	21°10.19'N	12:32:52	12:33:43	12:33:17	12	282	71	00m58.5s		
077°00.0'E	21°08.88'N	20°28.36'N	20°48.51'N	12:33:21	12:34:10	12:33:46	11	282	70	00m56.8s		
078°00.0'E	20°47.22'N	20°07.66'N	20°27.33'N	12:33:47	12:34:34	12:34:11	10	283	68	00m55.3s		
079°00.0'E	20°26.05'N	19°47.45'N	20°06.65'N	12:34:11	12:34:56	12:34:33	9	283	67	00m53.7s		
080°00.0'E	20°05.39'N	19°27.70'N	19°46.45'N	12:34:31	12:35:14	12:34:53	7	284	65	00m52.2s		
081°00.0'E	19°45.25'N	19°08.40'N	19°26.74'N	12:34:49	12:35:30	12:35:10	6	284	64	00m50.7s		
082°00.0'E	19°25.83'N	18°49.03'N	19°07.51'N	12:35:04	12:35:44	12:35:24	5	284	63	00m49.3s		
083°00.0'E	19°05.99'N	18°31.43'N	18°48.75'N	12:35:17	12:35:54	12:35:35	4	285	61	00m47.9s		
084°00.0'E	18°50.16'N	18°13.57'N	18°30.45'N	12:35:25	12:36:02	12:35:44	3	285	60	00m46.5s		
085°00.0'E	19°09.38'N	17°56.30'N	18°12.62'N	12:35:14	12:36:08	12:35:51	2	285	58	00m45.2s		
086°00.0'E	18°11.49'N	17°39.16'N	17°55.23'N	12:35:38	12:36:11	12:35:55	1	286	57	00m43.9s		

TABLE 8
MAPPING COORDINATES FOR THE ZONES OF GRAZING ECLIPSE
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Longitude	North Graze Zone		Northern	South Graze Zone		Southern	Path Azim	Elev Fact	Scale Fact	km ²
	Latitudes		Limit	Latitudes		Limit				
	Northern Limit	Southern Limit	Universal Time	Northern Limit	Southern Limit	Universal Time				
		h m s			h m s					
005 00.0W	50 36.96N	50 35.87N	10:12:42	49 40.76N	49 37.09N	10:12:10	93.3	-0.58	2.09	
004 00.0W	50 34.64N	50 33.59N	10:14:03	49 38.20N	49 34.52N	10:13:33	94.0	-0.58	2.09	
003 00.0W	50 31.86N	50 30.82N	10:15:25	49 35.16N	49 31.46N	10:14:57	94.7	-0.58	2.09	
002 00.0W	50 28.57N	50 27.56N	10:16:49	49 31.61N	49 27.93N	10:16:23	95.4	-0.58	2.09	
001 00.0W	50 24.78N	50 23.81N	10:18:14	49 27.57N	49 23.90N	10:17:51	96.1	-0.57	2.09	
000 00.0E	50 20.47N	50 19.55N	10:19:40	49 23.01N	49 19.37N	10:19:20	96.9	-0.57	2.08	
001 00.0E	50 15.65N	50 14.78N	10:21:07	49 17.94N	49 14.35N	10:20:50	97.6	-0.57	2.08	
002 00.0E	50 10.32N	50 09.49N	10:22:36	49 12.34N	49 08.80N	10:22:22	98.4	-0.57	2.08	
003 00.0E	50 04.46N	50 03.67N	10:24:07	49 06.21N	49 02.74N	10:23:56	99.1	-0.57	2.08	
004 00.0E	49 58.04N	49 57.30N	10:25:39	48 59.53N	48 56.16N	10:25:31	99.9	-0.57	2.08	
005 00.0E	49 51.09N	49 50.41N	10:27:12	48 52.28N	48 49.02N	10:27:07	100.6	-0.57	2.08	
006 00.0E	49 43.61N	49 42.95N	10:28:47	48 44.49N	48 41.35N	10:28:45	101.4	-0.57	2.08	
007 00.0E	49 35.56N	49 34.92N	10:30:24	48 36.14N	48 33.13N	10:30:25	102.2	-0.56	2.08	
008 00.0E	49 26.93N	49 26.33N	10:32:02	48 27.20N	48 24.35N	10:32:07	102.9	-0.56	2.08	
009 00.0E	49 17.70N	49 17.16N	10:33:41	48 17.67N	48 14.99N	10:33:50	103.7	-0.56	2.08	
010 00.0E	49 07.86N	49 07.40N	10:35:22	48 07.55N	48 05.05N	10:35:35	104.5	-0.56	2.08	
011 00.0E	48 57.49N	48 57.04N	10:37:05	47 56.82N	47 54.37N	10:37:21	105.2	-0.56	2.08	
012 00.0E	48 46.54N	48 46.07N	10:38:49	47 45.48N	47 43.08N	10:39:09	106.0	-0.56	2.08	
013 00.0E	48 34.95N	48 34.49N	10:40:35	47 33.51N	47 31.17N	10:40:59	106.8	-0.56	2.08	
014 00.0E	48 22.73N	48 22.28N	10:42:23	47 20.90N	47 18.39N	10:42:51	107.5	-0.57	2.08	
015 00.0E	48 09.90N	48 09.43N	10:44:12	47 07.65N	47 04.97N	10:44:44	108.3	-0.57	2.08	
016 00.0E	47 56.42N	47 55.93N	10:46:03	46 53.74N	46 50.91N	10:46:39	109.1	-0.57	2.08	
017 00.0E	47 42.27N	47 41.78N	10:47:56	46 39.18N	46 36.20N	10:48:36	109.8	-0.57	2.08	
018 00.0E	47 27.51N	47 26.98N	10:49:50	46 23.91N	46 20.83N	10:50:34	110.6	-0.57	2.08	
019 00.0E	47 12.11N	47 11.50N	10:51:46	46 07.99N	46 04.81N	10:52:35	111.3	-0.57	2.08	
020 00.0E	46 56.01N	46 55.33N	10:53:44	45 51.38N	45 48.11N	10:54:37	112.0	-0.57	2.09	
021 00.0E	46 39.21N	46 38.46N	10:55:43	45 34.07N	45 30.73N	10:56:40	112.7	-0.58	2.09	
022 00.0E	46 21.69N	46 20.89N	10:57:45	45 16.06N	45 12.67N	10:58:46	113.4	-0.58	2.09	
023 00.0E	46 03.44N	46 02.61N	10:59:47	44 57.35N	44 53.92N	11:00:53	114.1	-0.58	2.09	
024 00.0E	45 44.47N	45 43.63N	11:01:52	44 37.93N	44 34.47N	11:03:01	114.8	-0.59	2.10	
025 00.0E	45 24.77N	45 23.93N	11:03:58	44 17.80N	44 14.33N	11:05:11	115.5	-0.59	2.10	
026 00.0E	45 04.33N	45 03.51N	11:06:05	43 56.96N	43 53.49N	11:07:23	116.1	-0.59	2.10	
027 00.0E	44 43.16N	44 42.37N	11:08:14	43 35.40N	43 31.96N	11:09:36	116.8	-0.60	2.11	
028 00.0E	44 21.25N	44 20.51N	11:10:25	43 13.13N	43 09.73N	11:11:51	117.4	-0.60	2.11	
029 00.0E	43 58.65N	43 57.93N	11:12:37	42 50.12N	42 46.67N	11:14:06	118.0	-0.61	2.12	
030 00.0E	43 35.36N	43 34.63N	11:14:50	42 26.46N	42 22.90N	11:16:23	118.5	-0.61	2.12	
031 00.0E	43 11.38N	43 10.63N	11:17:04	42 02.11N	41 58.45N	11:18:41	119.0	-0.62	2.12	
032 00.0E	42 46.70N	42 45.93N	11:19:19	41 37.08N	41 33.34N	11:21:00	119.6	-0.62	2.13	
033 00.0E	42 21.37N	42 20.54N	11:21:36	41 11.40N	41 07.59N	11:23:19	120.0	-0.63	2.14	
034 00.0E	41 55.37N	41 54.49N	11:23:53	40 45.07N	40 41.20N	11:25:40	120.5	-0.63	2.14	
035 00.0E	41 28.70N	41 27.77N	11:26:11	40 18.13N	40 14.21N	11:28:00	120.9	-0.64	2.15	
036 00.0E	41 01.38N	41 00.42N	11:28:29	39 50.60N	39 46.59N	11:30:21	121.3	-0.64	2.15	
037 00.0E	40 33.45N	40 32.46N	11:30:48	39 22.50N	39 18.42N	11:32:42	121.6	-0.65	2.16	
038 00.0E	40 04.94N	40 03.93N	11:33:06	38 53.87N	38 49.72N	11:35:03	121.9	-0.66	2.17	
039 00.0E	39 35.86N	39 34.84N	11:35:25	38 24.72N	38 20.53N	11:37:23	122.2	-0.66	2.17	
040 00.0E	39 06.28N	39 05.26N	11:37:43	37 55.12N	37 50.89N	11:39:43	122.4	-0.67	2.18	

**TABLE 9
LOCAL CIRCUMSTANCES FOR THE UNITED STATES OF AMERICA
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.		Umbral Durat.	
				U.T. h	P m	V s	U.T. h	P m	V s	U.T. h	P m	V s	U.T. h	P m	V s	U.T. h	P m	V s	U.T. h	P m	Obs.	Durat.
CONNECTICUT																						
Bridgeport	41°11'N	073°11'W	3	-	-	-	10:27:42.5	100	149	5	09:59	Rise	-	-	0	69	0.526	0.421				
Hartford	41°46'N	072°41'W	52	-	-	-	10:28:11.9	101	149	5	09:54	Rise	-	-	0	68	0.611	0.520				
New Haven	41°18'N	072°56'W	12	-	-	-	10:27:47.0	100	149	5	09:57	Rise	-	-	0	69	0.556	0.455				
Stamford	41°03'N	073°32'W	11	-	-	-	10:27:137.2	100	149	4	10:00	Rise	-	-	0	69	0.500	0.382				
Waterbury	41°30'N	073°00'W	79	-	-	-	10:27:58.8	100	149	5	09:56	Rise	-	-	0	68	0.581	0.485				
DELAWARE																						
Dover	39°09'N	075°31'W	-	-	-	-	10:26:10.9	98	148	2	10:12	Rise	-	-	0	70	0.252	0.147				
DISTRICT OF COLUMBIA																						
Washington	38°53'N	077°02'W	4	-	-	-	10:26:16.3	98	148	1	10:19	Rise	-	-	0	70	0.142	0.063				
MAINE																						
Augusta	44°18'N	069°46'W	14	-	-	-	10:30:29.7	103	149	8	09:38	Rise	-	-	0	68	0.887	0.860				
Portland	43°40'N	070°17'W	14	-	-	-	10:29:53.1	102	149	8	09:41	Rise	-	-	0	68	0.848	0.811				
MARYLAND																						
Annapolis	38°58'N	076°29'W	-	-	-	-	10:26:13.3	98	148	2	10:17	Rise	-	-	0	70	0.177	0.087				
Baltimore	39°17'N	076°36'W	45	-	-	-	10:26:31.9	99	148	2	10:15	Rise	-	-	0	69	0.209	0.111				
MASSACHUSETTS																						
Boston	42°21'N	071°03'W	5	-	-	-	10:28:37.7	101	149	7	09:48	Rise	-	-	0	68	0.734	0.669				
Springfield	42°07'N	072°33'W	26	-	-	-	10:28:30.8	101	149	6	09:53	Rise	-	-	0	68	0.630	0.543				
Worcester	42°16'N	071°49'W	145	-	-	-	10:28:36.1	101	149	6	09:49	Rise	-	-	0	68	0.714	0.644				
MICHIGAN																						
Sault Ste. Marie	46°28'N	084°22'W	220	-	-	-	10:34:11.6	111	153	0	10:28	Rise	-	-	0	66	0.106	0.041				
NEW HAMPSHIRE																						
Concord	43°12'N	071°32'W	104	-	-	-	10:29:28.7	102	149	7	09:46	Rise	-	-	0	68	0.772	0.715				
NEW JERSEY																						
Elizabeth	40°40'N	074°13'W	6	-	-	-	10:27:21.4	100	148	4	10:04	Rise	-	-	0	69	0.430	0.317				
Jersey City	40°43'N	074°05'W	6	-	-	-	10:27:23.0	100	148	4	10:03	Rise	-	-	0	69	0.442	0.330				
Newark	40°44'N	074°10'W	2	-	-	-	10:27:24.7	100	148	4	10:04	Rise	-	-	0	69	0.434	0.321				
Paterson	40°55'N	074°10'W	30	-	-	-	10:27:35.0	100	149	4	10:02	Rise	-	-	0	69	0.458	0.347				
Trenton	40°13'N	074°44'W	11	-	-	-	10:27:01.3	99	148	3	10:07	Rise	-	-	0	69	0.374	0.259				
NEW YORK																						
Albany	42°39'N	073°45'W	84	-	-	-	10:29:08.6	102	149	5	09:56	Rise	-	-	0	68	0.592	0.497				
Buffalo	42°53'N	078°52'W	215	-	-	-	10:30:08.5	104	150	2	10:15	Rise	-	-	0	68	0.275	0.167				
New York	40°43'N	074°01'W	40	-	-	-	10:27:22.4	100	148	4	10:02	Rise	-	-	0	69	0.461	0.349				
Rochester	43°09'N	077°36'W	167	-	-	-	10:30:08.1	104	150	3	10:10	Rise	-	-	0	68	0.370	0.256				
Syracuse	43°02'N	076°08'W	125	-	-	-	10:29:147.9	104	150	4	10:05	Rise	-	-	0	68	0.457	0.346				
Yonkers	40°57'N	073°54'W	3	-	-	-	10:27:34.6	100	149	4	10:02	Rise	-	-	0	69	0.464	0.354				
OHIO																						
Youngstown	41°05'N	080°38'W	256	-	-	-	10:29:02.7	103	150	0	10:26	Rise	-	-	0	68	0.061	0.018				
PENNSYLVANIA																						
Allentown	40°35'N	075°30'W	78	-	-	-	10:27:29.4	100	149	3	10:08	Rise	-	-	0	69	0.360	0.246				
Erie	42°07'N	080°05'W	209	-	-	-	10:29:45.9	104	150	1	10:22	Rise	-	-	0	68	0.149	0.068				
Harrisburg	40°16'N	076°53'W	111	-	-	-	10:27:28.2	100	149	2	10:14	Rise	-	-	0	69	0.255	0.149				
Philadelphia	39°57'N	075°09'W	112	-	-	-	10:26:50.9	99	148	3	10:09	Rise	-	-	0	69	0.324	0.211				
Pittsburgh	40°26'N	079°59'W	228	-	-	-	10:28:20.3	102	149	0	10:25	Rise	-	-	0	69	0.067	0.021				
Scranton	41°24'N	075°39'W	221	-	-	-	10:28:15.3	101	149	3	10:05	Rise	-	-	0	68	0.416	0.302				
RHODE ISLAND																						
Providence	41°49'N	071°24'W	16	-	-	-	10:28:07.9	100	149	6	09:50	Rise	-	-	0	68	0.690	0.614				
VERMONT																						
Montpelier	44°15'N	072°34'W	148	-	-	-	10:30:32.3	104	150	6	09:47	Rise	-	-	0	67	0.759	0.699				
VIRGINIA																						
Alexandria	38°49'N	077°05'W	-	-	-	-	10:26:13.6	98	148	1	10:19	Rise	-	-	0	70	0.128	0.054				
Chesapeake	38°48'N	076°16'W	-	-	-	-	10:26:01.3	98	148	2	10:16	Rise	-	-	0	70	0.183	0.092				
Hampton	37°02'N	076°21'W	-	-	-	-	10:24:25.7	95	147	1	10:20	Rise	-	-	0	70	0.087	0.031				
Newport News	37°03'N	076°29'W	-	-	-	-	10:24:28.6	95	147	1	10:20	Rise	-	-	0	70	0.079	0.027				
Norfolk	36°50'N	076°17'W	3	-	-	-	10:24:13.7	95	147	1	10:19	Rise	-	-	0	70	0.088	0.031				
Portsmouth	36°50'N	076°19'W	3	-	-	-	10:24:14.2	95	147	1	10:20	Rise	-	-	0	70	0.085	0.030				
Richmond	37°33'N	077°27'W	50	-	-	-	10:25:11.0	96	147	0	10:22	Rise	-	-	0	70	0.062	0.019				
Virginia Beach	36°50'N	075°58'W	-	-	-	-	10:24:08.8	95	147	1	10:19	Rise	-	-	0	70	0.103	0.039				

TABLE 10
LOCAL CIRCUMSTANCES FOR CANADA
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.					
				U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °								
NEW BRUNSWICK Fredericton, NB	45°58' N	066°39' W	9	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:35:49.8	187	230	2	69	0.881	0.852	
NEWFOUNDLAND Saint John's, NF	47°34' N	052°43' W	64	08:39:18.1	273	314	2	—	—	—	—	—	—	—	—	—	—	—	09:36:05.1	187	231	11	80	0.930	0.916	
NOVA SCOTIA Halifax, NS	44°39' N	063°36' W	25	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:33:41.4	187	231	3	71	0.931	0.916	
ONTARIO Etobicoke, ON	43°39' N	079°34' W	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:19 Rise	—	—	—	0	68	0.215	0.116
Hamilton, ON	43°15' N	079°51' W	100	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:19 Rise	—	—	—	0	68	0.211	0.113
Kitchener, ON	43°27' N	080°29' W	335	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:19 Rise	—	—	—	0	67	0.210	0.112
London, ON	42°59' N	081°14' W	251	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:24 Rise	—	—	—	0	68	0.123	0.051
Mississauga, ON	43°35' N	079°37' W	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:19 Rise	—	—	—	0	68	0.208	0.110
North York, ON	43°46' N	079°25' W	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:18 Rise	—	—	—	0	68	0.231	0.129
Saint Catharines..	43°10' N	079°15' W	110	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:17 Rise	—	—	—	0	68	0.249	0.144
Scarborough, ON	43°44' N	079°16' W	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:18 Rise	—	—	—	0	68	0.240	0.136
Toronto, ON	43°39' N	079°23' W	116	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	10:16 Rise	—	—	—	0	68	0.268	0.160
PRINCE EDWARD ISLAND Charlottetown, ..	46°14' N	063°08' W	55	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:35:19.8	187	230	4	72	0.897	0.874	
QUEBEC Laval, QC	45°35' N	073°45' W	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:51 Rise	—	—	—	0	67	0.702	0.629
Montréal, QC	45°31' N	073°34' W	57	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:49 Rise	—	—	—	0	67	0.735	0.670
Ottawa, QC	45°25' N	075°42' W	114	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:57 Rise	—	—	—	0	67	0.609	0.517
Québec, QC	46°49' N	071°14' W	73	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:38:05.3	187	228	0	67	0.832	0.791	

TABLE 11
LOCAL CIRCUMSTANCES FOR THE NORTH ATLANTIC
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.					
				U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °								
AZORES Ponta Delgada	37°44' N	025°40' W	36	08:32:25.7	297	352	18	—	—	—	—	—	—	—	—	—	—	—	09:36:29.5	9	64	31	94	0.682	0.606	
BERMUDA Hamilton	32°17' N	064°46' W	46	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	09:40 Rise	—	—	—	0	71	0.608	0.517
CANARY ISLANDS Las Palmas G.Ca...	28°07' N	015°28' W	6	08:44:06.8	320	26	28	—	—	—	—	—	—	—	—	—	—	—	09:42:07.1	10	76	41	94	0.348	0.235	
Santa Cruz Tene...	28°25' N	016°16' W	—	08:42:46.3	319	25	27	—	—	—	—	—	—	—	—	—	—	—	09:40:57.0	10	76	40	94	0.360	0.246	
FAEROE ISLANDS Tórshavn	62°01' N	006°46' W	—	09:12:16.8	264	289	32	—	—	—	—	—	—	—	—	—	—	—	10:18:34.0	193	212	38	138	0.701	0.630	
GREENLAND Godthåb	64°11' N	051°44' W	20	09:06:07.8	254	280	11	—	—	—	—	—	—	—	—	—	—	—	09:59:10.7	189	216	17	89	0.574	0.477	
ICELAND Reykjavík	64°09' N	021°51' W	28	09:08:34.7	260	286	24	—	—	—	—	—	—	—	—	—	—	—	10:09:38.4	191	215	30	120	0.646	0.563	
ST. PIERRE & MIQUELON Saint-Pierre Is...	46°47' N	056°11' W	—	08:39:26.4	272	313	0	—	—	—	—	—	—	—	—	—	—	—	09:35:08.7	187	231	9	77	0.928	0.913	

**TABLE 12
LOCAL CIRCUMSTANCES FOR AFRICA
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Location Name	Latitude	Longitude	Elev.	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbra. Durat.								
				h	m	s	h	m	s	h	m	s	h	m	s	h	m	s				h	m	s					
ALGERIA																													
Algier (Algiers)	36°47'N	003°03'E	59	09:03:12.1	1	306	358	47	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.634	0.549				
Annaba (Bône)	36°54'N	007°46'E	20	09:11:11.9	9	305	355	52	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.661	0.581		
Constantine	36°22'N	006°37'E	—	09:09:32.2	2	306	357	51	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.638	0.554		
Wahran (Oran)	35°43'N	000°43'W	—	08:57:49.0	307	3	43	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.588	0.495		
DJIBOUTI																													
Djibouti	11°36'N	043°09'E	7	11:22:39.3	333	241	58	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.384	0.270	
EGYPT																													
Alexandria	31°12'N	029°54'E	32	10:04:51.8	309	310	74	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.712	0.644	
AI-Jizah (Giza)	30°01'N	031°13'E	—	10:10:05.2	310	302	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.698	0.626	
AI-Mahallah al-...	30°58'N	031°10'E	—	10:08:02.6	309	303	74	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.722	0.656	
AI-Manshah	26°28'N	031°48'E	—	10:19:31.0	316	293	78	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.612	0.523	
AI-Qahirah (Cairo)	30°03'N	031°15'E	116	10:10:05.2	310	302	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.699	0.628	
As-Suways (Suez)	29°58'N	032°33'E	—	10:13:01.4	310	294	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.714	0.646	
Asyut	27°11'N	033°11'E	—	10:16:27.2	315	299	78	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.623	0.556	
Bahim	30°08'N	031°17'E	—	10:09:59.1	310	302	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.702	0.631	
Bur Sa'id (Port...)	31°16'N	032°18'E	—	10:09:52.4	308	297	74	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.744	0.684	
Shubra al-Khaym...	30°06'N	031°15'E	—	10:09:59.0	310	302	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.700	0.630	
Tanta	30°47'N	031°00'E	—	10:08:03.1	309	304	74	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.715	0.648	
ERITREA																													
Asmera	15°20'N	038°53'E	2325	11:05:21.8	331	244	67	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.421	0.309	
ETHIOPIA																													
Adis Abeba (Add...)	09°02'N	038°42'E	2450	11:29:20.9	345	246	60	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.247	0.143	
LIBYA																													
Banghazi	32°07'N	020°04'E	25	09:41:47.3	312	350	68	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.621	0.534	
Tarabulus (Trip...)	32°54'N	013°11'E	22	09:25:58.1	313	2	60	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.578	0.483	
MAURITANIA																													
Nouakchott	18°06'N	015°57'W	21	09:18:12.2	356	74	35	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.027	0.005	
MOROCCO																													
Casablanca (Dar...)	33°39'N	007°35'W	50	08:49:08.2	310	10	36	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.514	0.409	
Fes	34°05'N	004°57'W	—	08:52:31.0	310	9	39	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.529	0.427	
Marrakech	31°38'N	008°00'W	460	08:50:07.1	314	16	36	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.449	0.339	
Meknes	33°53'N	005°37'W	—	08:51:41.9	310	9	38	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.522	0.419	
Oujda	34°41'N	003°45'W	—	08:56:52.1	309	6	42	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.553	0.454	
Rabat	34°02'N	006°51'W	65	08:49:53.7	310	9	37	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.526	0.423	
Salé	34°04'N	006°50'W	—	08:49:53.9	310	8	37	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.527	0.424	
Tanger (Tangier)	35°48'N	005°45'W	73	08:50:32.1	306	3	38	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.582	0.488	
QATAR																													
Aq-Dawhah (Doha)	25°17'N	051°32'E	—	10:58:07.4	303	235	56	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.847	0.812
SOMALIA																													
Mogdisho (Mogad...)	02°04'N	045°22'E	12	12:00:43.1	353	250	44	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.152	0.070
SUDAN																													
AI-Khartum	15°36'N	032°32'E	390	10:56:14.8	337	250	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.331	0.219	
AI-Khartum Bahri	15°38'N	032°33'E	—	10:56:07.8	337	250	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.333	0.220	
Umm Durman (Omd...)	15°38'N	032°30'E	—	10:56:04.0	337	250	75	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.332	0.219	
TUNISIA																													
Sfax	34°44'N	010°46'E	—	09:18:44.9	309	359	56	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.614	0.526	
Tunis	36°48'N	010°11'E	66	09:15:40.3	305	353	54	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.673	0.596	
WESTERN SAHARA																													
El Aaiun	27°09'N	013°12'W	—	08:48:30.9	323	31	31	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	—	0.311	0.200	

**TABLE 16 - continued
LOCAL CIRCUMSTANCES FOR EUROPE - FRANCE II
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.				
				U.T.	P	Alt	U.T.	P	Alt	U.T.	P	Alt	U.T.	P	Alt	U.T.	P	Alt				U.T.	P	Alt	
FRANCE																									
Nancy	48°41'N	006°12'E	-	09:08:59.1	285	320	45	-	-	-	-	11:51:57.6	106	103	57	10:29:05.3	15	35	54	150	1.000	1.000	0.982		
Nantes	47°13'N	001°33'W	-	08:58:19.0	287	328	39	-	-	-	-	11:38:52.7	101	111	57	10:16:08.4	14	45	50	133	0.935	0.925	0.864		
Nice	43°42'N	007°15'E	-	09:08:36.1	293	334	48	-	-	-	-	11:57:05.3	99	91	61	10:31:24.2	16	39	58	149	0.861	0.831	0.820		
Nîmes	43°50'N	004°21'E	-	09:04:18.6	293	335	45	-	-	-	-	11:51:09.6	98	97	61	10:25:48.1	15	43	57	142	0.853	0.820	0.820		
Noyon	49°35'N	003°00'E	-	09:05:30.2	283	319	42	10:22:55.5	104	128	10:25:06.9	285	309	11:46:00.2	106	108	56	10:24:01.2	195	218	51	144	1.027	1.000	0.2ml1s
Orléans	47°55'N	001°54'E	-	09:02:56.1	286	324	42	-	-	-	-	11:44:54.0	103	107	57	10:21:55.1	15	41	52	141	0.963	0.961	0.961		
Paris	48°52'N	002°20'E	50	09:04:08.6	284	321	42	-	-	-	-	11:45:14.3	105	108	56	10:22:49.3	15	39	51	142	0.992	0.994	0.994		
Perpignan	43°18'N	000°22'W	-	08:57:40.9	293	339	41	-	-	-	-	11:41:31.1	95	105	62	10:16:56.0	14	49	54	132	0.822	0.781	0.781		
Pau	42°41'N	002°53'E	-	09:01:57.9	295	339	44	-	-	-	-	11:48:25.8	95	98	63	10:22:57.7	15	46	56	138	0.813	0.770	0.770		
Poitiers	46°35'N	000°20'E	-	09:00:08.3	288	329	41	-	-	-	-	11:42:28.7	101	108	58	10:19:00.3	14	44	52	136	0.921	0.907	0.907		
Reims	49°15'N	004°02'E	-	09:06:32.4	284	319	43	10:24:36.3	131	154	10:26:35.7	259	281	11:47:55.5	106	106	56	10:25:36.0	15	37	52	146	1.028	1.000	0.1m59s
Rennes	48°05'N	001°41'W	-	08:58:49.1	285	325	39	-	-	-	-	11:38:29.3	102	113	57	10:16:16.1	14	44	50	134	0.960	0.956	0.956		
Rethel	49°31'N	004°22'E	-	09:07:09.2	283	318	43	10:25:04.7	94	116	10:27:15.1	296	317	11:48:19.3	107	106	56	10:26:09.9	195	217	52	147	1.028	1.000	0.2ml10s
Rouen	49°27'N	001°05'E	-	09:03:06.0	283	320	40	10:20:11.0	144	170	10:21:50.5	244	270	11:42:51.0	105	111	56	10:21:00.7	14	40	50	140	1.027	1.000	0.1m40s
Saint Avoild	49°06'N	006°42'E	-	09:09:54.8	284	318	45	10:28:47.3	109	128	10:31:02.1	282	301	11:52:29.9	107	103	56	10:29:54.6	15	35	53	152	1.028	1.000	0.2ml15s
Saint Étienne	45°26'N	004°24'E	-	09:04:53.3	290	330	45	-	-	-	-	11:50:36.1	100	99	60	10:25:54.7	15	41	55	144	0.900	0.880	0.880		
Saint Étienne d.	49°23'N	001°06'E	-	09:03:03.9	284	321	40	10:20:20.3	155	181	10:21:42.0	233	259	11:42:54.4	105	110	56	10:21:01.1	14	40	50	140	1.027	1.000	0.1m22s
Saint Nazaire	47°17'N	003°12'W	-	08:57:37.4	287	328	39	-	-	-	-	11:37:41.2	101	112	57	10:15:08.2	14	45	50	132	0.936	0.926	0.926		
Saint Quentin	49°51'N	003°17'E	-	09:06:03.5	283	318	42	10:23:39.3	66	89	10:25:22.2	323	346	11:46:19.5	107	108	55	10:24:30.7	195	217	51	145	1.027	1.000	0.1m43s
Sarcelles	49°00'N	002°23'E	-	09:04:18.2	284	321	42	-	-	-	-	11:45:15.6	105	108	56	10:22:55.8	15	39	51	143	0.996	0.997	0.997		
Sarrebourg	48°44'N	007°03'E	-	09:10:09.9	285	319	45	10:29:41.1	150	169	10:31:18.8	242	260	11:53:21.7	106	101	56	10:30:29.9	16	34	54	152	1.028	1.000	0.1m38s
Saverne	48°44'N	007°22'E	-	09:10:35.9	285	319	45	10:30:07.3	143	161	10:31:55.9	249	267	11:53:53.7	107	101	56	10:31:01.6	16	34	54	153	1.028	1.000	0.1m49s
Sedan	49°42'N	004°57'E	-	09:08:01.7	283	318	43	10:26:19.0	61	82	10:27:54.3	329	350	11:49:10.2	107	106	56	10:27:06.6	195	216	52	148	1.028	1.000	0.1m35s
Senlis	49°12'N	002°35'E	-	09:04:42.0	284	320	42	10:23:44.0	164	189	10:23:50.6	225	249	11:45:30.1	105	108	56	10:23:17.3	15	39	51	143	1.027	1.000	0.1m07s
Soissons	49°22'N	003°20'E	-	09:05:44.9	284	319	42	10:23:28.7	126	149	10:23:31.9	264	287	11:46:40.8	106	107	56	10:24:30.3	15	38	52	145	1.028	1.000	0.2m03s
Strasbourg	48°35'N	007°45'E	142	09:11:02.5	285	319	46	10:30:58.5	158	176	10:32:22.7	234	252	11:54:39.5	106	100	56	10:31:40.6	16	34	54	153	1.028	1.000	0.1m44s
Thionville	49°22'N	006°10'E	-	09:09:22.5	284	318	44	10:27:59.3	86	106	10:30:05.8	305	324	11:51:24.8	107	104	56	10:29:02.6	195	215	53	151	1.028	1.000	0.2m07s
Toulon	43°07'N	005°56'E	-	09:06:31.2	294	336	47	-	-	-	-	11:54:41.7	98	92	62	10:28:54.6	16	42	58	145	0.838	0.802	0.802		
Toulouse	43°36'N	001°26'E	164	09:00:10.0	293	337	43	-	-	-	-	11:45:13.9	96	102	62	10:20:18.2	15	47	55	136	0.836	0.799	0.799		
Tours	47°23'N	000°41'E	-	09:01:04.8	287	326	41	-	-	-	-	11:42:54.2	102	108	58	10:19:48.1	14	43	52	138	0.945	0.937	0.937		
Troyes	48°18'N	004°05'E	-	09:05:57.3	285	322	43	-	-	-	-	11:48:33.6	105	105	57	10:25:33.9	15	38	53	145	0.981	0.982	0.982		
Valence	44°56'N	004°54'E	-	09:05:24.7	291	331	45	-	-	-	-	11:51:48.0	100	98	60	10:26:49.3	15	41	56	144	0.887	0.864	0.864		
Valenciennes	50°21'N	003°32'E	-	09:06:45.8	282	317	42	-	-	-	-	11:46:26.4	107	109	55	10:24:59.3	195	217	51	146	0.990	0.992	0.992		
Verdun	49°10'N	005°23'E	-	09:08:12.9	284	319	44	10:26:41.5	121	142	10:28:50.3	269	290	11:50:14.9	106	104	56	10:27:45.9	15	36	53	149	1.028	1.000	0.2m09s
Villeurbanne	45°46'N	004°53'E	-	09:05:41.7	290	329	45	-	-	-	-	11:51:22.0	101	99	60	10:26:47.4	15	40	55	145	0.911	0.895	0.895		

TABLE 18
LOCAL CIRCUMSTANCES FOR EUROPE - GREECE, GUERNSEY, HUNGARY, IRELAND & N. IRELAND
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact		Second Contact		Third Contact		Fourth Contact		Maximum Eclipse		Eclip. Mag.	Eclip. Obs.	Umbra Durat.
				h	m	h	m	h	m	h	m	h	m			
GREECE																
Athínai (Athens)	37°58' N	023°43' E	107	09:41:25.6	300 324 65	—	—	—	—	12:34:29.7	100 56 54	11:10:14.1	21 1 66 204	0.822	0.781	
Iraklion	35°20' N	025°09' E	—	09:47:39.2	304 325 68	—	—	—	—	12:40:18.9	98 47 54	11:16:50.5	21 353 67 214	0.766	0.711	
Kallithéa	37°57' N	023°42' E	—	09:41:24.7	300 324 65	—	—	—	—	11:10:13.3	100 56 55	11:10:13.3	21 1 66 204	0.821	0.780	
Lárisa	39°38' N	022°25' E	—	09:37:13.7	298 323 63	—	—	—	—	12:30:04.9	102 61 55	11:05:29.4	20 6 65 197	0.854	0.822	
Pátraí	38°15' N	021°44' E	—	09:37:04.2	300 328 63	—	—	—	—	12:30:34.8	100 58 56	11:05:40.0	20 7 66 197	0.810	0.766	
Peristérión	38°01' N	023°42' E	—	09:41:20.5	300 323 65	—	—	—	—	12:34:22.0	101 56 54	11:10:04.9	21 1 66 204	0.823	0.783	
Piraiévs (Piraeus...)	37°57' N	023°38' E	—	09:41:16.4	300 324 65	—	—	—	—	12:34:22.0	101 56 54	11:10:04.9	21 1 66 204	0.820	0.780	
Tessaloníki (S...)	40°38' N	022°56' E	24	09:37:29.4	296 319 62	—	—	—	—	12:29:38.0	104 64 54	11:05:23.2	20 6 64 198	0.886	0.863	
Volos	39°21' N	022°56' E	—	09:38:30.4	298 322 63	—	—	—	—	12:31:20.6	102 60 55	11:06:53.5	20 5 65 199	0.851	0.819	
GUERNSEY																
Saint Peter Port	49°27' N	002°32' W	—	08:59:03.3	283 322 38	—	—	—	—	11:36:45.3	104 115 55	10:15:34.6	14 43 48 134	0.997	0.999	0.1m41s
Alderney Island	49°43' N	002°12' W	—	08:59:39.3	283 321 38	—	—	10:15:20.6	140 169	10:17:01.8	247 275	10:16:11.2	14 42 48 135	1.027	1.000	
HUNGARY																
Ajka	47°07' N	017°34' E	—	09:25:38.5	286 312 53	10:48:07.8	98 101	10:50:27.7	297 299	12:12:13.9	109 86 54	10:49:17.8	198 200 58 177	1.029	1.000	0.2m20s
Baja	46°11' N	018°57' E	—	09:27:58.7	287 312 55	10:51:55.7	173 172	10:52:56.6	223 222	12:15:37.5	108 82 54	10:52:26.2	18 17 59 181	1.029	1.000	0.1m01s
Békéscsaba	46°41' N	021°06' E	—	09:31:41.4	286 309 54	—	—	—	—	12:18:09.2	110 82 53	10:55:52.0	198 193 58 187	0.998	0.999	
Budapest	47°30' N	019°05' E	120	09:28:12.9	285 309 54	—	—	—	—	12:14:00.8	110 86 53	10:51:41.6	198 197 58 181	0.991	0.993	
Debrecen	47°32' N	021°38' E	—	09:32:34.1	285 304 55	—	—	—	—	12:17:38.8	112 84 52	10:56:04.0	198 193 58 188	0.972	0.971	
Dombóvár	46°23' N	018°08' E	—	09:26:34.1	287 313 54	10:50:23.8	178 178	10:51:13.8	218 219	12:14:04.9	108 84 55	10:50:48.8	18 19 59 179	1.029	1.000	0.0m50s
Dunaújváros	46°58' N	018°57' E	—	09:27:58.1	286 310 54	10:50:51.8	75 74	10:52:50.2	322 320	12:14:33.2	110 85 54	10:51:51.0	198 197 58 181	1.029	1.000	0.1m58s
Győr	47°42' N	017°38' E	—	09:25:49.0	285 310 53	—	—	—	—	12:11:33.8	110 85 54	10:49:02.0	198 200 58 177	0.996	0.998	
Hódmezővásárhely	46°25' N	020°20' E	—	09:30:22.2	287 309 55	10:53:34.7	95 92	10:55:53.3	302 298	12:17:24.5	110 82 53	10:54:44.0	198 195 59 185	1.029	1.000	0.2m19s
Kalocsa	46°32' N	018°59' E	—	09:28:01.5	287 311 55	10:51:04.6	122 121	10:53:23.2	274 273	12:15:11.9	109 83 54	10:52:13.9	18 17 59 181	1.029	1.000	0.2m19s
Kaposvár	46°22' N	017°47' E	—	09:25:58.3	287 314 54	—	—	—	—	12:13:33.1	108 84 55	10:50:11.6	18 17 59 178	0.998	0.999	
Kecskemét	46°54' N	019°42' E	—	09:29:15.1	286 309 55	10:52:31.0	55 53	10:53:55.8	341 339	12:15:46.5	110 84 53	10:53:13.5	198 196 58 183	1.029	1.000	0.1m25s
Keszthely	46°46' N	017°15' E	—	09:25:05.1	287 313 53	10:48:03.0	148 150	10:49:52.4	248 250	12:12:11.5	108 85 55	10:48:57.7	18 20 59 177	1.029	1.000	0.1m49s
Kiskunfélegyháza	46°43' N	019°52' E	—	09:29:32.6	286 309 55	10:52:40.2	75 73	10:54:39.2	321 318	12:16:17.0	110 83 53	10:53:39.8	198 196 59 184	1.029	1.000	0.1m59s
Kiskunhalas	46°26' N	019°30' E	—	09:28:55.2	287 311 55	10:52:03.5	118 116	10:54:24.3	279 276	12:16:07.6	109 83 54	10:53:14.0	18 16 59 183	1.029	1.000	0.2m21s
Kőrmend	47°01' N	016°37' E	—	09:24:03.1	287 314 55	10:46:37.3	135 139	10:48:43.2	260 263	12:10:52.4	108 87 55	10:47:40.3	18 21 58 175	1.029	1.000	0.2m06s
Kézeg	47°23' N	016°33' E	—	09:23:59.3	286 313 52	10:46:10.6	96 100	10:48:29.2	299 302	12:10:18.3	109 88 55	10:47:19.9	198 201 58 175	1.029	1.000	0.2m19s
Makó	46°13' N	020°29' E	—	09:30:39.1	287 310 56	10:53:58.7	112 108	10:56:21.3	285 281	12:17:55.2	109 81 53	10:55:10.1	18 14 59 186	1.029	1.000	0.2m23s
Muskolc	48°06' N	020°47' E	—	09:31:07.1	284 305 54	—	—	—	—	12:15:36.1	112 86 52	10:54:09.3	198 194 57 185	0.964	0.961	
Nyíregyháza	47°59' N	021°43' E	—	09:32:42.2	284 303 55	—	—	—	—	12:17:04.6	112 85 52	10:55:49.6	198 193 57 188	0.960	0.956	
Paks	46°39' N	018°53' E	—	09:27:51.1	287 311 54	10:50:46.8	112 112	10:53:09.0	284 283	12:14:53.2	109 84 54	10:51:57.9	18 17 59 181	1.029	1.000	0.2m22s
Pécs	47°19' N	017°28' E	—	09:25:29.7	286 312 53	10:47:57.8	78 81	10:50:00.8	317 319	12:11:48.5	109 87 54	10:48:59.4	198 200 58 177	1.028	1.000	0.2m03s
Pécs	46°05' N	018°13' E	—	09:26:42.9	288 314 54	—	—	—	—	12:14:36.5	108 83 55	10:51:10.8	18 19 59 179	0.993	0.995	
Sárvár	47°15' N	016°57' E	—	09:24:37.8	286 313 53	10:46:56.9	100 104	10:49:17.4	295 298	12:11:06.0	109 87 55	10:48:07.2	198 201 58 176	1.028	1.000	0.2m20s
Siófok	46°54' N	018°04' E	—	09:26:28.1	286 312 54	10:49:08.7	108 109	10:51:30.9	288 288	12:13:17.5	109 85 54	10:50:19.8	18 19 58 179	1.028	1.000	0.2m22s
Sopron	47°41' N	016°36' E	—	09:24:07.0	286 312 54	10:46:31.1	56 59	10:47:57.9	339 343	12:10:00.1	109 88 54	10:47:14.5	198 201 58 175	1.028	1.000	0.1m27s
Szeged	46°15' N	020°09' E	—	09:30:03.8	287 310 55	10:53:22.2	118 115	10:55:43.0	279 275	12:17:22.2	109 82 54	10:54:32.6	18 15 59 185	1.029	1.000	0.2m21s
Székesfehérvár	47°12' N	018°25' E	—	09:27:04.5	286 311 54	10:49:55.8	61 62	10:51:33.1	334 334	12:13:25.5	110 86 54	10:50:44.5	198 198 58 180	1.028	1.000	0.1m37s
Székeszárd	46°21' N	018°42' E	—	09:27:52.5	287 312 55	10:51:01.9	154 154	10:52:41.2	242 241	12:15:00.6	109 83 54	10:51:51.6	18 18 59 181	1.029	1.000	0.1m39s
Szentest	46°39' N	020°16' E	—	09:30:14.4	286 309 55	10:53:30.2	69 66	10:55:20.7	327 324	12:16:58.5	110 83 53	10:54:25.5	198 195 59 185	1.029	1.000	0.1m50s
Szombathely	47°14' N	016°38' E	—	09:24:06.3	286 313 53	10:46:23.4	110 112	10:48:44.9	285 288	12:10:37.5	109 87 55	10:47:34.2	18 21 58 175	1.028	1.000	0.2m22s
Vápalota	47°12' N	018°09' E	—	09:26:37.6	286 311 54	10:49:19.5	71 74	10:51:13.0	325 325	12:13:01.1	109 86 54	10:50:16.3	198 199 58 179	1.028	1.000	0.1m54s
Veszprém	47°06' N	017°55' E	—	09:26:13.6	286 312 53	10:48:48.0	90 92	10:51:03.3	305 306	12:12:47.7	109 86 54	10:49:55.6	198 199 58 178	1.029	1.000	0.2m15s
Zalaegerszeg	46°51' N	016°51' E	—	09:24:25.4	287 314 53	10:47:18.7	150 153	10:49:04.3	246 248	12:11:27.2	108 86 55	10:48:11.5	18 21 58 176	1.029	1.000	0.1m46s
IRELAND																
Cork	51°54' N	008°28' W	17	08:56:03.2	279 316 33	—	—	—	—	11:27:06.8	106 124 51	10:09:06.5	193 224 43 127	0.966	0.963	
Dublin (Baile Á..)	53°20' N	006°15' W	47	08:59:42.8	277 312 34	—	—	—	—	11:30:16.9	109 123 50	10:12:49.4	193 221 43 131	0.925	0.912	
IRELAND, NORTH																
Belfast, NI	54°35' N	005°55' W	17	09:01:57.0	275 309 34	—	—	—	—	11:30:31.8	111 124 49	10:14:03.7	193 219 43 133	0.891	0.868	

TABLE 19
LOCAL CIRCUMSTANCES FOR EUROPE - ITALY
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.	
				U.T.	P	V	U.T.	P	V	U.T.	P	V	U.T.	P	V	U.T.	P	V				
ITALY																						
Ancona	43°38'N	013°30'E	—	09:18:45.1	293	327	53	—	—	—	—	12:09:17.5	102	81	59	10:43:46.3	18	28	61	166	0.894	0.873
Bari	41°07'N	016°52'E	—	09:25:28.8	296	320	57	—	—	—	—	12:18:11.9	100	70	59	10:52:23.5	19	21	64	177	0.847	0.901
Bergamo	45°41'N	009°43'E	—	09:12:45.9	290	325	49	—	—	—	—	12:00:27.5	103	92	59	10:35:39.5	16	33	58	157	0.930	0.919
Bologna	44°29'N	011°20'E	—	09:15:08.1	291	327	51	—	—	—	—	12:04:27.6	102	86	59	10:39:07.3	17	32	60	160	0.905	0.887
Bolzano	46°31'N	011°22'E	—	09:15:31.4	298	321	50	—	—	—	—	12:02:40.6	105	91	58	10:38:26.3	17	30	58	161	0.963	0.960
Brescia	45°33'N	013°15'E	—	09:18:21.0	289	322	52	—	—	—	—	12:06:55.8	105	87	58	10:42:18.5	17	28	59	166	0.946	0.939
Cagliari	39°13'N	009°07'E	1	09:12:17.1	301	346	52	—	—	—	—	12:03:19.4	93	77	65	10:36:40.4	17	40	64	151	0.738	0.677
Catania	37°30'N	015°06'E	—	09:24:28.4	303	344	59	—	—	—	—	12:07:51.6	94	61	63	10:51:31.8	19	26	68	172	0.729	0.666
Catanzaro	38°54'N	016°26'E	—	09:25:58.1	300	337	59	—	—	—	—	12:19:32.2	97	64	61	10:53:21.0	19	22	66	176	0.781	0.730
Como	45°47'N	009°05'E	—	09:11:49.4	289	326	48	—	—	—	—	11:59:13.0	103	93	59	10:34:27.6	16	34	57	155	0.930	0.919
Cosenza	39°17'N	016°15'E	—	09:25:21.0	300	336	58	—	—	—	—	12:18:48.4	97	65	61	10:52:36.5	19	23	66	175	0.790	0.742
Ferrara	44°50'N	011°35'E	—	09:15:34.1	291	326	51	—	—	—	—	12:04:37.0	103	87	59	10:39:27.8	17	31	59	161	0.916	0.901
Firenze (Floren...)	43°46'N	011°15'E	—	09:14:58.3	293	330	51	—	—	—	—	12:04:54.6	101	85	60	10:39:15.0	17	32	60	160	0.884	0.860
Foggia	41°27'N	015°34'E	—	09:22:55.8	296	331	56	—	—	—	—	12:15:21.7	100	73	60	10:49:24.0	18	24	64	173	0.847	0.813
Forlì	44°13'N	012°03'E	—	09:16:17.8	292	327	51	—	—	—	—	12:06:02.2	102	85	59	10:40:37.6	17	30	60	162	0.902	0.883
Genova (Genoa)	44°25'N	008°57'E	97	09:11:19.9	292	330	49	—	—	—	—	11:59:58.8	101	90	60	10:34:32.5	16	36	59	154	0.890	0.868
La Spezia	44°07'N	009°50'E	—	09:12:41.5	292	330	50	—	—	—	—	12:01:54.3	101	88	60	10:36:20.4	17	34	59	156	0.886	0.863
Latina	41°28'N	012°52'E	—	09:18:03.7	297	335	54	—	—	—	—	12:10:00.1	99	76	61	10:43:42.2	18	30	63	164	0.828	0.789
Lecce	40°23'N	018°11'E	—	09:28:22.2	297	330	59	—	—	—	—	12:21:29.4	100	67	58	10:55:49.0	19	18	65	182	0.837	0.801
Livorno (Leghor...)	43°33'N	010°19'E	—	09:13:26.5	293	331	50	—	—	—	—	12:03:16.3	100	86	61	10:37:23.2	17	34	60	157	0.873	0.846
Massa	44°01'N	013°09'E	—	09:13:11.5	292	330	50	—	—	—	—	12:02:35.5	101	87	60	10:36:59.5	17	34	60	157	0.885	0.862
Messina	38°11'N	015°34'E	—	09:24:49.4	302	341	59	—	—	—	—	12:18:20.4	95	63	62	10:52:01.1	19	25	67	173	0.753	0.695
Mestre	45°29'N	012°15'E	—	09:16:43.2	290	323	51	—	—	—	—	12:05:13.4	104	88	58	10:40:27.7	17	29	59	163	0.939	0.908
Milano (Milan)	45°28'N	009°12'E	—	09:11:55.3	290	327	48	—	—	—	—	11:59:40.6	103	92	59	10:34:45.1	16	34	58	155	0.922	0.900
Modena	44°40'N	010°55'E	—	09:14:28.7	291	327	50	—	—	—	—	12:03:31.4	102	87	59	10:38:14.9	17	32	59	159	0.908	0.891
Monza	45°35'N	009°16'E	—	09:12:03.1	290	326	50	—	—	—	—	11:59:42.6	103	92	59	10:34:50.8	16	34	58	155	0.925	0.913
Napoli (Naples)	40°51'N	014°17'E	25	09:14:09.5	297	335	56	—	—	—	—	12:13:23.5	99	72	61	10:47:06.6	16	37	64	168	0.820	0.779
Novara	45°28'N	008°38'E	—	09:11:03.5	290	327	48	—	—	—	—	11:58:37.3	102	93	59	10:33:41.7	16	35	58	154	0.919	0.905
Padova	45°25'N	011°53'E	—	09:16:07.2	290	324	51	—	—	—	—	12:04:37.7	104	88	58	10:39:48.0	17	30	59	162	0.935	0.924
Palermo	38°07'N	013°22'E	108	09:20:39.2	302	344	57	—	—	—	—	12:13:34.2	94	66	63	10:47:00.9	18	31	67	165	0.734	0.671
Parma	44°48'N	010°20'E	—	09:13:33.8	291	327	50	—	—	—	—	12:02:19.2	102	89	59	10:37:05.0	17	33	59	158	0.908	0.891
Perugia	43°08'N	012°22'E	—	09:16:52.1	294	330	52	—	—	—	—	12:07:35.9	101	81	60	10:41:46.0	17	30	61	163	0.873	0.846
Pescara	42°28'N	014°13'E	—	09:20:11.0	295	330	54	—	—	—	—	12:11:46.2	101	77	60	10:45:54.2	18	27	62	168	0.866	0.837
Piacenza	45°01'N	009°40'E	—	09:12:32.9	291	327	49	—	—	—	—	12:00:53.8	102	90	59	10:35:44.9	17	34	58	156	0.911	0.895
Pisa	43°43'N	010°23'E	—	09:13:33.2	293	331	50	—	—	—	—	12:03:16.5	101	86	60	10:37:33.4	17	34	60	157	0.878	0.852
Prato	43°53'N	011°06'E	—	09:14:43.6	292	329	51	—	—	—	—	12:04:31.5	101	85	60	10:38:54.4	17	32	60	159	0.887	0.863
Ravenna	44°25'N	012°12'E	—	09:16:32.9	291	327	51	—	—	—	—	12:06:08.1	103	85	59	10:40:49.7	17	30	60	163	0.908	0.891
Reggio di Calab.	38°06'N	015°39'E	—	09:25:03.1	302	341	59	—	—	—	—	12:18:35.2	95	62	62	10:52:16.9	19	24	67	174	0.751	0.693
Reggio nell'Em...	44°43'N	010°36'E	—	09:13:58.6	291	327	50	—	—	—	—	12:03:53.3	102	88	59	10:37:37.4	17	33	59	158	0.908	0.890
Rimini	44°04'N	012°34'E	—	09:17:09.2	292	327	52	—	—	—	—	12:07:08.4	102	84	59	10:41:42.4	17	30	60	164	0.900	0.881
Roma (Rome)	41°54'N	012°29'E	115	09:17:16.4	296	334	53	—	—	—	—	12:08:52.4	99	78	61	10:42:39.1	18	31	63	163	0.838	0.802
Salerno	40°41'N	014°47'E	—	09:21:48.8	298	334	56	—	—	—	—	12:14:32.9	99	71	61	10:48:18.0	18	26	64	170	0.819	0.778
Sassari	40°41'N	008°34'E	—	09:10:51.0	298	342	50	—	—	—	—	12:01:24.7	95	80	60	10:34:53.0	17	40	62	150	0.780	0.729
Savona	44°17'N	008°30'E	—	09:10:36.8	292	331	49	—	—	—	—	11:59:12.1	101	90	60	10:33:42.5	16	36	59	153	0.884	0.860
Siracusa (Syrac...)	37°04'N	015°18'E	—	09:25:15.0	304	345	59	—	—	—	—	12:18:37.4	94	59	63	10:52:22.6	19	25	68	172	0.719	0.652
Siracusa	37°04'N	015°18'E	—	09:25:15.0	304	345	59	—	—	—	—	12:18:37.4	94	59	63	10:52:22.6	19	25	68	172	0.719	0.652
Taranto	40°28'N	017°15'E	—	09:26:32.3	298	331	58	—	—	—	—	12:19:37.1	100	68	59	10:53:45.4	19	20	65	179	0.832	0.794
Terni	42°34'N	012°37'E	—	09:17:22.3	295	332	53	—	—	—	—	12:08:34.6	100	80	60	10:42:34.1	18	30	62	163	0.858	0.827
Torino (Turin)	45°03'N	007°40'E	—	09:09:29.6	291	329	47	—	—	—	—	11:57:05.2	101	93	60	10:31:58.0	16	37	57	151	0.902	0.883
Torre del Greco	40°47'N	014°22'E	—	09:21:00.3	297	335	56	—	—	—	—	12:13:37.2	98	72	67	10:47:20.1	18	27	64	169	0.819	0.778
Trento	46°04'N	011°08'E	—	09:15:03.1	289	323	50	—	—	—	—	12:02:41.0	104	90	58	10:38:09.8	17	31	58	160	0.949	0.942
Trieste (Triest)	45°46'N	013°46'E	—	09:19:12.5	289	331	52	—	—	—	—	12:07:42.7										

TABLE 20
LOCAL CIRCUMSTANCES FOR EUROPE -
LATVIA, LIECHTENSTEIN, LITHUANIA, MOLDOVA, MONACO & NETHERLANDS
LATVIA, LIECHTENSTEIN, LITHUANIA, LUXEMBOURG, MACEDONIA, MOLDOVA, MONACO & NETHERLANDS
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.
				h	m	s	h	m	s	h	m	s	h	m	s	h	m	s			
LATVIA																					
Daugavpils	55°53' N	026°32' E	—	09:41:45.2	271	279	49	—	—	—	12:09:18.7	124	102	44	10:56:23.4	197	189	49	194	0.737	0.674
Liepaja	56°31' N	021°01' E	—	09:34:15.4	271	285	47	—	—	—	12:02:45.4	122	106	46	10:48:52.9	197	195	49	183	0.754	0.696
Riga	56°57' N	024°06' E	—	09:38:43.5	270	280	47	—	—	—	12:05:06.8	124	106	45	10:52:31.3	197	192	48	189	0.727	0.662
LIECHTENSTEIN																					
Vaduz	47°09' N	009°31' E	—	09:12:54.9	287	321	48	—	—	—	11:58:52.5	105	95	58	10:34:56.4	16	32	56	157	0.971	0.970
LITHUANIA																					
Kaunas	54°54' N	023°54' E	—	09:37:34.1	273	284	49	—	—	—	12:08:26.1	122	101	46	10:53:42.5	197	192	50	189	0.775	0.722
Klaipeda (Memel)	55°43' N	021°07' E	—	09:33:57.1	272	286	48	—	—	—	12:04:08.7	121	104	47	10:49:27.8	197	195	50	183	0.772	0.719
Panevezys	55°44' N	024°21' E	—	09:38:32.7	271	282	49	—	—	—	12:07:27.6	123	103	45	10:53:41.8	197	192	49	190	0.753	0.695
Stauliai	55°56' N	023°19' E	—	09:37:09.3	271	283	48	—	—	—	12:06:04.7	122	104	46	10:52:12.8	197	193	49	187	0.755	0.697
Vilnius	54°41' N	025°19' E	—	09:39:35.3	273	283	50	—	—	—	12:10:15.8	122	100	46	10:55:46.3	197	190	50	192	0.772	0.718
LUXEMBOURG																					
Differdange	49°32' N	005°52' E	—	09:09:05.6	283	317	44	10:27:39.9	69	89	10:29:27.8	321	341	—	10:28:33.8	195	215	53	150	1.028	1.000
Dudelange	49°28' N	006°05' E	—	09:09:19.9	283	317	44	10:27:57.1	74	94	10:29:52.2	316	335	—	10:28:54.6	195	215	53	150	1.028	1.000
Esch-Sur-Alzette	49°30' N	005°59' E	—	09:09:13.4	283	317	44	10:27:49.3	72	92	10:29:40.8	319	338	—	10:28:45.0	195	215	53	150	1.028	1.000
Luxembourg	49°36' N	006°09' E	334	09:09:30.4	283	317	44	10:28:21.2	52	72	10:29:41.1	339	358	—	10:29:01.1	195	215	53	151	1.028	1.000
MACEDONIA																					
Bitola	41°01' N	021°20' E	—	09:34:04.9	296	322	61	—	—	—	12:26:25.0	103	66	55	11:01:43.8	20	12	63	190	0.882	0.857
Gostivar	41°47' N	020°54' E	—	09:32:47.8	294	321	60	—	—	—	12:24:40.6	104	69	55	11:00:03.7	19	12	63	190	0.899	0.874
Kumanovo	42°08' N	021°43' E	—	09:34:10.7	294	318	60	—	—	—	12:25:34.6	105	69	55	11:01:19.3	19	10	63	190	0.915	0.900
Skopje	41°59' N	021°26' E	240	09:33:42.7	294	319	60	—	—	—	12:25:18.6	104	69	55	11:00:55.1	19	11	63	191	0.909	0.892
Tetovo	42°01' N	020°58' E	—	09:32:48.1	294	320	60	—	—	—	12:24:29.0	104	69	55	11:00:59.5	19	12	63	190	0.905	0.888
MOLDOVA																					
Bel'cy	47°46' N	029°56' E	—	09:43:37.4	283	292	57	—	—	—	12:25:23.6	115	82	48	11:06:15.8	199	183	56	203	0.916	0.901
Bendery	46°48' N	029°29' E	—	09:46:44.0	284	290	58	—	—	—	12:28:55.4	115	79	47	11:09:52.3	199	180	56	208	0.926	0.914
Kisin'ov (Kishi...)	47°00' N	028°50' E	—	09:45:28.4	283	291	58	—	—	—	12:27:48.6	115	80	47	11:08:35.3	199	181	56	206	0.927	0.915
Tiraspol'	46°51' N	029°38' E	—	09:46:59.3	283	290	58	—	—	—	12:29:00.3	115	79	47	11:10:03.0	200	180	56	208	0.924	0.911
MONACO																					
Monaco	43°42' N	007°23' E	55	09:08:48.3	293	333	48	—	—	—	11:57:21.4	99	91	61	10:31:39.8	16	39	59	149	0.862	0.832
NETHERLANDS																					
Alkmaar	52°37' N	004°44' E	—	09:10:11.2	278	309	42	—	—	—	11:46:46.3	111	111	53	10:27:09.5	195	213	50	150	0.924	0.911
Amersfoort	52°09' N	005°24' E	—	09:10:33.6	279	310	42	—	—	—	11:48:07.0	111	109	53	10:28:03.6	195	213	50	151	0.935	0.925
Amsterdam	52°22' N	004°54' E	2	09:10:09.2	279	310	42	—	—	—	11:47:12.3	111	110	53	10:27:21.6	195	213	50	150	0.931	0.919
Apeldoorn	52°13' N	005°58' E	—	09:11:18.6	279	310	43	—	—	—	11:48:54.7	111	109	53	10:28:54.0	195	213	50	152	0.931	0.920
Arnhem	51°59' N	005°55' E	—	09:11:02.8	279	310	43	—	—	—	11:49:01.4	111	109	53	10:28:48.2	195	213	51	152	0.938	0.928
Breda	51°35' N	004°46' E	—	09:09:17.9	280	312	42	—	—	—	11:47:34.5	110	109	54	10:27:03.0	195	214	50	149	0.952	0.947
Dordrecht	51°49' N	004°40' E	—	09:09:22.8	280	312	42	—	—	—	11:47:15.3	110	110	54	10:26:56.2	195	214	50	149	0.946	0.939
Eindhoven	51°26' N	005°28' E	—	09:10:01.7	280	312	43	—	—	—	11:48:45.8	110	108	54	10:28:04.8	195	214	51	150	0.954	0.949
Enschede	52°12' N	006°53' E	—	09:12:25.7	279	309	43	—	—	—	11:50:17.4	111	108	53	10:30:15.0	195	212	51	154	0.929	0.917
Geele	50°58' N	005°52' E	—	09:10:08.9	281	313	43	—	—	—	11:49:44.2	109	107	54	10:28:38.7	195	214	51	151	0.966	0.963
Groningen	53°13' N	006°33' E	—	09:12:54.7	278	307	43	—	—	—	11:48:56.0	113	110	52	10:29:49.9	195	211	50	154	0.902	0.883
Haarlem	52°23' N	004°38' E	—	09:09:51.1	279	310	42	—	—	—	11:46:47.6	111	111	53	10:26:58.6	195	214	50	149	0.931	0.920
Haarlemmermeer	52°15' N	004°38' E	—	09:09:43.8	279	311	42	—	—	—	11:46:53.5	111	110	53	10:26:57.3	195	214	50	149	0.935	0.924
Heerlen	50°54' N	005°59' E	—	09:10:14.7	281	313	43	—	—	—	11:49:58.2	109	107	54	10:28:49.2	195	214	52	151	0.967	0.965
Leiden	52°09' N	004°30' E	—	09:09:28.8	279	311	43	—	—	—	11:46:45.8	110	110	53	10:26:44.7	195	214	50	149	0.938	0.928
Maastricht	50°52' N	005°43' E	—	09:09:52.9	281	314	43	—	—	—	11:49:34.6	109	107	54	10:28:24.5	195	214	51	151	0.969	0.968
Nijmegen	51°50' N	005°50' E	—	09:10:49.0	280	311	43	—	—	—	11:49:00.9	110	108	53	10:28:40.0	195	213	51	151	0.942	0.934
Rottterdam	51°55' N	004°28' E	—	09:09:11.3	280	312	42	—	—	—	11:46:52.7	110	110	53	10:26:39.5	195	214	50	149	0.944	0.936
's-Gravenhage (...)	52°06' N	004°18' E	—	09:09:11.8	279	311	42	—	—	—	11:46:29.8	110	111	53	10:26:26.7	195	214	50	149	0.940	0.931
's-Hertogenbosch	51°41' N	005°19' E	—	09:10:03.2	280	312	43	—	—	—	11:48:20.7	110	109	54	10:27:52.9	195	214	51	150	0.948	0.941
Tilburg	51°34' N	005°05' E	—	09:09:40.1	280	312	42	—	—	—	11:48:04.4	110	109	54	10:27:31.2	195	214	51	150	0.952	0.946
Utrecht	52°05' N	005°08' E	—	09:10:10.7	279	311	42	—	—	—	11:47:46.0	110	110	53	10:27:39.6	195	214	50	149	0.938	0.928
Zaandam	52°26' N	004°49' E	—	09:10:06.9	279	310	42	—	—	—	11:47:01.8	111	111	53	10:27:15.0	195	214	50	150	0.929	0.917
Zoetermeer	52°03' N	004°30' E	—	09:09:23.4	279	311	42	—	—	—	11:46:50.1	110	110	53	10:26:43.7	195	214	50	149	0.941	0.932

TABLE 21
LOCAL CIRCUMSTANCES FOR EUROPE - NORWAY, POLAND, PORTUGAL
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.	
				U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °	U.T. h m s	P °	V °				
NORWAY																						
Bergen	60°23'N	005°20'E	43	09:19:41.7	267	289	38	—	—	—	—	—	—	—	—	—	—	—	—	—	0.722	0.656
Oslo	59°55'N	010°45'E	94	09:24:38.1	267	286	41	—	—	—	—	—	—	—	—	—	—	—	—	—	0.718	0.651
Stavanger	58°58'N	005°45'E	—	09:18:17.2	269	292	39	—	—	—	—	—	—	—	—	—	—	—	—	—	0.756	0.698
Trondheim	63°25'N	010°25'E	127	09:28:29.4	262	278	38	—	—	—	—	—	—	—	—	—	—	—	—	—	0.637	0.553
POLAND																						
Białystok	53°09'N	023°09'E	—	09:35:51.1	276	289	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.822	0.781
Bielsko-Biala	49°49'N	019°02'E	—	09:28:29.2	282	303	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.931	0.920
Bydgoszcz	53°08'N	018°00'E	—	09:28:09.6	277	296	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.853	0.820
Bytom (Beuthen)	50°22'N	018°54'E	—	09:28:25.3	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.918	0.903
Chorzow	50°19'N	018°07'E	—	09:28:29.1	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.919	0.904
Czestochowa	50°49'N	019°06'E	—	09:28:52.1	280	300	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.905	0.887
Elblag (Elbing)	54°10'N	019°25'E	—	09:30:43.9	275	292	48	—	—	—	—	—	—	—	—	—	—	—	—	—	0.819	0.778
Gdansk (Danzig)	54°23'N	018°40'E	11	09:29:46.7	275	292	48	—	—	—	—	—	—	—	—	—	—	—	—	—	0.818	0.776
Gdynia	54°32'N	018°33'E	—	09:29:41.9	274	292	48	—	—	—	—	—	—	—	—	—	—	—	—	—	0.815	0.772
Gliwice (Gleiwitz...)	50°17'N	018°40'E	—	09:28:01.8	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.921	0.908
Gorzow Wielkopolo...	52°44'N	015°15'E	—	09:23:59.7	278	300	48	—	—	—	—	—	—	—	—	—	—	—	—	—	0.878	0.852
Grudziadz	53°29'N	018°45'E	—	09:29:25.5	276	294	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.840	0.804
Kalisz	51°46'N	018°06'E	—	09:27:40.8	279	299	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.887	0.864
Katowice	50°16'N	019°00'E	—	09:28:33.0	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.920	0.906
Kielce	50°52'N	020°37'E	—	09:31:16.0	280	298	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.894	0.873
Koszalin (Köslin...)	54°12'N	016°09'E	—	09:26:09.1	275	296	47	—	—	—	—	—	—	—	—	—	—	—	—	—	0.836	0.799
Krakow	50°03'N	019°58'E	220	09:30:02.0	281	301	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.919	0.905
Legnica (Liegnitz...)	51°13'N	016°09'E	—	09:24:31.1	280	303	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.912	0.896
Lodz	51°46'N	019°30'E	—	09:29:48.3	278	297	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.878	0.853
Lublin	51°15'N	022°35'E	—	09:34:28.9	279	294	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.872	0.844
Olseztyn (Alliens...)	53°48'N	020°29'E	—	09:32:06.2	275	291	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.822	0.781
Opole (Oppeln)	50°41'N	017°55'E	—	09:26:59.3	280	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.916	0.900
Plock	52°33'N	019°43'E	—	09:30:25.9	277	295	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.858	0.826
Poznan	52°25'N	016°55'E	—	09:26:13.3	278	299	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.877	0.851
Radom	51°25'N	021°10'E	—	09:32:16.7	279	296	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.877	0.851
Ruda Slaska	50°18'N	018°51'E	—	09:28:19.4	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.920	0.906
Rybnik	50°06'N	018°32'E	—	09:27:46.1	281	303	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.927	0.915
Rzeszow	50°03'N	022°00'E	—	09:33:19.8	281	298	53	—	—	—	—	—	—	—	—	—	—	—	—	—	0.905	0.887
Slupsk (Stolp)	54°28'N	017°01'E	—	09:27:31.1	275	294	47	—	—	—	—	—	—	—	—	—	—	—	—	—	0.825	0.785
Sosnowiec	50°18'N	019°08'E	—	09:28:46.2	281	302	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.918	0.903
Szczecin (Stettin...)	53°24'N	014°32'E	—	09:23:24.3	277	299	47	—	—	—	—	—	—	—	—	—	—	—	—	—	0.864	0.835
Tarnow	50°01'N	021°00'E	—	09:31:41.7	281	299	53	—	—	—	—	—	—	—	—	—	—	—	—	—	0.913	0.897
Torun	53°02'N	019°35'E	—	09:28:57.7	277	296	49	—	—	—	—	—	—	—	—	—	—	—	—	—	0.852	0.819
Tychy	50°09'N	018°59'E	—	09:28:29.6	281	302	52	—	—	—	—	—	—	—	—	—	—	—	—	—	0.923	0.909
Walbrzych (Wald...)	50°46'N	016°17'E	—	09:24:31.5	281	304	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.923	0.910
Warszawa (Warsaw...)	52°15'N	021°00'E	90	09:32:16.3	277	294	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.857	0.826
Wloclawek	52°39'N	019°02'E	—	09:29:27.1	277	296	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.859	0.828
Wodzislaw Slaski	50°00'N	018°28'E	—	09:27:38.1	281	303	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.930	0.919
Wroclaw (Bresla...)	51°06'N	017°00'E	147	09:25:44.6	280	302	50	—	—	—	—	—	—	—	—	—	—	—	—	—	0.910	0.894
Zabrze	50°18'N	018°46'E	—	09:28:11.5	281	302	51	—	—	—	—	—	—	—	—	—	—	—	—	—	0.920	0.906
Zielona Gora (G...)	51°56'N	015°31'E	—	09:23:56.1	279	302	48	—	—	—	—	—	—	—	—	—	—	—	—	—	0.897	0.877
PORTUGAL																						
Lisboa (Lisbon)	38°43'N	009°08'W	95	08:46:01.9	300	354	34	—	—	—	—	—	—	—	—	—	—	—	—	—	0.672	0.595
Porto	41°11'N	008°36'W	—	08:47:10.8	296	347	34	—	—	—	—	—	—	—	—	—	—	—	—	—	0.748	0.688

TABLE 23
LOCAL CIRCUMSTANCES FOR EUROPE - SCOTLAND & SPAIN
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclips. Obs.	Umbral Durat.			
				U.T.	P.	V.	U.T.	P.	V.	U.T.	P.	V.	U.T.	P.	V.	U.T.	P.	V.						
SCOTLAND																								
Aberdeen	57°10' N	002°04' W	—	09:08:27.9	271	300	361	—	—	—	—	—	—	—	—	—	—	10:20:04.6	193	214	43	141	0.818	0.776
Ayr	55°28' N	003°38' W	—	09:03:53.8	274	306	35	—	—	—	—	—	—	—	—	—	—	10:16:08.1	193	218	43	136	0.866	0.837
Dundee	56°28' N	003°00' W	—	09:06:40.8	272	303	35	—	—	—	—	—	—	—	—	—	—	10:18:37.3	193	216	43	139	0.838	0.801
Dunfermline	56°04' N	003°29' W	—	09:05:42.6	273	304	35	—	—	—	—	—	—	—	—	—	—	10:17:50.0	193	216	43	138	0.849	0.815
Edinburgh	55°57' N	003°13' W	134	09:05:47.3	273	304	35	—	—	—	—	—	—	—	—	—	—	10:18:04.4	193	216	43	138	0.852	0.818
Falkirk	56°00' N	003°48' W	—	09:05:20.6	273	304	35	—	—	—	—	—	—	—	—	—	—	10:17:25.7	193	216	43	138	0.851	0.818
Glasgow	55°53' N	004°15' W	—	09:04:47.6	273	305	35	—	—	—	—	—	—	—	—	—	—	10:16:50.3	193	217	43	137	0.855	0.822
Greenock	55°57' N	004°45' W	—	09:04:27.3	273	305	35	—	—	—	—	—	—	—	—	—	—	10:16:18.5	193	217	43	136	0.853	0.821
Kirkcaldy	56°07' N	003°10' W	—	09:06:03.5	273	304	35	—	—	—	—	—	—	—	—	—	—	10:18:13.7	193	216	43	139	0.847	0.813
SPAIN																								
Albacete	38°59' N	001°51' W	—	08:55:07.6	301	353	41	—	—	—	—	—	—	—	—	—	—	10:13:03.7	14	56	55	124	0.687	0.614
Alcalá de Henar...	40°29' N	003°22' W	—	08:53:08.7	298	348	39	—	—	—	—	—	—	—	—	—	—	10:10:32.7	14	55	53	123	0.731	0.667
Alcorcón	40°21' N	003°50' W	—	08:52:32.0	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:37.4	13	56	52	122	0.726	0.661
Alicante	38°21' N	000°29' W	—	08:57:09.6	302	354	42	—	—	—	—	—	—	—	—	—	—	10:15:45.0	14	56	56	126	0.671	0.594
Almería	36°50' N	002°27' W	—	08:54:44.5	305	359	41	—	—	—	—	—	—	—	—	—	—	10:11:20.7	14	60	55	120	0.619	0.532
Avilés	43°33' N	005°55' W	—	08:51:07.0	292	339	36	—	—	—	—	—	—	—	—	—	—	10:07:24.1	13	53	49	122	0.821	0.779
Badajoz	38°53' N	006°58' W	—	08:48:32.5	300	353	36	—	—	—	—	—	—	—	—	—	—	10:03:13.9	13	60	50	115	0.678	0.602
Badalona	41°27' N	002°15' E	—	09:00:54.5	297	343	44	—	—	—	—	—	—	—	—	—	—	10:21:39.1	15	49	57	135	0.774	0.721
Barcelona	41°23' N	002°11' E	95	09:00:48.3	297	344	44	—	—	—	—	—	—	—	—	—	—	10:21:30.8	15	49	57	135	0.772	0.718
Bilbao	43°15' N	002°58' W	—	08:54:23.2	293	340	39	—	—	—	—	—	—	—	—	—	—	10:12:16.1	14	51	51	127	0.815	0.773
Burgos	42°21' N	003°42' W	—	08:53:10.2	295	343	38	—	—	—	—	—	—	—	—	—	—	10:10:37.7	13	53	51	124	0.787	0.737
Cádiz	36°32' N	006°18' W	—	08:49:35.2	305	357	37	—	—	—	—	—	—	—	—	—	—	10:03:25.9	13	63	51	113	0.605	0.515
Cartagena	37°36' N	000°59' W	—	08:56:37.1	304	357	42	—	—	—	—	—	—	—	—	—	—	10:14:35.3	14	53	56	128	0.646	0.564
Castellón de la...	39°59' N	000°02' W	—	08:57:36.6	299	349	42	—	—	—	—	—	—	—	—	—	—	10:16:54.9	14	53	56	128	0.723	0.657
Córdoba	38°51' N	004°46' W	—	08:51:19.4	303	356	38	—	—	—	—	—	—	—	—	—	—	10:06:56.7	13	60	52	117	0.649	0.589
Elche	38°15' N	000°42' W	—	08:56:52.0	302	355	42	—	—	—	—	—	—	—	—	—	—	10:15:16.9	14	56	56	125	0.667	0.589
El Ferrol del C...	43°29' N	008°14' W	—	08:48:37.3	292	340	34	—	—	—	—	—	—	—	—	—	—	10:03:39.5	12	54	47	118	0.817	0.775
Fuenlabrada	40°17' N	003°48' W	—	08:52:34.0	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:39.7	13	56	52	122	0.724	0.659
Getafe	40°18' N	003°43' W	—	08:52:40.5	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:49.4	13	56	52	122	0.725	0.659
Gi-jón	43°32' N	005°40' W	—	08:51:23.1	292	339	36	—	—	—	—	—	—	—	—	—	—	10:07:48.3	13	53	49	122	0.820	0.779
Granada	37°13' N	003°41' W	—	08:52:54.0	304	358	39	—	—	—	—	—	—	—	—	—	—	10:08:53.3	14	60	54	118	0.629	0.543
Hospitalet	41°22' N	002°08' E	—	09:00:44.0	297	344	44	—	—	—	—	—	—	—	—	—	—	10:21:24.8	15	49	57	135	0.771	0.717
Huelva	37°16' N	006°57' W	—	08:48:37.3	303	358	36	—	—	—	—	—	—	—	—	—	—	10:02:30.1	13	62	50	113	0.628	0.541
Jaén	37°46' N	003°47' W	—	08:52:38.8	303	357	39	—	—	—	—	—	—	—	—	—	—	10:08:51.0	13	60	53	119	0.646	0.564
Jerez de la Fro...	36°41' N	006°08' W	—	08:49:45.9	305	0	37	—	—	—	—	—	—	—	—	—	—	10:03:49.0	13	63	51	114	0.610	0.520
La Coruña	43°22' N	008°23' W	—	08:48:24.2	292	340	34	—	—	—	—	—	—	—	—	—	—	10:03:21.1	12	54	47	118	0.814	0.771
Leganés	40°19' N	003°45' W	—	08:52:38.1	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:46.0	13	56	52	122	0.725	0.660
León	42°36' N	005°34' W	—	08:51:04.1	294	342	37	—	—	—	—	—	—	—	—	—	—	10:07:30.2	13	54	50	122	0.792	0.744
Lérida	41°37' N	000°37' E	—	08:58:36.5	296	344	43	—	—	—	—	—	—	—	—	—	—	10:18:27.1	15	50	55	132	0.774	0.721
Logroño	42°28' N	002°27' W	—	08:54:45.0	295	342	39	—	—	—	—	—	—	—	—	—	—	10:12:54.5	14	52	52	127	0.793	0.744
Madrid	40°24' N	003°41' W	667	08:52:43.1	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:54.8	13	56	52	122	0.728	0.663
Málaga	36°43' N	004°25' W	—	08:52:02.0	305	0	39	—	—	—	—	—	—	—	—	—	—	10:07:14.7	13	62	53	116	0.612	0.524
Mataró	40°32' N	002°27' E	—	09:01:12.2	297	343	44	—	—	—	—	—	—	—	—	—	—	10:22:03.3	15	48	57	136	0.777	0.725
Móstoles	40°19' N	003°51' W	—	08:52:30.5	298	349	39	—	—	—	—	—	—	—	—	—	—	10:09:34.8	13	56	52	122	0.725	0.660
Murcia	37°59' N	001°07' W	—	08:56:19.5	303	355	42	—	—	—	—	—	—	—	—	—	—	10:14:22.0	14	57	56	124	0.658	0.578
Orense	42°20' N	007°51' W	—	08:48:26.3	294	343	34	—	—	—	—	—	—	—	—	—	—	10:03:34.4	12	55	48	118	0.783	0.732
Pavia	43°22' N	005°50' W	—	08:51:07.0	290	340	36	—	—	—	—	—	—	—	—	—	—	10:07:26.6	13	53	49	122	0.815	0.772
Palma (de Mallo...	39°24' N	002°39' E	—	09:01:35.8	300	349	45	—	—	—	—	—	—	—	—	—	—	10:22:26.1	15	51	58	134	0.718	0.652
Pamploña	42°49' N	001°38' W	—	08:55:53.5	294	341	40	—	—	—	—	—	—	—	—	—	—	10:14:29.9	14	51	53	129	0.805	0.760
Sabadell	41°33' N	002°06' E	—	09:00:41.9	297	343	44	—	—	—	—	—	—	—	—	—	—	10:21:21.5	15	49	57	135	0.777	0.724
Salamanca	40°58' N	005°39' W	—	08:50:25.7	297	347	37	—	—	—	—	—	—	—	—	—	—	10:06:34.7	13	56	50	120	0.743	0.682
San Cristóbal d...	28°29' N	016°19' W	—	08:42:38.1	319	24	27	—	—	—	—	—	—	—	—	—	—	09:40:54.5	10	76	40	94	0.362	0.248
San Sebastián	43°19' N	001°59' W	—	08:5																				

TABLE 24

LOCAL CIRCUMSTANCES FOR EUROPE - SLOVAKIA, SLOVENIA, SWEDEN, SWITZERLAND, WALES & YUGOSLAVIA
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact			Second Contact			Third Contact			Fourth Contact			Maximum Eclipse			Eclip. Mag.	Eclip. Obs.	Umbral Durat.		
				U.T.	P	V	U.T.	P	V	U.T.	P	V	U.T.	P	V	U.T.	P	V				Alt	Alt Azm
				h	m	s	h	m	s	h	m	s	h	m	s	h	m	s	o	o	o		
SLOVAKIA																							
Bratislava	48°09' N	017°07' E	—	09:25:02.4	285	310	52	—	—	—	—	—	12:10:11.2	110	89	54	10:47:51.5	198	200	57	176	0.987	0.989
Kosice	48°43' N	021°15' E	—	09:31:56.5	283	302	54	—	—	—	—	—	12:15:19.0	113	88	51	10:54:26.8	198	194	56	186	0.944	0.937
SLOVENIA																							
Ljubljana	46°03' N	014°31' E	—	09:20:28.8	288	319	52	—	—	—	—	—	12:08:34.9	106	86	57	10:44:24.8	17	25	59	169	0.968	0.966
Maribor	46°33' N	015°39' E	—	09:22:24.1	287	316	52	—	—	—	—	—	12:09:53.4	107	86	56	10:46:12.6	18	23	59	172	0.989	0.991
SWEDEN																							
Borås	57°43' N	012°55' E	—	09:24:48.3	270	290	43	—	—	—	—	—	11:52:14.7	121	113	47	10:38:11.8	196	203	47	167	0.763	0.707
Göteborg	57°43' N	011°58' E	17	09:23:40.9	270	291	43	—	—	—	—	—	11:51:10.4	120	114	47	10:37:01.2	195	203	47	166	0.767	0.712
Helsingborg	56°03' N	012°42' E	—	09:23:01.4	273	294	44	—	—	—	—	—	11:54:03.7	118	110	49	10:38:10.3	196	203	49	167	0.805	0.760
Jönköping	57°47' N	014°11' E	—	09:26:23.3	270	289	43	—	—	—	—	—	11:53:34.0	121	112	47	10:39:45.2	196	201	47	170	0.756	0.699
Lindköping	58°25' N	015°37' E	—	09:28:41.4	269	286	44	—	—	—	—	—	11:54:15.4	122	113	46	10:41:22.0	196	200	47	172	0.735	0.672
Malmö	55°36' N	013°00' E	—	09:23:01.0	273	295	45	—	—	—	—	—	11:54:58.9	118	109	49	10:38:39.0	196	203	49	167	0.815	0.773
Norrköping	58°36' N	016°11' E	—	09:29:32.2	269	285	44	—	—	—	—	—	11:54:35.5	123	113	46	10:42:00.0	196	199	47	173	0.728	0.664
Örebro	59°17' N	015°13' E	—	09:29:00.0	268	284	43	—	—	—	—	—	11:52:39.1	123	115	45	10:40:41.3	196	200	46	172	0.716	0.649
Stockholm	59°20' N	018°03' E	45	09:32:22.5	267	281	44	—	—	—	—	—	11:55:24.3	124	113	45	10:43:58.9	196	198	46	177	0.703	0.633
Uppsala	59°52' N	017°38' E	—	09:32:22.5	267	281	44	—	—	—	—	—	11:54:12.8	125	114	44	10:43:19.6	196	198	45	176	0.693	0.620
Västerås	59°37' N	016°33' E	—	09:30:52.4	267	282	43	—	—	—	—	—	11:53:31.3	124	115	45	10:42:09.4	196	199	46	174	0.703	0.633
SWITZERLAND																							
Basel (Bâle)	47°33' N	007°35' E	—	09:10:16.8	287	322	46	—	—	—	—	—	11:55:10.4	105	98	58	10:31:29.2	16	35	55	152	0.973	0.973
Bern (Berne)	46°57' N	007°26' E	572	09:09:47.4	288	324	46	—	—	—	—	—	11:55:20.9	104	98	58	10:31:16.8	16	36	56	152	0.956	0.951
Genève (Geneva)	46°12' N	006°09' E	405	09:07:39.0	289	327	46	—	—	—	—	—	11:53:30.3	102	98	59	10:29:03.3	16	38	56	148	0.929	0.917
Lausanne	46°31' N	006°38' E	—	09:08:28.3	288	326	46	—	—	—	—	—	11:54:11.8	103	98	59	10:29:54.2	16	37	56	150	0.940	0.931
Lucern	47°03' N	008°18' E	—	09:11:05.4	287	323	47	—	—	—	—	—	11:56:49.1	104	96	58	10:32:48.2	16	34	56	154	0.962	0.959
Sankt Gallen	47°25' N	009°23' E	—	09:12:49.3	287	321	48	—	—	—	—	—	11:58:24.9	106	96	57	10:34:38.9	16	32	56	157	0.978	0.978
Winterthur	47°30' N	008°43' E	—	09:11:52.9	287	321	47	—	—	—	—	—	11:57:11.4	105	97	57	10:33:27.9	16	33	56	155	0.977	0.977
Zürich	47°23' N	008°32' E	493	09:11:33.6	287	322	47	—	—	—	—	—	11:56:57.8	105	97	57	10:33:09.6	16	34	56	155	0.973	0.972
WALES																							
Cardiff	51°29' N	003°13' W	62	09:00:22.5	280	316	37	—	—	—	—	—	11:35:11.5	107	118	53	10:15:36.9	193	221	46	135	0.973	0.972
Newport	51°35' N	003°00' W	—	09:00:41.9	280	316	37	—	—	—	—	—	11:35:30.2	107	118	53	10:15:57.6	193	221	46	135	0.970	0.969
Port Talbot	51°36' N	003°47' W	—	08:59:56.6	280	316	36	—	—	—	—	—	11:34:17.0	107	119	53	10:14:54.8	193	221	46	134	0.971	0.969
Swansea	51°38' N	003°57' W	—	08:59:49.1	280	316	36	—	—	—	—	—	11:34:01.1	107	119	53	10:14:42.5	193	221	46	134	0.970	0.968
YUGOSLAVIA																							
Beograd (Belgra...)	44°50' N	020°30' E	138	09:30:54.1	289	313	57	—	—	—	—	—	12:19:54.0	108	78	54	10:56:23.5	19	14	60	186	0.977	0.977
Kikinda	45°50' N	020°28' E	—	09:30:40.0	288	311	56	—	—	—	—	—	12:18:26.4	109	80	54	10:55:27.7	19	14	59	186	1.029	1.000
Kragujevac	43°01' N	020°55' E	—	09:31:53.0	291	315	58	—	—	—	—	—	12:21:41.5	107	75	54	10:57:55.0	19	13	61	188	0.959	0.955
Nis	43°19' N	021°54' E	—	09:33:58.2	292	314	59	—	—	—	—	—	12:24:13.9	106	72	54	11:00:27.9	19	10	62	192	0.948	0.942
Novi Sad	45°15' N	019°50' E	—	09:29:37.6	289	313	56	—	—	—	—	—	12:18:16.7	108	79	54	10:54:47.9	19	15	60	184	0.983	0.984
Platina	42°39' N	021°10' E	—	09:32:52.5	293	318	59	—	—	—	—	—	12:23:58.2	105	71	55	10:59:43.1	19	12	62	190	0.924	0.911
Subotica	46°06' N	019°39' E	—	09:29:12.1	287	311	55	—	—	—	—	—	12:16:49.2	109	82	54	10:53:46.0	18	16	59	183	1.029	1.000

TABLE 30
LOCAL CIRCUMSTANCES FOR IRAQ
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact		Second Contact		Third Contact		Fourth Contact		Maximum Eclipse		Eclip. Obs.	Umbra Durat.
				U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °		
IRAQ															
Al-'Amarah	31°50'N	047°09'E	-	10:38:22.2	297 246 62	-	-	-	-	13:12:00.7	108 46 30	11:59:56.7	22 322 45 259	0.946	0.938
Al-Basrah (Basr...)	30°30'N	047°47'E	-	10:41:53.5	298 244 61	-	-	-	-	13:14:29.4	107 43 29	12:03:00.9	23 320 44 262	0.923	0.909
Al-Hillah	32°29'N	044°25'E	-	10:32:10.2	298 253 64	-	-	-	-	13:09:01.5	107 46 33	11:55:21.3	23 324 48 256	0.927	0.915
Al-Mawsil (Mosul...)	36°20'N	043°08'E	223	10:23:29.0	293 259 64	-	-	11:47:08.4	216 164	13:01:34.7	111 55 35	11:46:53.7	22 331 50 249	1.027	1.000
An-Najaf	31°59'N	044°20'E	-	10:32:53.6	298 252 64	-	-	-	-	13:09:44.8	107 45 33	11:56:07.2	23 324 48 256	0.915	0.899
Ar-Nasiriyah	31°02'N	046°10'E	-	10:38:11.7	298 247 63	-	-	-	-	13:12:39.5	107 44 31	12:00:15.7	23 321 46 260	0.917	0.901
Ar-Ramadi	33°25'N	043°17'E	-	10:28:24.0	292 256 65	-	-	-	-	13:06:37.4	108 48 34	11:52:11.5	22 326 50 253	0.935	0.925
As-Sulaymaniyah	33°33'N	045°26'E	-	10:29:03.6	292 253 62	-	-	11:52:16.0	277 223	13:04:39.3	111 54 33	11:51:18.2	22 328 48 253	1.026	1.000
Baghdad	33°21'N	044°25'E	34	10:30:41.0	296 253 64	-	-	-	-	13:07:37.9	108 48 33	11:53:50.2	22 326 48 255	0.947	0.940
Ba 'Qubah	33°45'N	044°38'E	-	10:30:25.5	296 253 64	-	-	-	-	13:07:08.3	109 49 33	11:53:25.3	22 326 48 254	0.959	0.955
Dabuk	36°52'N	043°00'E	-	10:22:26.6	292 259 64	11:44:45.7	122 72	11:46:48.2	281 231	13:00:31.8	110 56 36	11:45:47.1	22 332 50 248	1.027	1.000
Halabjah	35°10'N	045°59'E	-	10:30:41.1	293 251 62	11:51:46.2	137 83	11:53:34.0	267 212	13:05:42.0	111 53 32	11:52:40.3	22 327 47 254	1.026	1.000
Irbil	36°11'N	044°01'E	-	10:25:24.5	292 256 63	11:47:25.9	139 87	11:49:15.6	265 213	13:02:31.3	111 55 34	11:48:20.9	22 330 49 250	1.026	1.000
Karbala'	32°36'N	044°02'E	-	10:31:14.3	298 254 65	-	-	-	-	13:08:32.3	107 46 33	11:54:38.4	23 325 49 255	0.925	0.912
Kirkuk	35°28'N	044°28'E	-	10:27:21.5	293 255 63	-	-	-	-	13:04:06.4	111 53 34	11:50:11.8	22 329 48 252	0.996	0.998
KuyseanJaq	36°05'N	044°38'E	-	10:26:14.1	292 255 63	11:48:19.4	116 63	11:50:20.7	288 235	13:03:09.1	112 55 34	11:49:20.2	22 329 48 251	1.026	1.000
Zakhu	37°08'N	042°41'E	-	10:21:26.7	292 260 64	11:43:50.5	112 63	11:45:55.7	292 242	12:59:48.2	112 56 36	11:44:53.3	22 332 50 247	1.027	1.000

TABLE 31
LOCAL CIRCUMSTANCES FOR PAKISTAN
TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

Location Name	Latitude	Longitude	Elev. m	First Contact		Second Contact		Third Contact		Fourth Contact		Maximum Eclipse		Eclip. Obs.	Umbra Durat.
				U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °	U.T. h m s	P ° V ° Alt °		
PAKISTAN															
Chamburi Kalat	26°09'N	064°43'E	-	11:13:43.3	292 223 41	12:23:14.3	65 357	12:24:09.8	337 269	13:25:49.2	110 45 11	12:23:42.2	201 133 95 276	1.020	1.000
Faisalabad (Iya...)	31°25'N	073°05'E	-	11:14:13.3	291 221 37	-	-	-	-	13:16:56.0	120 61 7	12:18:43.3	200 138 19 277	0.807	0.761
Gharo	24°44'N	067°35'E	-	11:18:50.0	291 221 33	12:26:22.8	84 15	12:27:28.6	317 248	13:27:32.7	110 43 8	12:26:55.8	201 132 22 278	1.019	1.000
Gujranwala	32°26'N	074°33'E	-	11:14:08.5	277 216 32	-	-	-	-	13:14:55.1	122 64 7	12:17:31.5	200 139 18 277	0.774	0.718
Hab Nadi Chowki	25°01'N	066°53'E	-	11:17:43.6	291 221 38	12:25:41.6	88 19	12:26:50.8	313 244	13:27:14.2	110 44 9	12:26:16.3	201 132 22 277	1.020	1.000
Hirok Sami	26°02'N	063°25'E	-	11:12:29.1	293 224 42	12:22:55.8	173 105	12:23:34.2	229 161	13:25:57.5	110 44 12	12:23:15.1	21 313 26 275	1.021	1.000
Hoshab	26°01'N	063°56'E	-	11:13:05.1	293 223 42	12:22:54.7	134 66	12:24:09.0	268 200	13:25:59.7	110 44 12	12:23:31.9	21 313 26 275	1.021	1.000
Hyderabad	25°22'N	068°22'E	-	11:18:35.2	290 220 36	-	-	-	-	13:26:38.4	111 45 8	12:28:17.3	201 132 21 278	0.983	0.983
Islamabad	33°42'N	073°10'E	-	11:11:15.4	276 216 34	-	-	-	-	13:13:07.7	123 66 8	12:15:13.1	200 140 20 275	0.761	0.702
Jati	24°21'N	068°16'E	-	11:20:03.1	291 220 36	12:27:06.9	96 27	12:28:17.0	305 235	13:27:57.6	110 43 7	12:27:42.0	201 131 21 278	1.019	1.000
Jungshahi	24°51'N	067°46'E	-	11:18:49.4	291 221 37	12:26:34.8	44 335	12:27:04.2	357 288	13:27:22.6	110 44 8	12:26:49.6	201 132 21 278	1.019	1.000
Kandrach	25°29'N	065°29'E	-	11:15:35.9	292 222 40	12:24:23.0	110 42	12:25:40.6	291 223	13:26:41.9	110 44 10	12:25:01.9	201 132 24 276	1.020	1.000
Karachi	24°52'N	067°03'E	4	11:18:07.3	292 221 38	12:25:56.4	99 30	12:27:09.3	303 234	13:27:25.1	110 43 9	12:26:32.9	201 132 22 277	1.020	1.000
Keti Bandar	24°08'N	067°27'E	-	11:19:38.7	292 221 37	-	-	-	-	13:28:18.9	109 42 8	12:27:44.2	21 311 21 278	0.999	1.000
Kotri Allahrakht...	24°24'N	067°50'E	-	11:19:34.9	292 221 37	12:26:52.9	116 46	12:28:05.6	286 216	13:27:56.6	110 43 8	12:27:29.3	21 311 21 278	1.019	1.000
Lahore	31°35'N	074°18'E	-	11:15:02.8	278 216 32	-	-	-	-	13:16:22.1	121 62 6	12:18:45.6	200 138 18 277	0.793	0.743
Mirpur Batoro	24°44'N	068°16'E	-	11:19:27.8	291 220 36	-	-	-	-	13:27:28.4	110 43 8	12:27:09.5	201 132 21 278	0.998	0.999
Mirpur Sakro	24°33'N	067°37'E	-	11:19:09.0	292 221 37	12:26:35.4	108 39	12:27:48.7	293 224	13:27:46.5	110 44 8	12:27:12.1	201 131 21 278	1.019	1.000
Multan	30°11'N	071°29'E	122	11:14:28.1	292 218 34	-	-	-	-	13:19:13.5	118 57 8	12:20:09.3	200 137 20 276	0.848	0.812
Naka Kharari	25°15'N	066°44'E	-	11:17:13.0	291 221 38	12:25:28.5	60 352	12:26:16.5	341 273	13:26:56.7	110 44 9	12:25:52.6	201 132 23 277	1.020	1.000
Nirwano	26°22'N	062°43'E	-	11:11:08.5	293 221 43	12:22:02.0	171 103	12:22:43.7	231 163	13:25:28.4	110 44 13	12:22:23.0	201 131 27 274	1.021	1.000
Peshawar	34°01'N	071°33'E	-	11:09:18.1	277 218 35	-	-	-	-	13:12:54.8	123 65 10	12:14:13.6	200 141 22 274	0.769	0.713
Quetta	30°12'N	067°00'E	-	11:09:58.0	285 221 39	-	-	-	-	13:19:48.1	116 55 11	12:18:31.2	201 137 24 274	0.890	0.866
Rahim ki Bazar	24°19'N	069°09'E	-	11:20:53.0	291 220 35	-	-	-	-	13:27:53.4	110 43 7	12:28:01.0	200 131 20 278	0.999	1.000
Rawalpindi	33°36'N	073°04'E	511	11:11:18.2	276 217 34	-	-	-	-	13:13:19.5	123 65 8	12:15:21.1	200 140 20 275	0.763	0.706
Sargodha	32°05'N	072°40'E	-	11:12:57.2	278 217 34	-	-	-	-	13:15:56.7	110 42 8	12:17:35.4	200 139 20 276	0.797	0.748
Shahbandar	24°10'N	067°54'E	-	11:20:00.3	292 221 36	12:27:20.7	146 76	12:28:19.9	255 186	13:28:13.8	109 42 8	12:27:50.4	21 311 21 278	1.019	1.000
Shahbaz Kalat	26°42'N	063°58'E	-	11:14:01.0	291 223 42	-	-	-	-	13:25:03.7	111 45 12	12:22:30.9	201 134 26 275	0.998	0.999
Shikhot	32°30'N	074°31'E	-	11:14:01.6	277 216 32	-	-	-	-	13:14:49.1	122 64 7	12:17:25.0	200 139 19 277	0.773	0.717
Somiani	25°26'N	065°36'E	-	11:16:47.9	291 221 38	-	-	-	-	13:26:42.7	110 44 9	12:25:33.5	201 132 93 277	0.999	1.000
Sujawal	24°36'N	068°05'E	-	11:19:30.1	291 221 36	12:26:49.7	70 1	12:27:44.7	331 262	13:27:39.8	110 43 8	12:27:17.3	201 131 21 278	1.019	1.000
Tatta	24°45'N	067°55'E	-	11:19:07.0	291 221 37	12:26:41.1	54 345	12:27:21.1	347 278	13:27:29.4	110 44 8	12:27:01.2	201 132 21 278	1.019	1.000

TABLE 37

SOLAR ECLIPSES OF SAROS SERIES 145

First Eclipse: 1639 Jan 04 Duration of Series: 1370.3 yrs.
 Last Eclipse: 3009 Apr 17 Number of Eclipses: 77
 Saros Summary: Partial: 34 Annular: 1 Total: 41 Hybrid: 1

Eclipse					Eclipse				
Date	Type	Gamma	Magn. / Width	Center Durat.	Date	Type	Gamma	Magn. / Width	Center Durat.
1639	Jan 04	Pb	1.565	0.001	2360	Mar 18	T	-0.017	181 04n03s
1657	Jan 14	P	1.554	0.017	2378	Mar 29	T	-0.048	193 04n51s
1675	Jan 25	P	1.543	0.034	2396	Apr 09	T	-0.085	206 05m2s
1693	Feb 05	P	1.527	0.059	2414	Apr 20	T	-0.128	217 05n03s
1711	Feb 17	P	1.507	0.091	2432	Apr 30	T	-0.178	229 05n56s
1729	Feb 27	P	1.481	0.134	2450	May 12	T	-0.233	241 06m9s
1747	Mar 11	P	1.450	0.186	2468	May 22	T	-0.293	252 06m41s
1765	Mar 21	P	1.412	0.251	2486	Jun 02	T	-0.358	263 06n59s
1783	Apr 01	P	1.367	0.329	2504	Jun 14	T	-0.427	275 07m0s
1801	Apr 13	P	1.315	0.420	2522	Jun 25	T	-0.499	287 07m2s
1819	Apr 24	P	1.258	0.522	2540	Jul 05	T	-0.572	300 07n04s
1837	May 04	P	1.193	0.638	2558	Jul 16	T	-0.646	315 06m43s
1855	May 16	P	1.125	0.762	2576	Jul 27	T	-0.720	334 06m2s
1873	May 26	P	1.051	0.897	2594	Aug 07	T	-0.792	361 05n02s
1891	Jun 06	A	0.975	33 00m06s	2612	Aug 18	T	-0.862	406 04m45s
1909	Jun 17	H	0.896	51 00m24s	2630	Aug 30	T	-0.930	512 03n54s
1927	Jun 29	T	0.816	77 00m50s	2648	Sep 09	Ts	-0.992	- 02m49s
1945	Jul 09	T	0.736	92 01m15s	2666	Sep 20	P	-1.050	0.919
1963	Jul 20	T	0.657	101 01m40s	2684	Oct 01	P	-1.103	0.817
1981	Jul 31	T	0.579	108 02m02s	2702	Oct 13	P	-1.150	0.727
1999	Aug 11	T	0.506	112 02m23s	2720	Oct 23	P	-1.191	0.648
2017	Aug 21	T	0.437	115 02m40s	2738	Nov 04	P	-1.225	0.584
2035	Sep 02	T	0.373	116 02m54s	2756	Nov 14	P	-1.255	0.528
2053	Sep 12	T	0.314	116 03m04s	2774	Nov 25	P	-1.278	0.486
2071	Sep 23	T	0.262	116 03m11s	2792	Dec 06	P	-1.297	0.451
2089	Oct 04	T	0.217	115 03m14s	2810	Dec 17	P	-1.311	0.426
2107	Oct 16	T	0.178	114 03m16s	2828	Dec 27	P	-1.322	0.405
2125	Oct 26	T	0.146	112 03m15s	2847	Jan 08	P	-1.331	0.389
2143	Nov 07	T	0.121	111 03m14s	2865	Jan 18	P	-1.339	0.374
2161	Nov 17	T	0.101	110 03m13s	2883	Jan 30	P	-1.348	0.359
2179	Nov 28	T	0.087	110 03m12s	2901	Feb 10	P	-1.359	0.339
2197	Dec 09	T	0.077	111 03m13s	2919	Feb 21	P	-1.371	0.316
2215	Dec 21	T	0.070	114 03m14s	2937	Mar 04	P	-1.389	0.285
2233	Dec 31	T	0.065	117 03m18s	2955	Mar 15	P	-1.411	0.245
2252	Jan 12	T	0.061	123 03m23s	2973	Mar 25	P	-1.439	0.194
2270	Jan 22	T	0.056	129 03m29s	2991	Apr 06	P	-1.472	0.134
2288	Feb 02	T	0.050	138 03m08s	3009	Apr 17	Pe	-1.514	0.059
2306	Feb 14	T	0.040	147 03m49s					
2324	Feb 25	Th	0.026	158 04m02s					
2342	Mar 08	T	0.008	169 04m16s					

Eclipse Type: P - Partial Pb - Partial Eclipse (Saros Series Begins)
 A - Annular Pe - Partial Eclipse (Saros Series Ends)
 T - Total Ts - Total Eclipse (no southern limit)
 H - Hybrid (Annular/Total)

Note: Magn./Width column gives either the eclipse magnitude (for partial eclipses) or the umbral path width in kilometers (for total and annular eclipses).

TABLE 38

**CLIMATOLOGICAL STATISTICS FOR AUGUST ALONG THE ECLIPSE PATH OF
THE TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Station	Latitude	Longitude	T max °F	T min °F	Days with Rain	Days with TKW	Days with fast Cloud	POR	Hours of sun- shine	% of Possible sun- shine	Mean Cloud Cover (10ths)	Days with Fog
England												
Falmouth	50°08N	5°02W	67	56	7.5	0.7	1.0	4				
The Lizard	49°57N	5°12W	67	56	4.5	0.4	6.5	8				
Plymouth	50°22N	4°07W	66	55	7.3	1	5.8	11	6.4	43		
Bill of Portland	50°32N	2°27W	65	58	5.3	0.8	6.5	11				
Cardrose	50°05N	5°15W	67	56	4.5	0.4	6.5	8				
Bournemouth	50°47N	1°50W	68	55	6.6	2.0	6.7	11				
Guernsey	49°26N	2°36W	68	57	4.6	0.3	8.3	11				
France												
Clerbourg	49°39N	1°28W	68	58	7.4	2.0	n/a	n/a				
Caen	49°10N	0°26W	73	53	6.0	3.0	n/a	n/a				
Le Touquet	50°31N	1°37E	70	56	8.2	n/a	6.6	9				
Le Havre	42°39N	0°05E	73	56	7.0	2.2	10.0	8			8.0	2.4
Bureux	49°01N	1°13E	71	53	5.0	2.5	2.7	14				
Boisbous	49°08N	2°02E	74	55	5.9	3.0	5.4	11				
Beauvais	49°27N	2°06E	73	53	6.1	4.0	2.0	4				
Creil	49°15N	2°31E	73	54	6.8	4.0	3.5	2				
Paris - Le Bourget	48°58N	2°28E	75	56	6.5	4.0	7.2	9	6.6	45	6.9	1.5
Paris - Orly	48°43N	2°24E	75	56	6.7	3.2	5.4	11				
Lao	49°38N	3°33E	72	54	6.2	4.4	2.3	15				
Leims	49°19N	4°03E	76	54	6.2	4.0	5.1	9				
Laippes	49°09N	4°38E	75	55	7.6	2.0	3.7	5				
Voisiers	49°16N	4°45E	75	55	7.1	2.0	3.7	5				
Ehuis	49°14N	5°40E	71	53	5.4	3.4	3.3	12				
Metz	49°04N	6°08E	75	55	6.6	3.0	5.3	4				
Dizier	48°38N	4°58E	75	54	7.1	n/a	8.5	3				
Toul	48°47N	5°59E	72	54	8.2	4.7	3.2	16				
Nancy	48°42N	6°14E	73	54	6.6	5.0	2.2	13				
Flakbourg	48°46N	7°12E	71	55	8.6	5.0	3.6	13				
Strasbourg	48°32N	7°37E	76	55	7.7	5.0	8.7	12	7.0	48	7.4	3.5
Clamby	49°01N	5°52E	69	53	6.2	2.1	2.9	11				
Gros-Tourgis	49°01N	6°43E	75	55	6.6	3.0	5.3	4				
Montmedy	49°27N	5°25E	71	53	8.3	3.4	3.3	12				
Kocroi	49°54N	4°25E	69	54	8.6	2.0	1.5	8				
Belgium												
Charleroi	50°27N	4°27E	69	54	7.7	2.0	1.5	8				
Florennes	50°14N	4°39E	69	54	9.5	2.0	1.5	8				
St-Hibert	50°02N	5°24E	67	51	8.5	4.0	1.8	8				
Luxembourg												
Luxembourg	49°37N	6°12E	70	54	8.3	4.5	3.3	8	6.5	44	5.9	4

Abbreviations:

Tmax - average daily maximum temperature (°F).

Tmin - average daily minimum temperature (°F).

TRW - thunderstorms.

POR - period of record (years).

TABLE 38 - continued

**CLIMATOLOGICAL STATISTICS FOR AUGUST ALONG THE ECLIPSE PATH OF
THE TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Station	Latitude	Longitude	T max T	T min T	Days with Rain	Days with TKW	Days with Scat Cloud	POB	Hours of sun- shine	% of Possible sun- shine	Mean Cloud Cover (10ths)	Days with Fog
Germany												
Trier	49°43'N	6°36'E	74	52	7.2	4.3	4.6	6				
Kamstein	49°26'N	7°36'E	72	57	7.3	3.8	2.5	16				
Zembaich	49°30'N	7°52'E	70	53	8.1	4.2	2.4	15				
Stthart	48°41'N	9°12'E	74	55	6.8	4.0	4.1	15				
Karlsruhe	49°01'N	8°23'E	74	56	7.6	6.8	6.8	14	7.0	48	61	1.7
Oltrisa	49°12'N	9°31'E	76	55	7.2	ah	6.0	6				
Wurzberg	49°48'N	9°54'E	73	54	6.0	4.0	6.1	6				
Stots	48°40'N	9°52'E	69	55	8.2	4.7	6.3	6				
Ulm	48°24'N	9°59'E	72	52	7.9	6.5	5.9	10				
Wessersberg	49°02'N	10°58'E	72	53	7.5	4.4	6.1	8				
Aszberg	48°23'N	10°51'E	72	54	8.4	5.5	6.9	11	7.2	49	51	1.7
Fischerfeld	48°12'N	11°16'E	73	54	8.6	5.4	4.6	12				
Marsberg (Kern)	48°08'N	11°42'E	72	54	8.9	6.0	7.5	10				
Endau	48°19'N	11°56'E	74	54	8.0	4.5	4.9	12				
Pasau	48°35'N	13°29'E	72	54	8.3	3.4	7.1	10				
Austria												
Linz	48°14'N	14°11'E	73	56	8.8	4.0	9.0	9				
Falkenberg	47°47'N	13°00'E	73	55	9.9	4.0	8.0	11	7.0	48	55	1.6
Alpe	47°32'N	14°08'E	74	49	10.4	3.7	9.4	9				
Zellmoos	47°12'N	14°44'E	73	48	9.3	ah	5.2	5				
Wien (Vienna)	48°15'N	16°22'E	73	58	7.0	6.0	6.8	5	8.1	56	4.9	0.8
Sollnegg	48°07'N	16°34'E	79	57	2.4	1.9	11.0	5				
Hungary												
Szombathely	47°17'N	16°37'E	77	55	6.2	5.7	7.5	8	7.7	53	5.3	0.5
Keszthely	46°46'N	17°12'E	81	60	2.4	0.7	14.0	5				
Pecs	46°06'N	18°13'E	81	57	5.3	3.3	10.8	5	9.0	63	4.0	0.8
Budapest	47°31'N	19°02'E	81	61	4.7	4.4	8.2	8	8.7	60	4.3	0.1
Szolnok	47°11'N	20°13'E	84	58	2.8	1.8	10.9	8				
Szeged	46°15'N	20°06'E	82	58	4.1	4.6	12.0	8	9.2	64	3.9	0.3
Romania												
Bucarest	44°30'N	26°06'E	86	60	3.6	5.4	14.5	16			3.8	1.4
Cimpas	45°17'N	25°02'E	75	50	4.9	5.4	10.8	5				
Deva	45°53'N	22°54'E	83	56	6.4	6.3	13.2	16				
Sibiu	45°48'N	24°09'E	78	55	7.1	6.4	11.5	10				
Oradea	45°38'N	25°27'E	49	39	7.4	7.8	2.9	16				
Oradea	47°03'N	21°56'E	82	58	5.8	5.7	11.5	9				
Arad	46°11'N	21°19'E	82	58	4.8	6.4	11.5	16				
Timisoara	45°45'N	21°14'E	84	58	5.2	4.7	9.5	11			3.5	0.4
Trasi-Severin	44°38'N	22°38'E	86	61	4.5	4.6	14.6	16				
Constata	44°11'N	28°40'E	80	63	2.8	2.8	18.0	16			2.9	1.8
Trasi-Magurele	43°45'N	24°52'E	86	61	3.3	2.6	17.3	10				
Bulgaria												
Varna	43°12'N	27°55'E	81	64	1.5	2.3	17.0	16	10.1	71	3.0	0.8
Burgas	42°29'N	27°29'E	82	64	2.6	1.8	19.4	12				
Blvna	42°41'N	26°16'N	84	63	2.5	2.5	15.2	15				
Kozle	43°52'N	25°58'E	86	64	3.4	5.3	17.5	15				

TABLE 38 - continued

**CLIMATOLOGICAL STATISTICS FOR AUGUST ALONG THE ECLIPSE PATH OF
THE TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11**

Station	Latitude	Longitude	T max °	T min °	Days with Rain	Days with TKW	Days with Scat Cloud	POB	Hours of Sun- shine	% of Possible Sun- shine	Mean Cloud Cover (10ths)	Days with Fog
Turkey												
Zonguldak	42°27'N	31°40'E	78	64	5.1	4.4	16.2	7				
Erzurum	42°01'N	35°09'E	78	68	3.6	3.6	17.6	7				
Yozgat	39°49'N	36°54'E	82	51	0.5	1.2	19.1	6		86		0.2
Ankara	39°57'N	32°42'E	88	58	0.7	1.2	16.1	10		87		0.1
Amasya	41°17'N	36°20'E	80	65	2.3	3.0	13.2	10		66		0
Yesilkoy (Istanbul)	40°58'N	28°49'E	84	66	2.6	1.8	16.6	11		80		0.6
Kastamonu	41°22'N	33°46'E	80	54	3.4	3.2	13.3	6				
Merzifon	40°49'N	35°31'E	85	58	0.6	1.7	15.2	11				
Erbaa	38°28'N	38°09'E	94	65	0.1	0.0	24.8	10				
Erzincan	39°42'N	39°31'E	89	59	0.7	3.0	16.8	11				
Erzurum	39°57'N	41°10'E	80	59	1.2	3.4	11.7	10		82		0
Van	38°28'N	43°20'E	82	56	0.3	1.6	24.2	7				
Mahya	38°21'N	38°19'E	94	65	0.1	0.0	24.8	10				
Diyarbakir	37°53'N	40°13'E	101	70	0.0	0.0	23.1	10		49?		0
Batman	37°55'N	41°07'E	101	70	0.0	0.0	23.1	10				
Siirt	37°53'N	41°52'E	98	73	0.2	0.6	27.6	6				
Syria												
AlQamishli	37°01'N	41°11'E	105	75	0.0	0.0	29.3	11		91		0
Iraq												
Mosul	36°18'N	43°08'E	110	70	0.1	0.1	27.3	14				
Kirkuk	35°28'N	44°21'E	108	79	0.1	0.2	28.3	13				
Klatana	34°18'N	45°29'E	110	78	0.1	0.1	26.6	11				
Iran												
Kozangol	37°32'N	45°09'E	90	63	0.2	0.8	26.2	5				
Kermanshal	34°19'N	47°07'E	99	55	0.1	0.0	25.7	10				
Hamadan	34°38'N	48°31'E	90	57	0.4	0.2	21.7	6				
Bafq	32°37'N	51°42'E	96	63	0.2	0.2	24.4	10				
Dezful	32°29'N	48°24'E	113	80	0.0	0.0	26.4	14				
Yazd	31°54'N	54°24'E	100	71	0.0	0.0	25.3	7				
Kerman	30°15'N	56°57'E	95	62	0.1	0.4	20.6	7				
Shiraz	29°32'N	52°39'E	97	65	0.0	0.4	21.6	9				
Zahedan	29°27'N	60°54'E	96	63	0.0	0.1	24.2	8				
Pakistan												
Imran	25°04'N	61°48'E	88	79	0.2	0.0	12.8	10				
Karachi	24°54'N	67°09'E	88	79	2.8	1.8	4.9	5				
India												
Bij	23°16'N	69°40'E	88	76	4.8	1.0	2.7	10			5.9	0
Kajkot	22°18'N	70°47'E	89	75	7.8	1.0	1.6	8			6.5	0
Amalabad	23°03'N	72°37'E	90	77	11.0	1.0	0.3	10	4.3	33	6.7	0
Surat	21°12'N	72°50'E	87	77	10.4	0.4	0.8	9				
Baroda	22°19'N	73°13'E	90	77	11.9	1.0	0.3	6				
Aurangabad	19°52'N	75°24'E	85	70	7.4	2.0	0.2	10				
Indore	22°43'N	75°48'E	83	71	11.0	3.0	0.2	10				
Nagpur	21°05'N	79°02'E	87	75	13.9	5.0	0.3	10				
Kalpur	21°10'N	81°44'E	86	75	15.0	4.0	0.7	10			6.9	0
Jadapur	19°05'N	82°02'E	83	71	2.8	12.5	0.2	7			6.9	0
Vadkapatnam	17°43'N	83°13'E	89	79	7.8	5.0	1.5	10			6.5	0
Gopalpur	19°16'N	84°52'E	88	79	10.7	9.6	0.6	10				

TABLE 39

35 MM FIELD OF VIEW AND SIZE OF SUN'S IMAGE FOR VARIOUS PHOTOGRAPHIC FOCAL LENGTHS

<u>Focal Length</u>	<u>Field of View</u>	<u>Size of Sun</u>
28 mm	49° x 74°	0.2 mm
35 mm	39° x 59°	0.3 mm
50 mm	27° x 40°	0.5 mm
105 mm	13° x 19°	1.0 mm
200 mm	7° x 10°	1.8 mm
400 mm	3.4° x 5.1°	3.7 mm
500 mm	2.7° x 4.1°	4.6 mm
1000 mm	1.4° x 2.1°	9.2 mm
1500 mm	0.9° x 1.4°	13.8 mm
2000 mm	0.7° x 1.0°	18.4 mm
2500 mm	0.6° x 0.8°	22.9 mm

Image Size of Sun (mm) = Focal Length (mm) / 109

TABLE 40

SOLAR ECLIPSE EXPOSURE GUIDE

<u>ISO</u>	<u>f/Number</u>								
	1.4	2	2.8	4	5.6	8	11	16	22
25	1.4	2	2.8	4	5.6	8	11	16	22
50	2	2.8	4	5.6	8	11	16	22	32
100	2.8	4	5.6	8	11	16	22	32	44
200	4	5.6	8	11	16	22	32	44	64
400	5.6	8	11	16	22	32	44	64	88
800	8	11	16	22	32	44	64	88	128
1600	11	16	22	32	44	64	88	128	176

<u>Subject</u>	<u>Q</u>	<u>Shutter Speed</u>								
Solar Eclipse										
Partial ¹ - 4.0 ND	11	—	—	—	1/4000	1/2000	1/1000	1/500	1/250	1/125
Partial ¹ - 5.0 ND	8	1/4000	1/2000	1/1000	1/500	1/250	1/125	1/60	1/30	1/15
Baily's Beads ²	11	—	—	—	1/4000	1/2000	1/1000	1/500	1/250	1/125
Chromosphere	10	—	—	1/4000	1/2000	1/1000	1/500	1/250	1/125	1/60
Prominences	9	—	1/4000	1/2000	1/1000	1/500	1/250	1/125	1/60	1/30
Corona - 0.1 R	7	1/2000	1/1000	1/500	1/250	1/125	1/60	1/30	1/15	1/8
Corona - 0.2 R ³	5	1/500	1/250	1/125	1/60	1/30	1/15	1/8	1/4	1/2
Corona - 0.5 R	3	1/125	1/60	1/30	1/15	1/8	1/4	1/2	1 sec	2 sec
Corona - 1.0 R	1	1/30	1/15	1/8	1/4	1/2	1 sec	2 sec	4 sec	8 sec
Corona - 2.0 R	0	1/15	1/8	1/4	1/2	1 sec	2 sec	4 sec	8 sec	15 sec
Corona - 4.0 R	-1	1/8	1/4	1/2	1 sec	2 sec	4 sec	8 sec	15 sec	30 sec
Corona - 8.0 R	-3	1/2	1 sec	2 sec	4 sec	8 sec	15 sec	30 sec	1 min	2 min

Exposure Formula: $t = f^2 / (I \times 2^Q)$

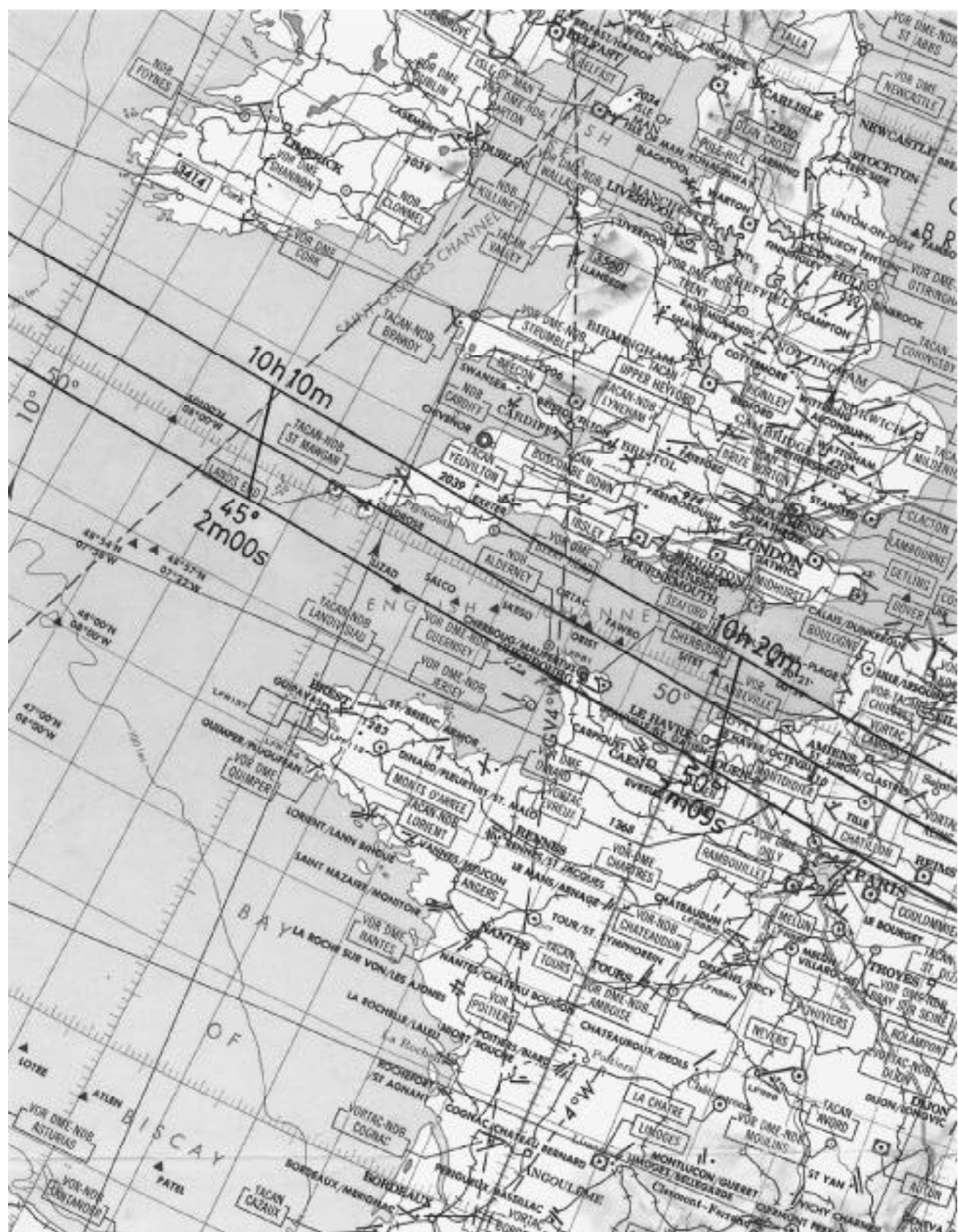
where: t = exposure time (sec)
 f = f/number or focal ratio
 I = ISO film speed
 Q = brightness exponent

Abbreviations: ND = Neutral Density Filter.
 R_s = Solar Radii.

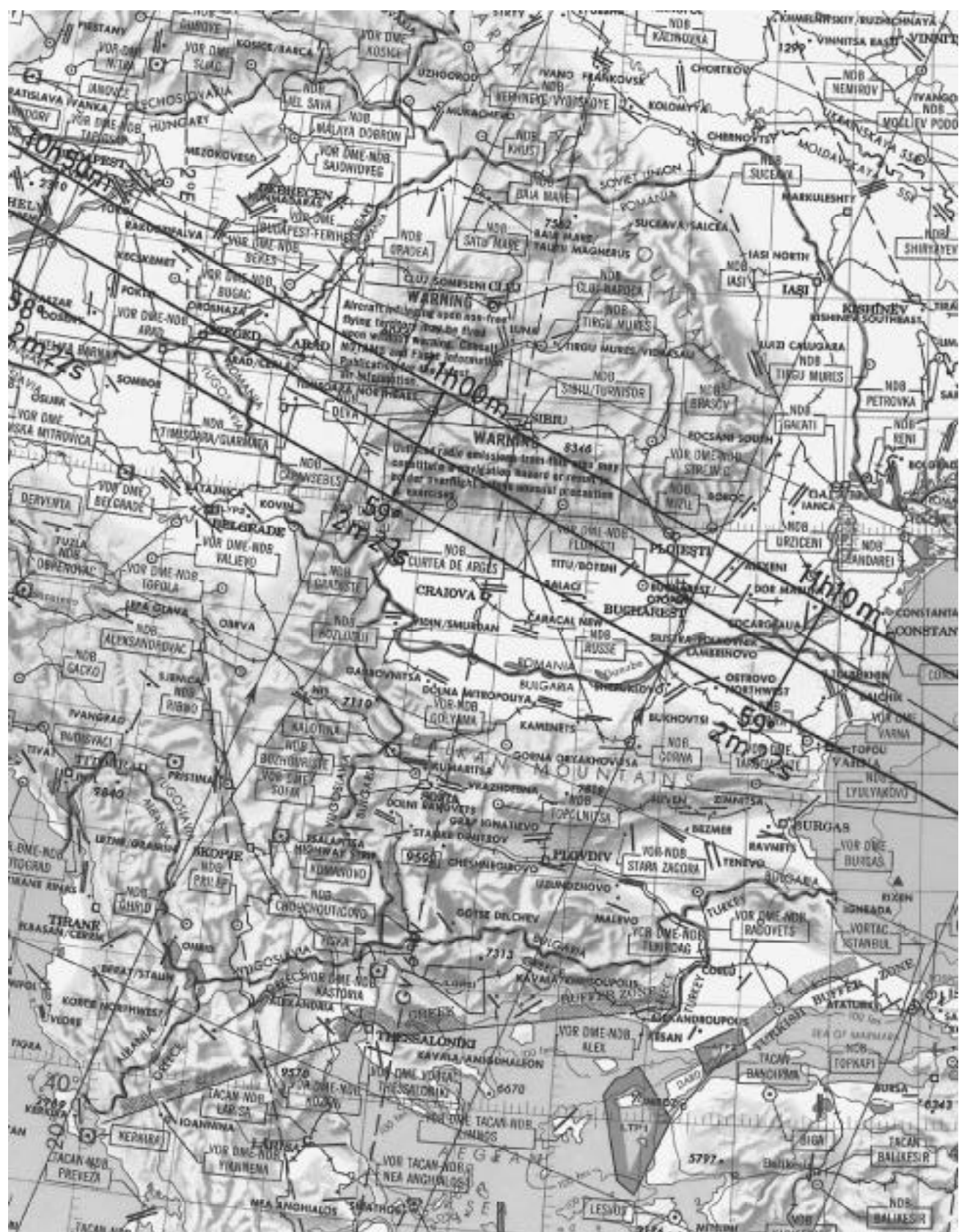
Notes: ¹ Exposures for partial phases are also good for annular eclipses.
² Baily's Beads are extremely bright and change rapidly.
³ This exposure also recommended for the 'Diamond Ring' effect.

TOTAL SOLAR ECLIPSE OF 1999 AUGUST 11

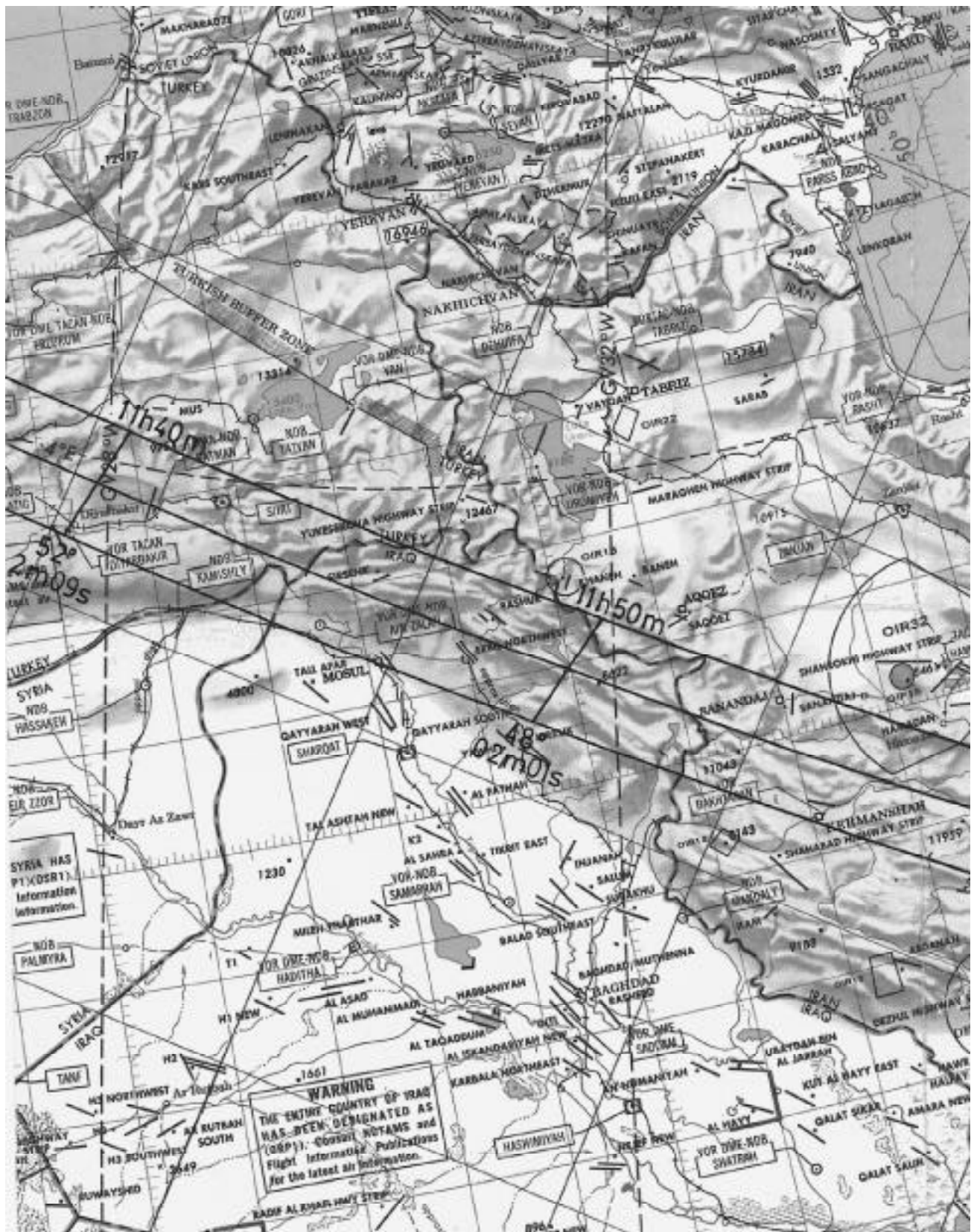
**MAPS
OF THE
UMBRA L PATH**

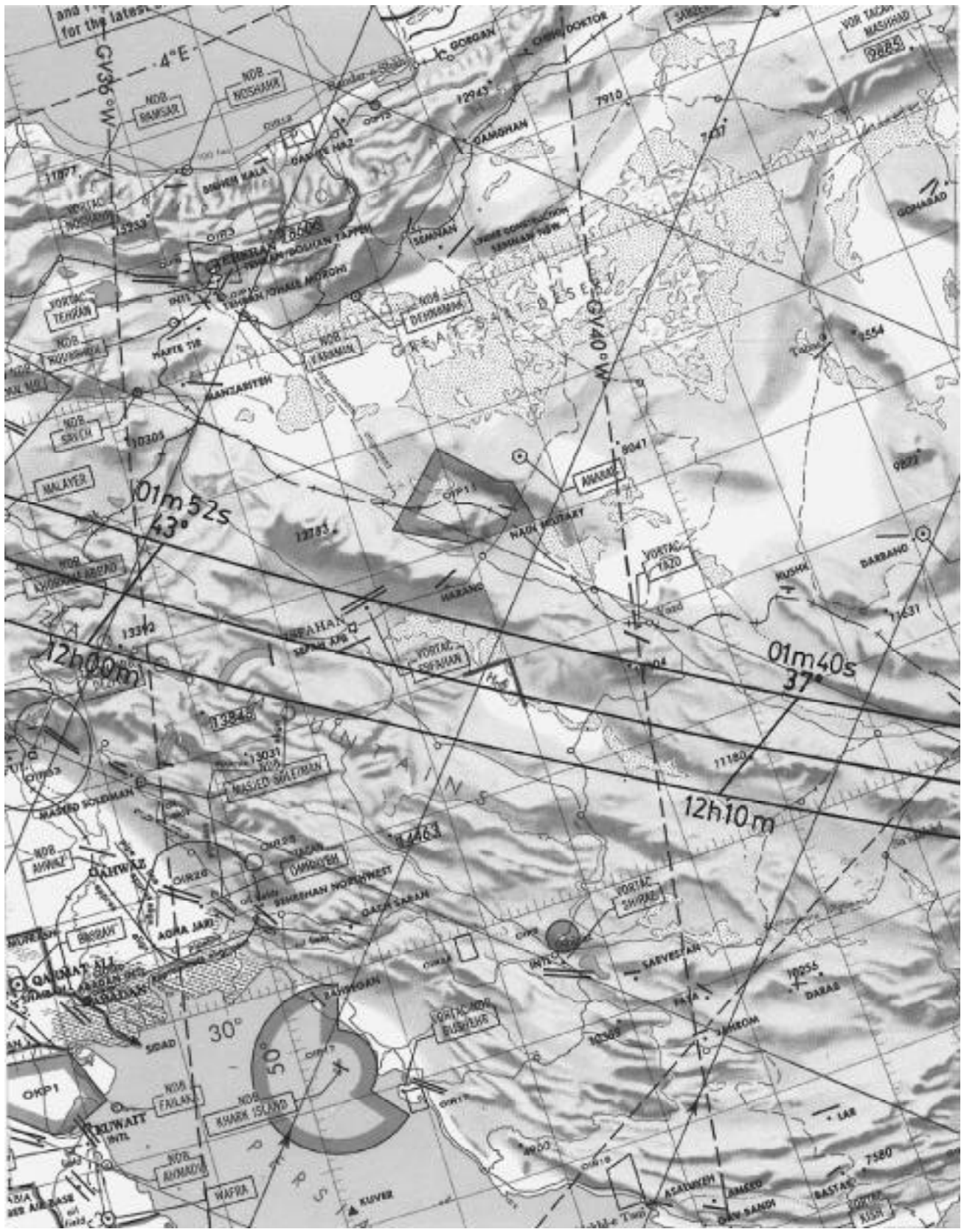


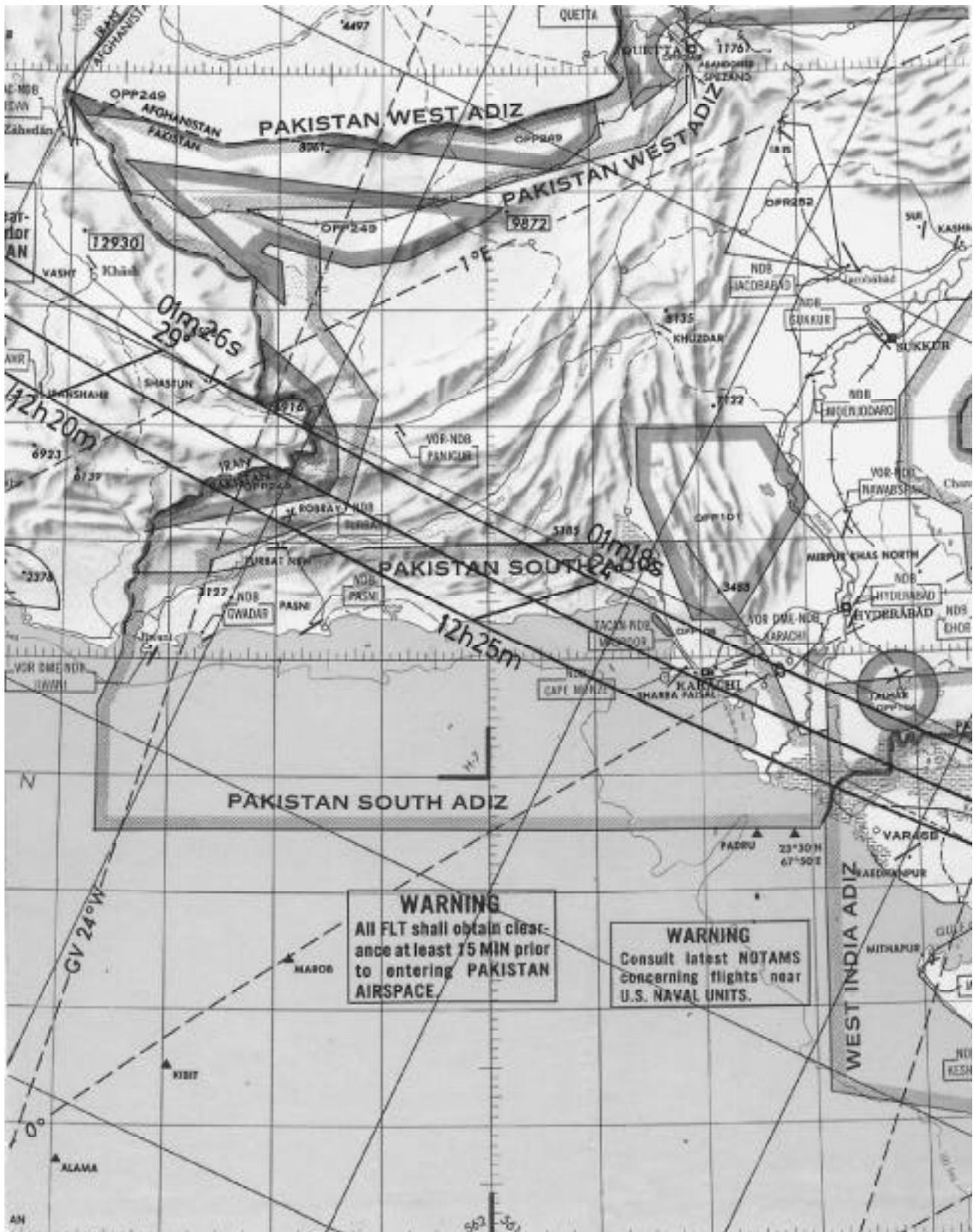


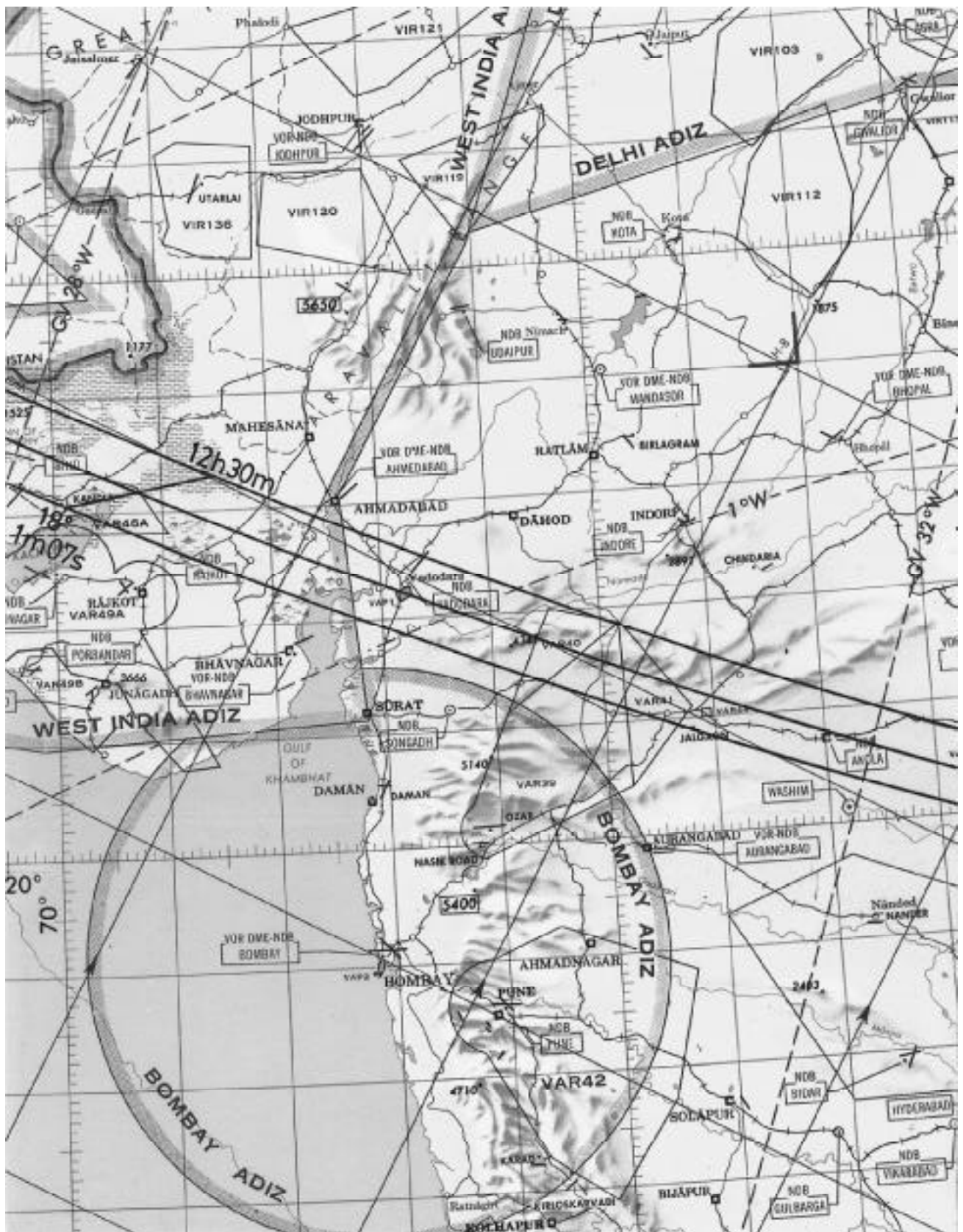


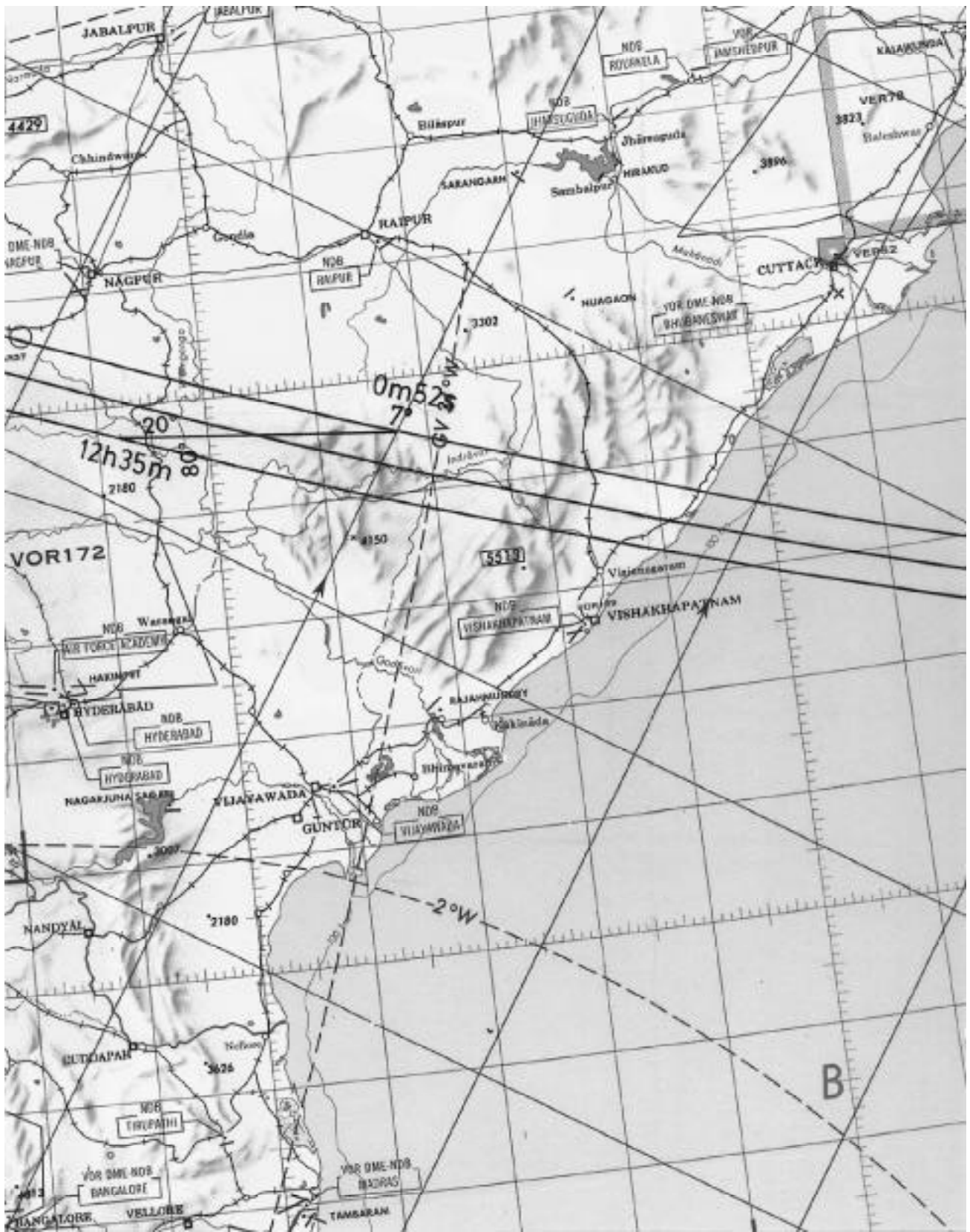












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